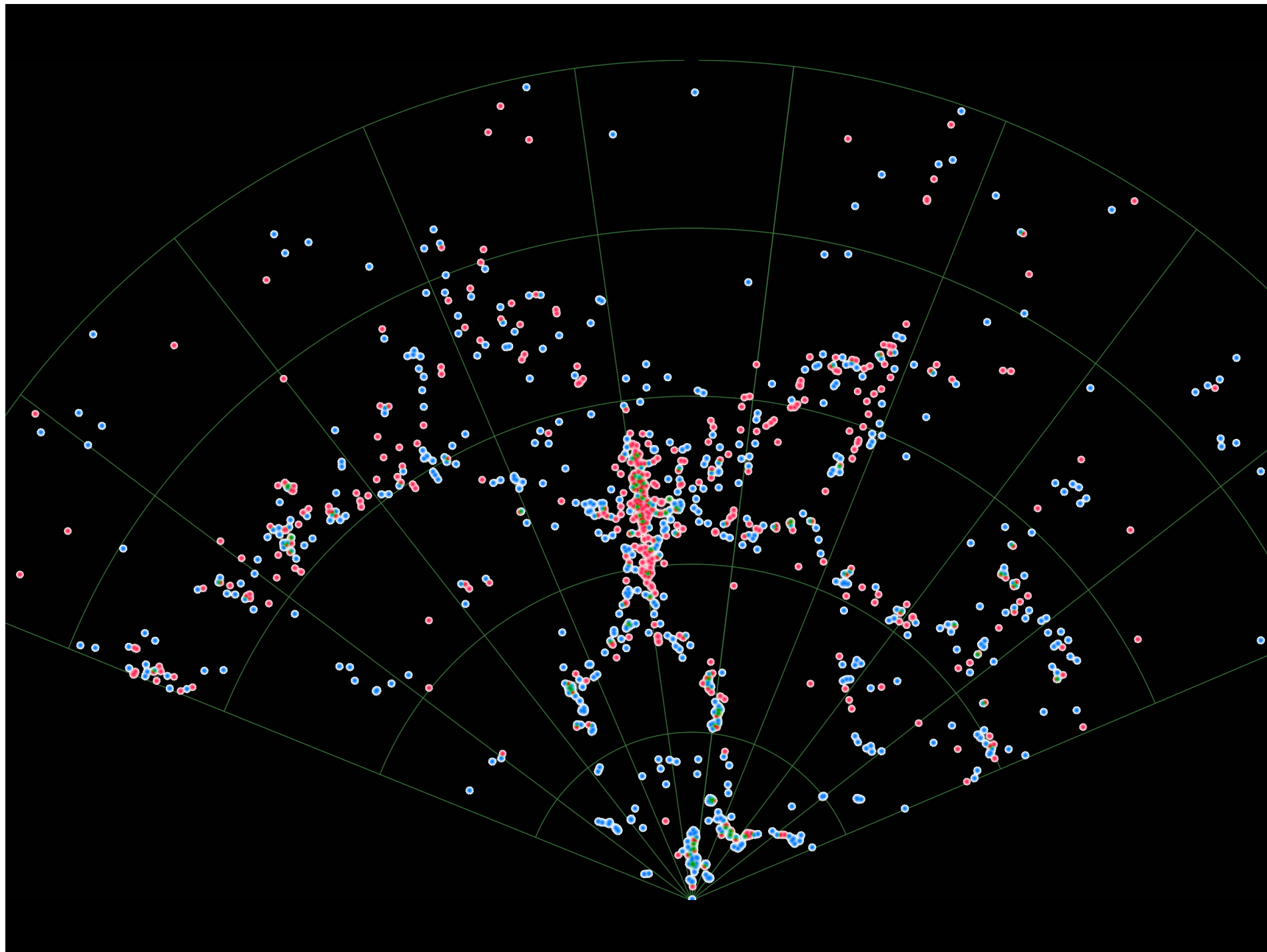
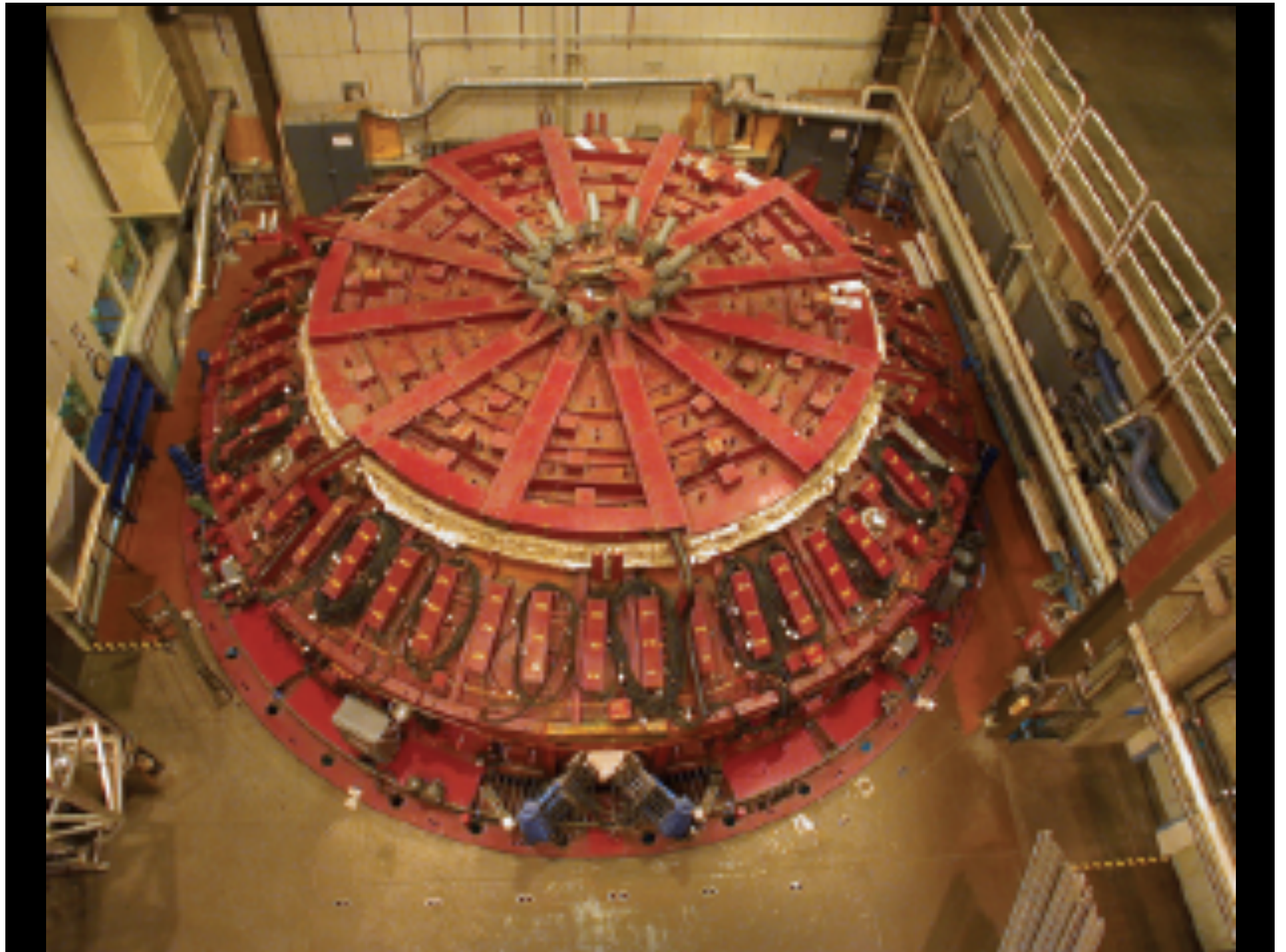


Building the HectoMAP 1986-2012+

Michael J. Kurtz

Harvard-Smithsonian Center for Astrophysics



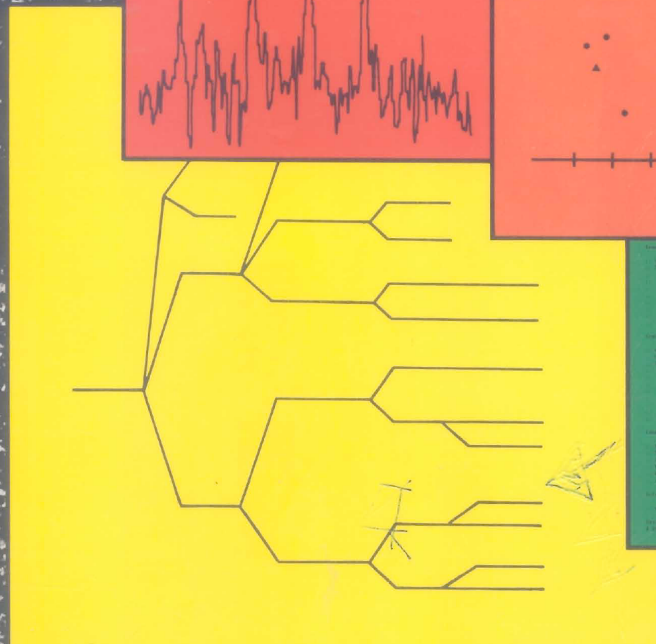
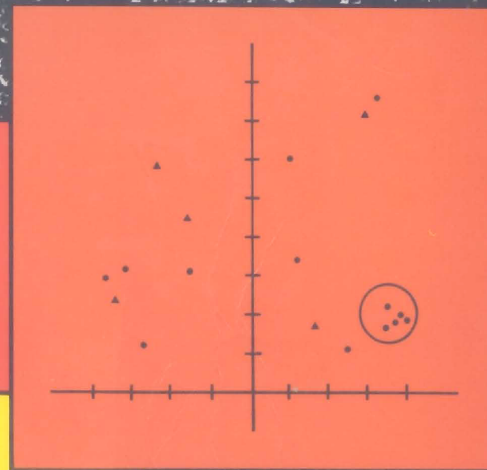
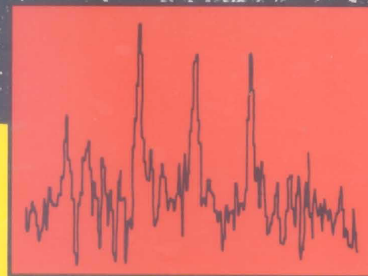




ASTRONOMY FROM LARGE DATABASES

SCIENTIFIC OBJECTIVES AND METHODOLOGICAL APPROACHES

Garching, 12–14 October 1987



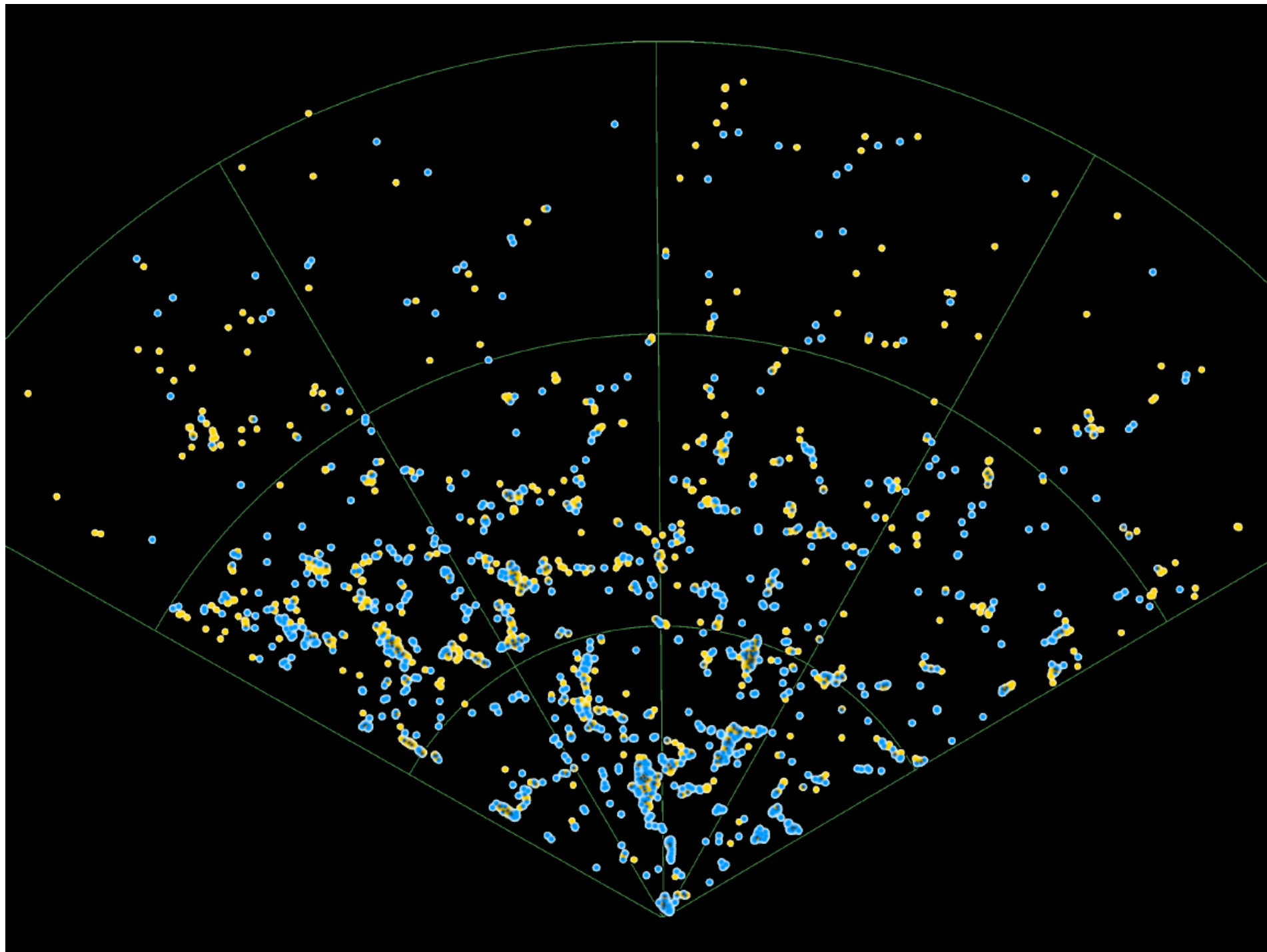
Procedures for the analysis of astronomical data
The analysis of astronomical data is a complex task that requires the use of sophisticated statistical methods. This paper describes the procedures used in the analysis of data from the European Space Telescope (EST) and discusses the challenges involved in the analysis of large databases.

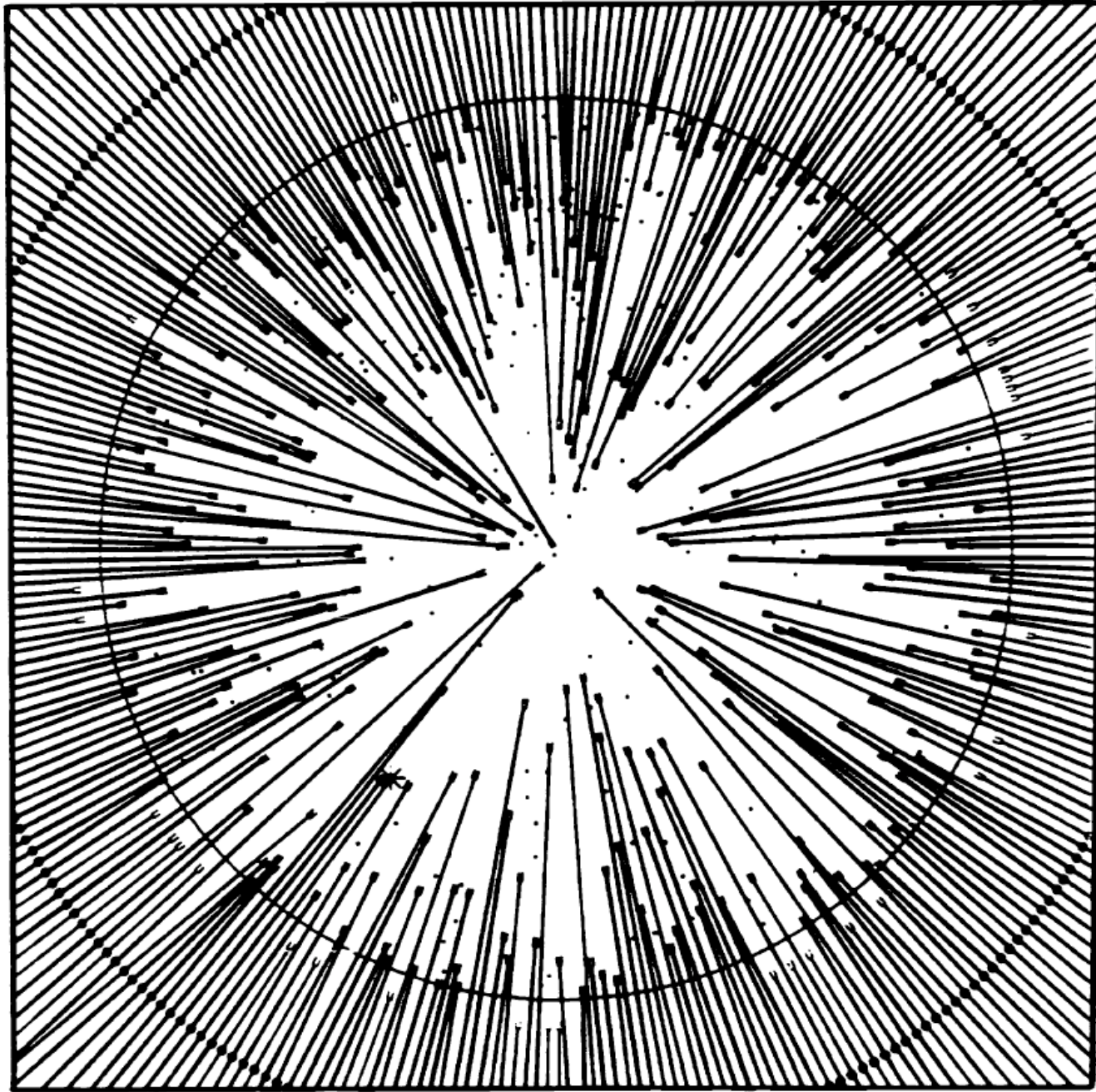
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Certainly our knowledge of structures at larger scales does not permit a detailed characterization of them at this time. The recent discovery by de Lapparent *et al.* 1986 of galaxy-filled surfaces surrounding large empty regions similar to the void in Bootes (Kirshner *et al.* 1987) can give rise however to some wild speculation. Perhaps sometime in the next century astronomers will have distinguished several different kinds of surfaces and voids, and will be able to attempt a comparison of the distribution of the various types with galaxy clusters, with each other, and with the distribution of fluctuations (by then discovered?) in the cosmic microwave background. From this, knowledge about the VERY large scale structure of the universe may result.





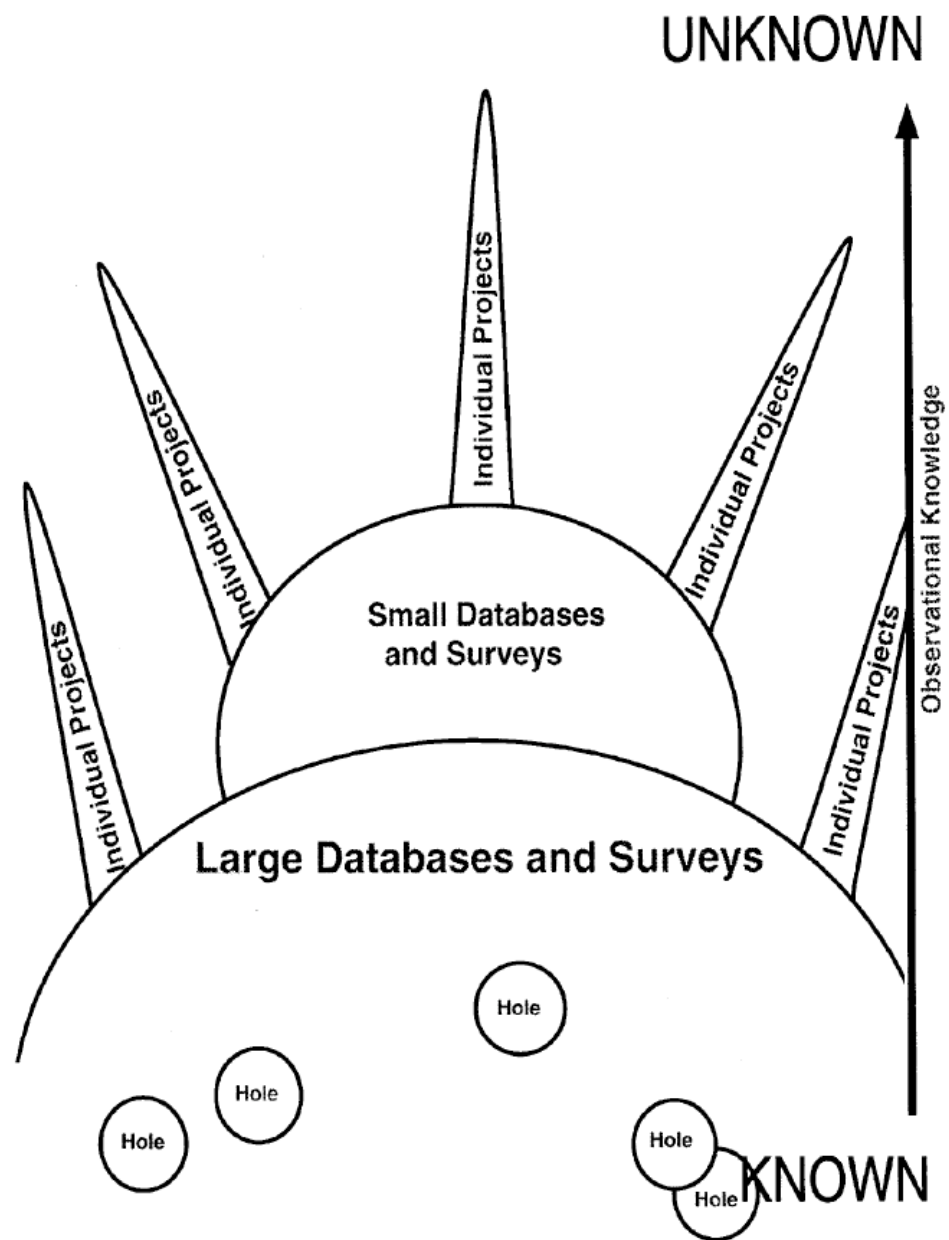


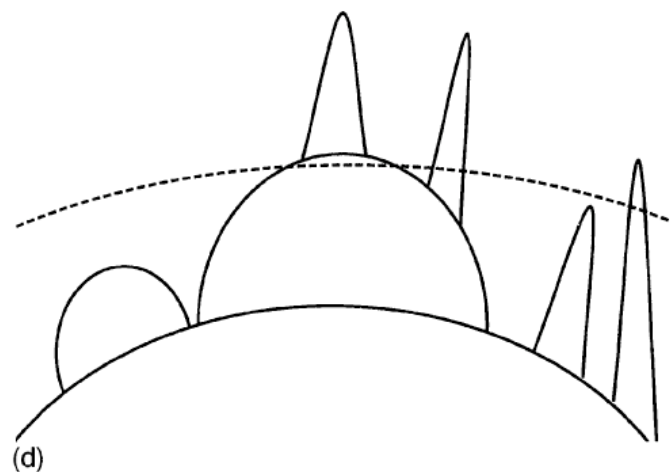
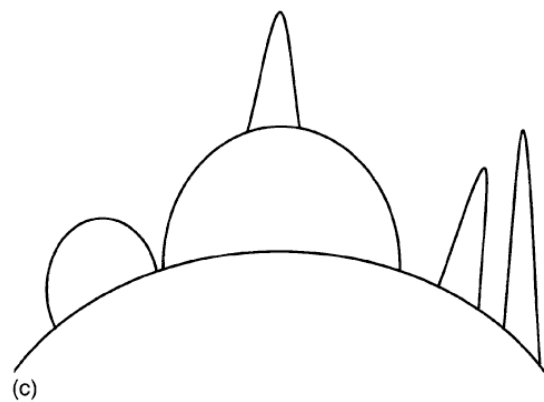
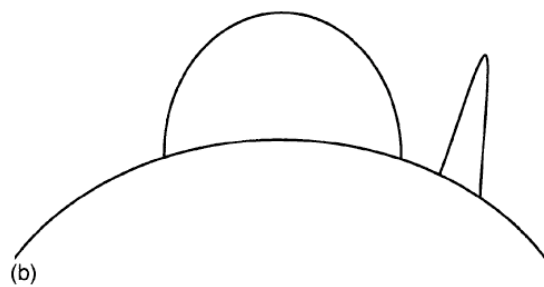
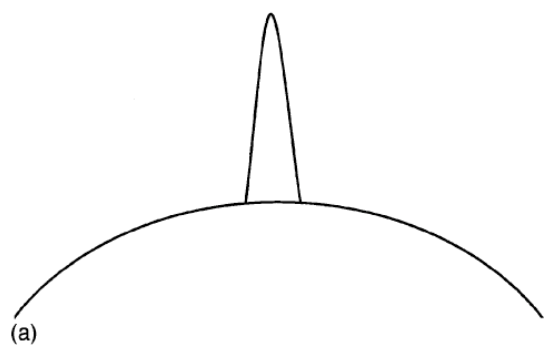
16 The Need for Automation in the Reduction and Analysis of Stellar Spectra

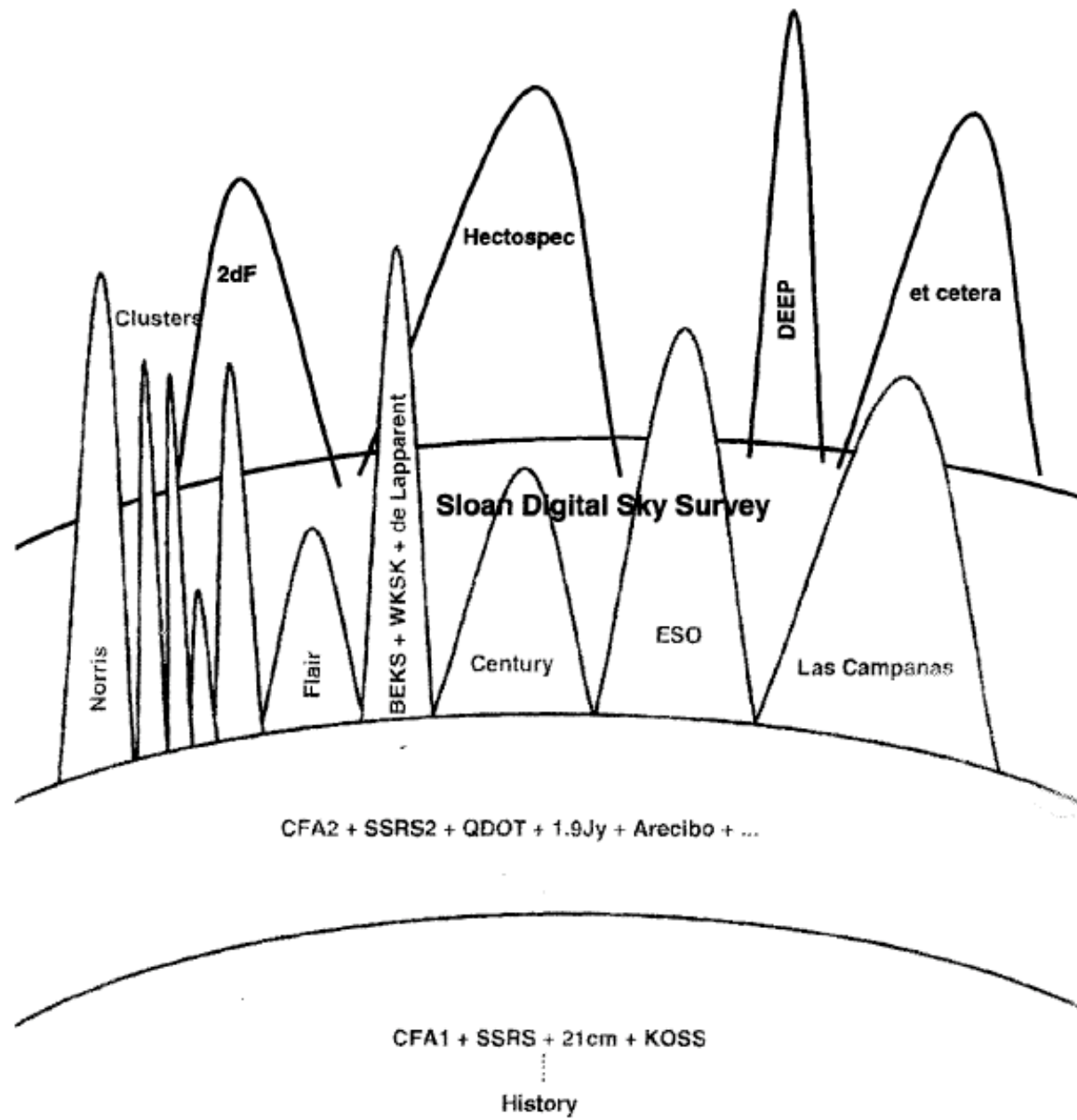
M. J. Kurtz

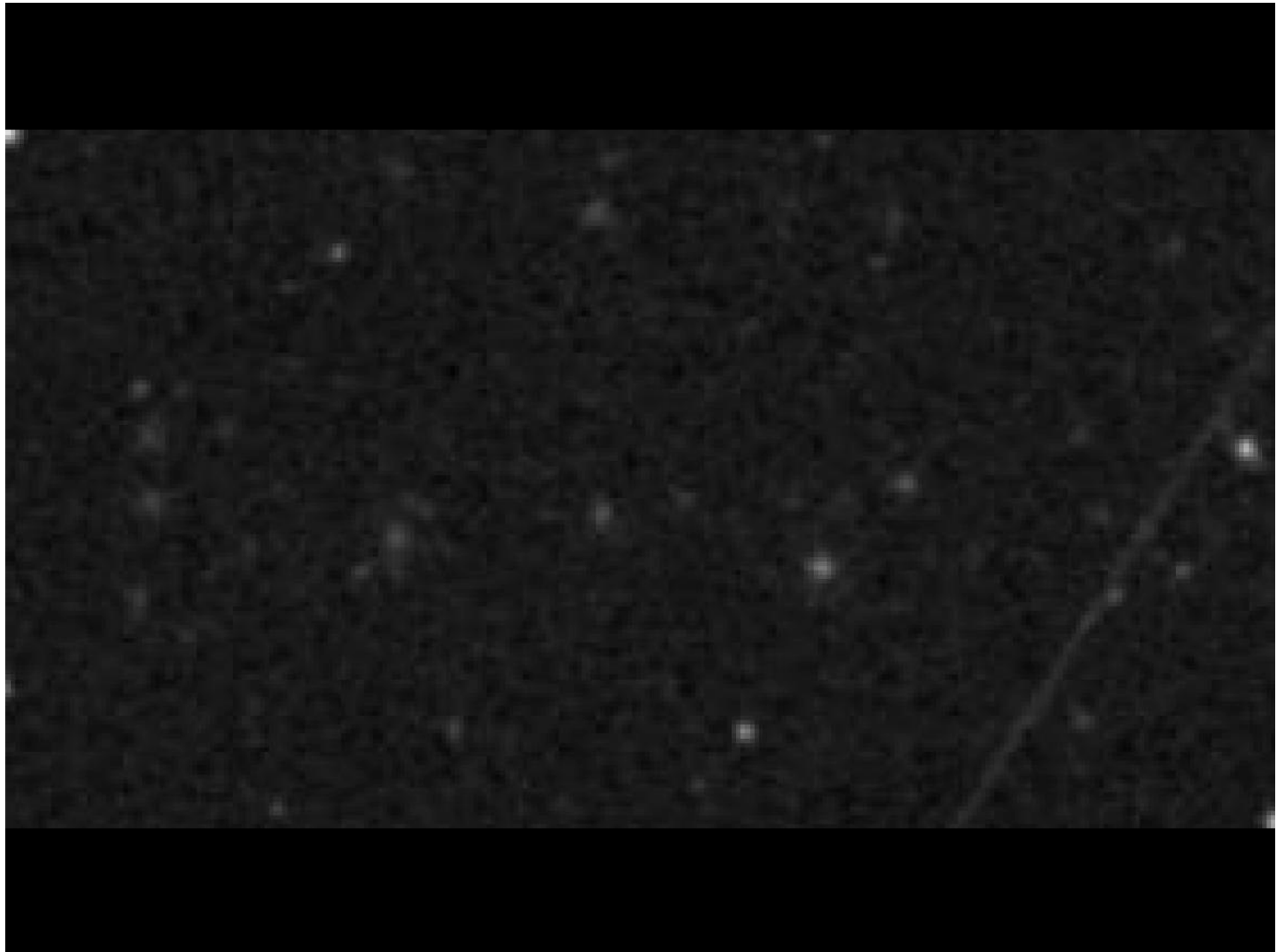
Advances in instrumentation and optics have brought us to the edge of a new era in spectroscopy. Beginning in 1995, it will be possible for a single instrument, the upgraded MMT, using a 300-fiber spectrograph, to obtain 1000 classification quality spectra per hour for stars with $V < 14$. Several other similar projects, both larger and smaller, are planned.

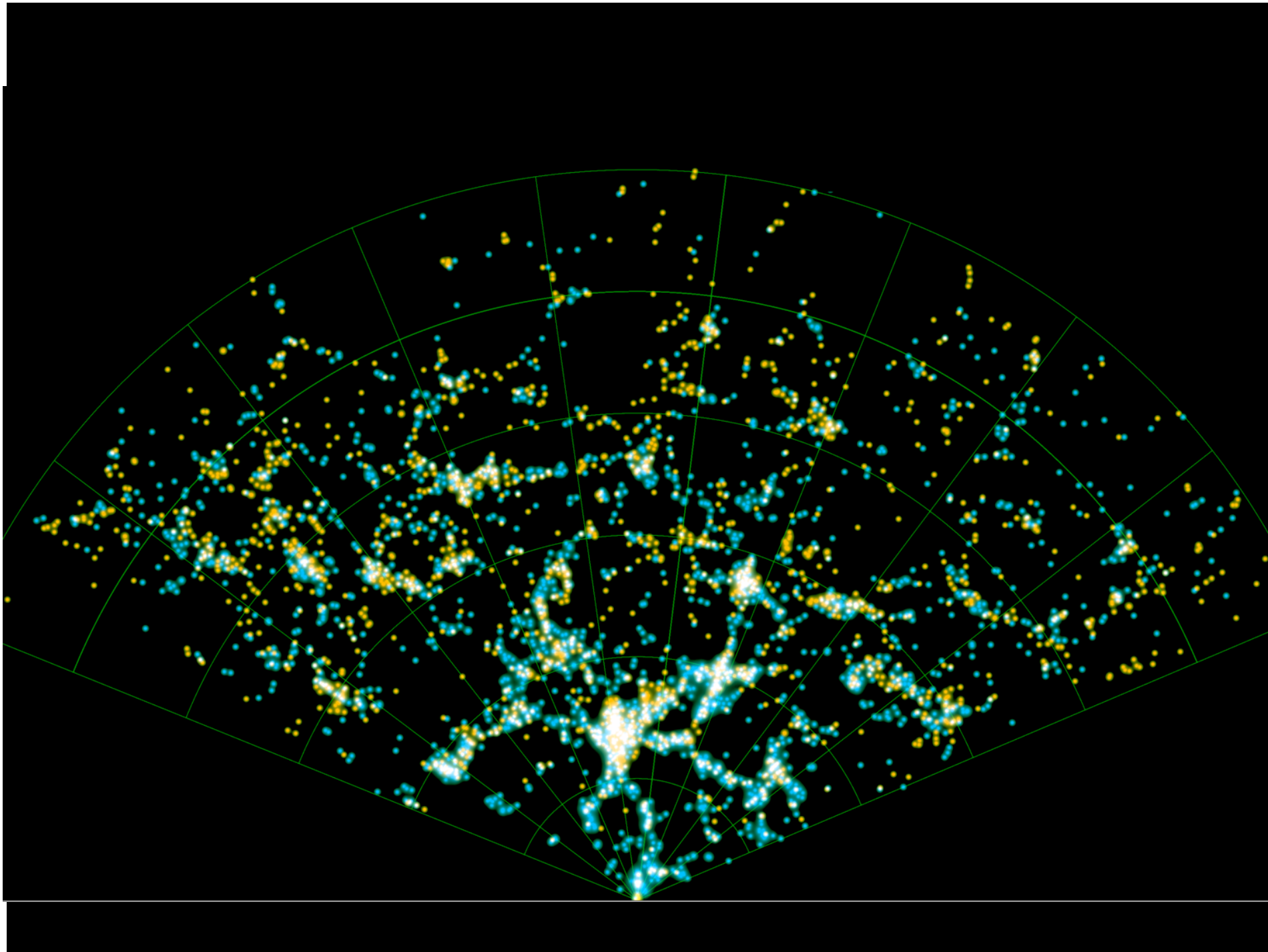
The ability of individual researchers to obtain 100,000 spectra per year will cause substantial changes in the way they are used. For results to be understood, a standard nomenclature describing normal spectra must be used, and it must be possible to obtain these descriptors automatically. The MK system can and should serve as the basis for this system of description; an automated procedure for obtaining MK types must be developed.

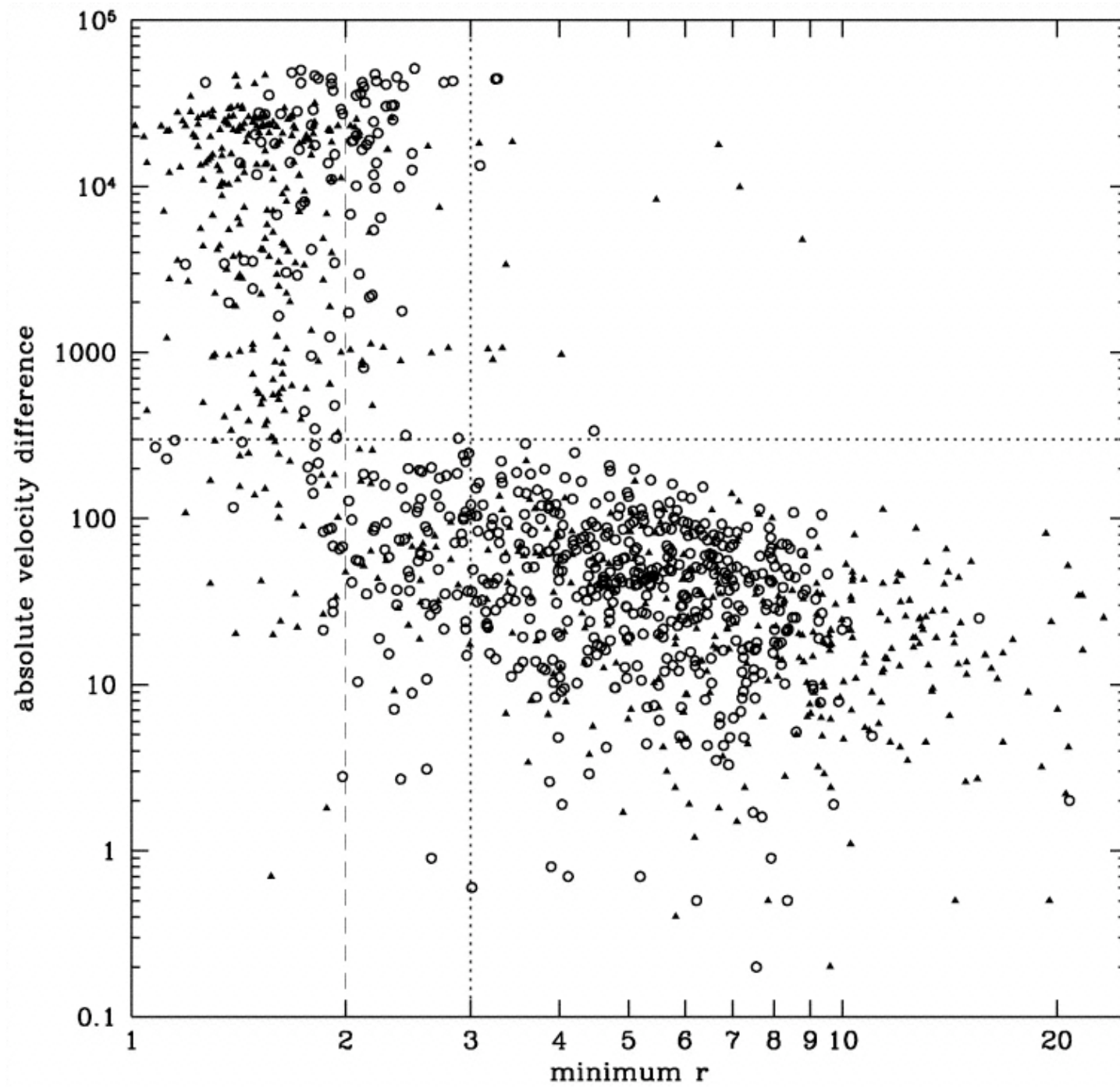


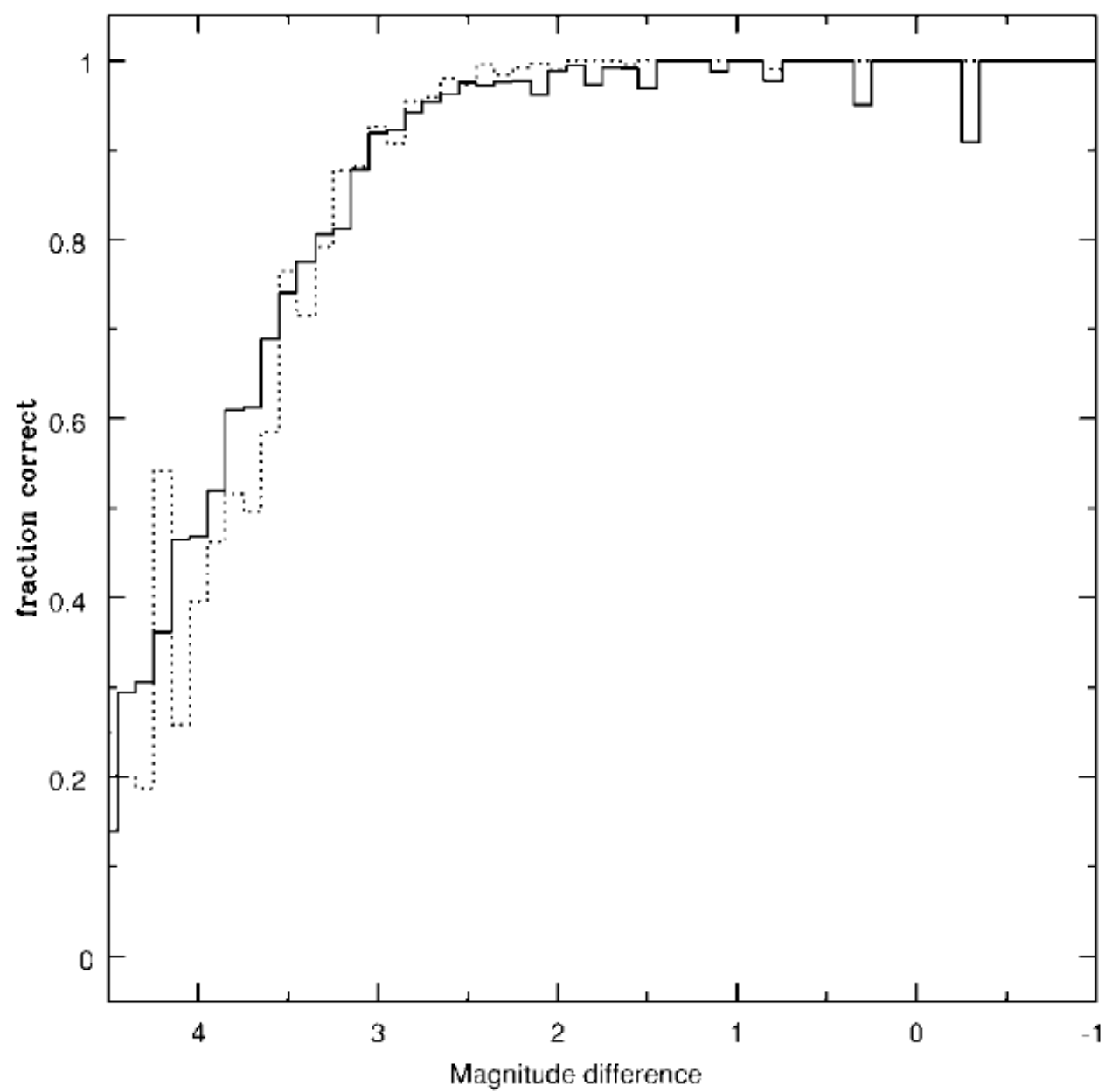


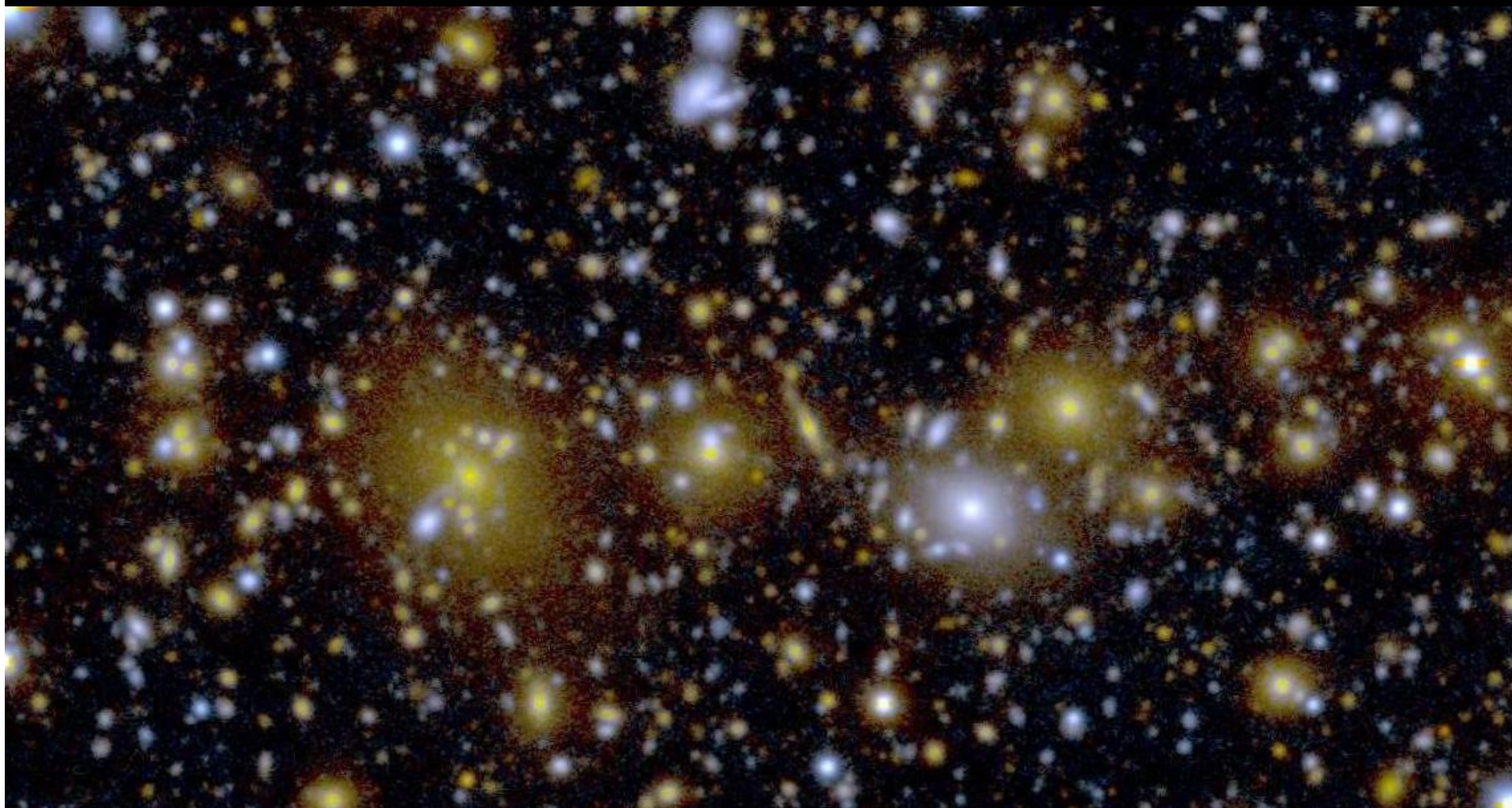


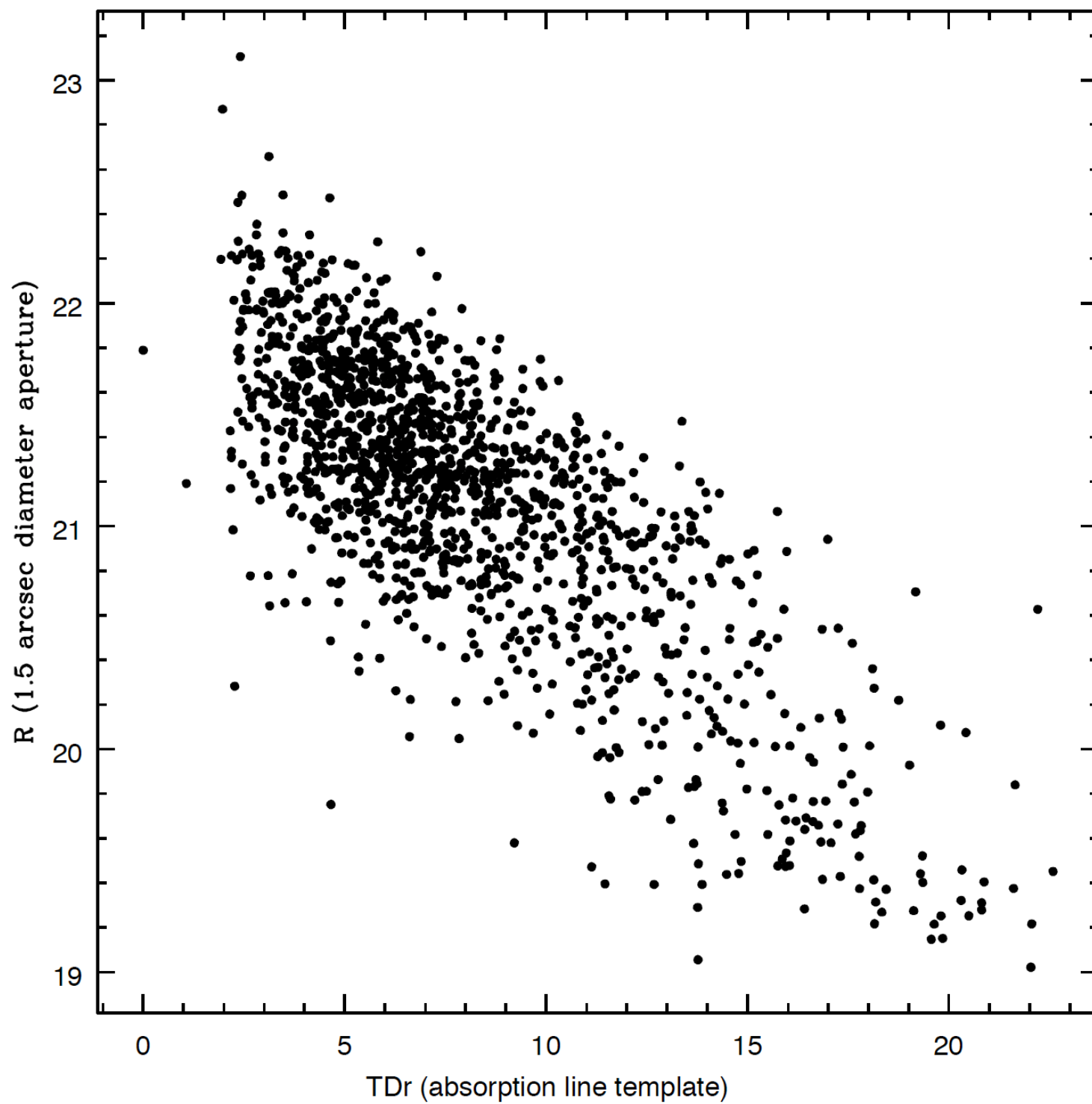


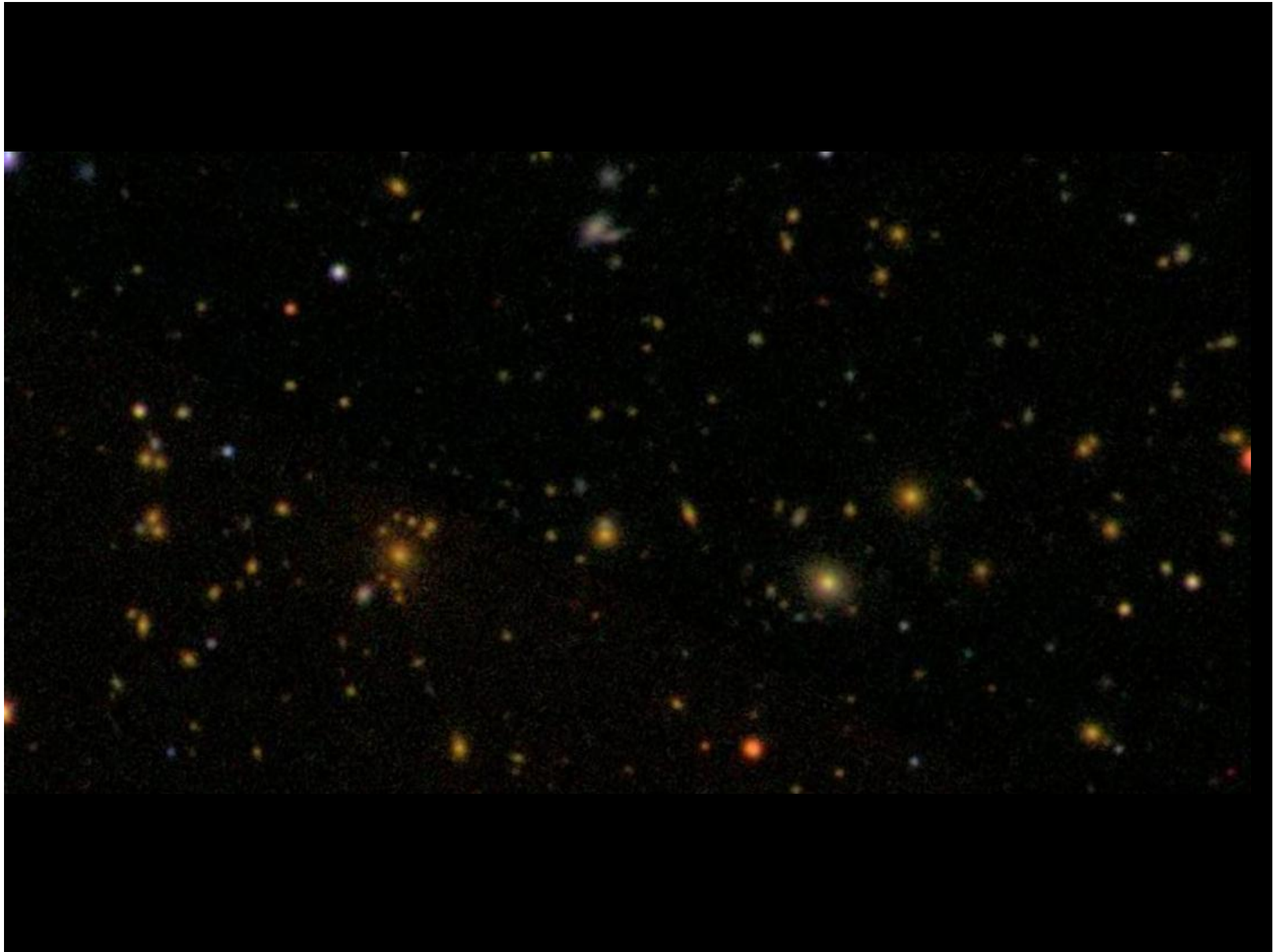


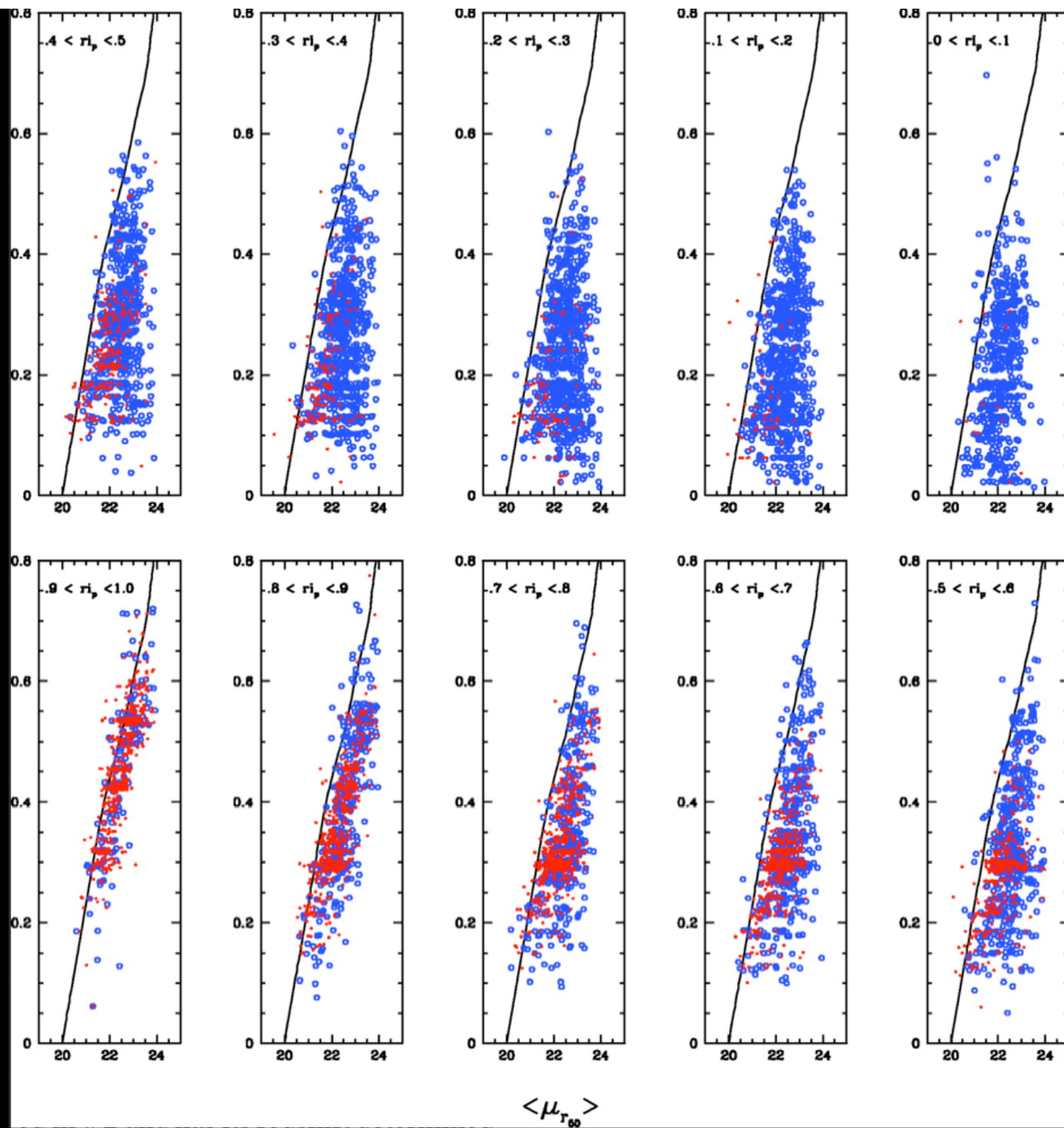


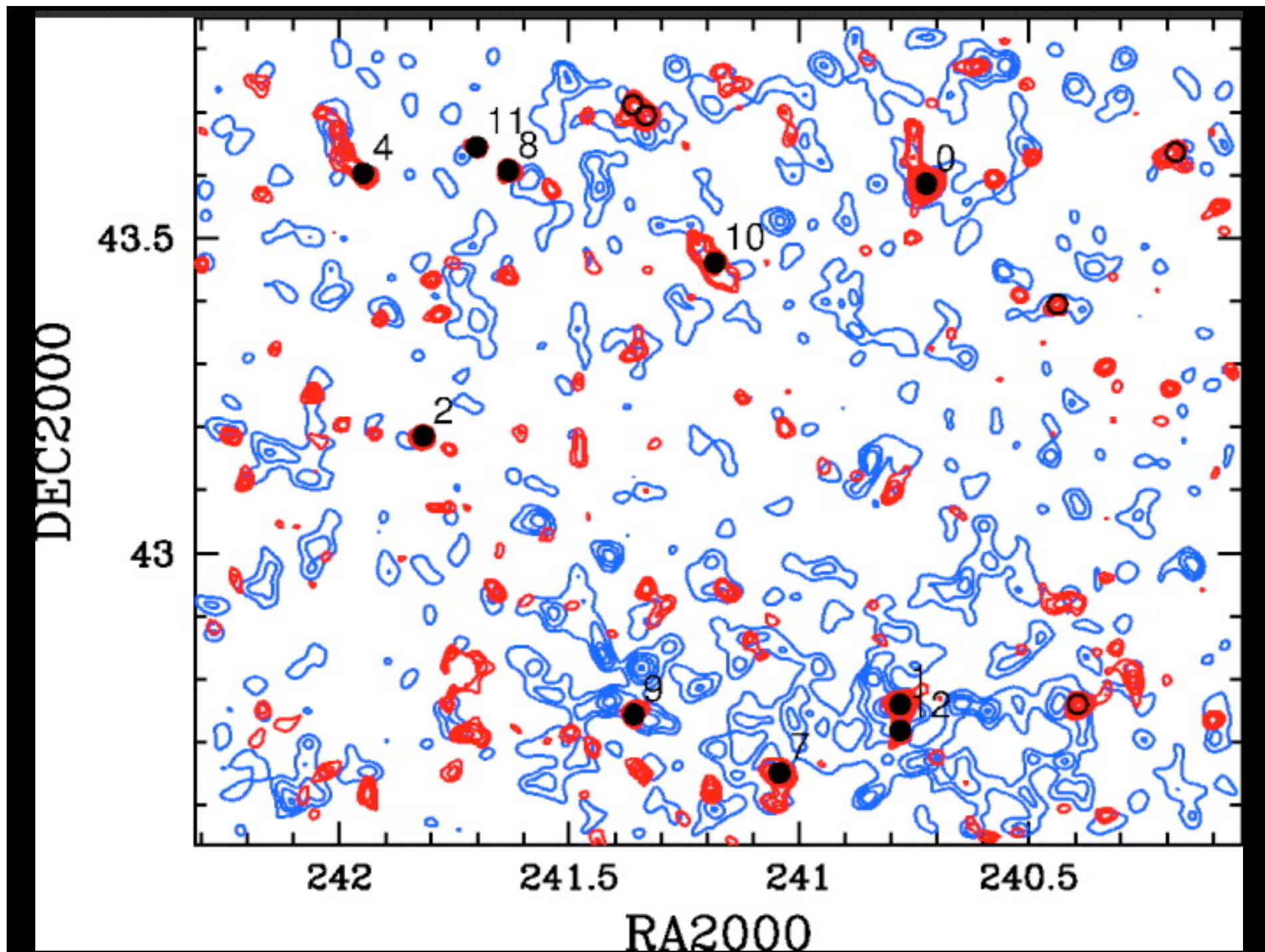














HectoMAP

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Smithsonian Astrophysical Observatory

