

# PLANETARY DEBRIS AROUND WHITE DWARFS IN ZTF

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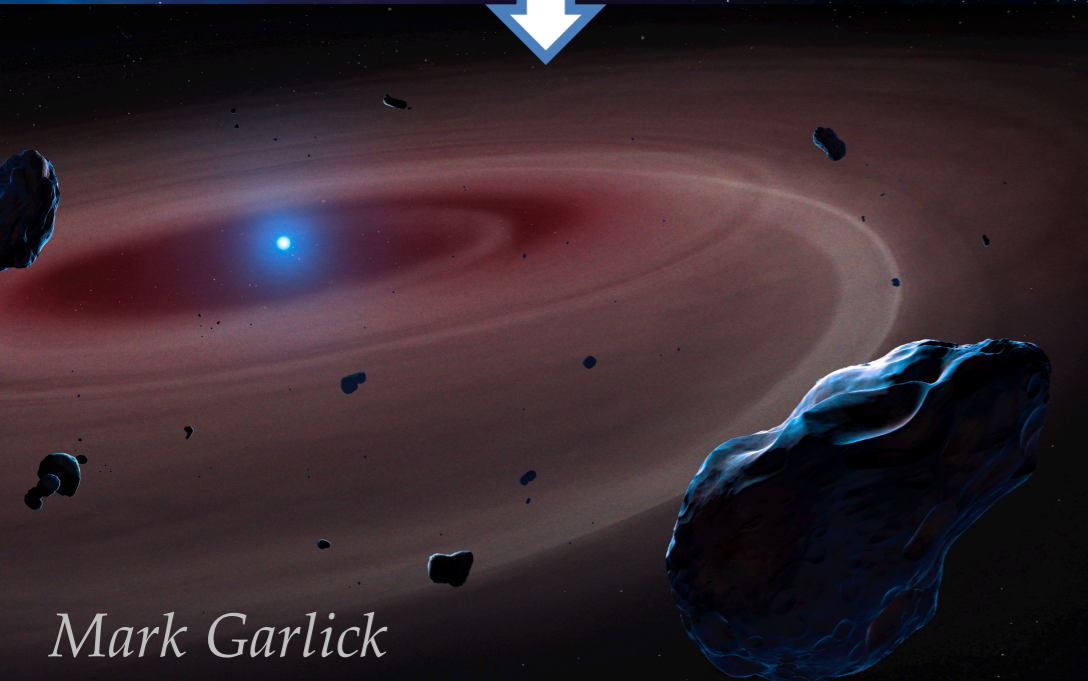
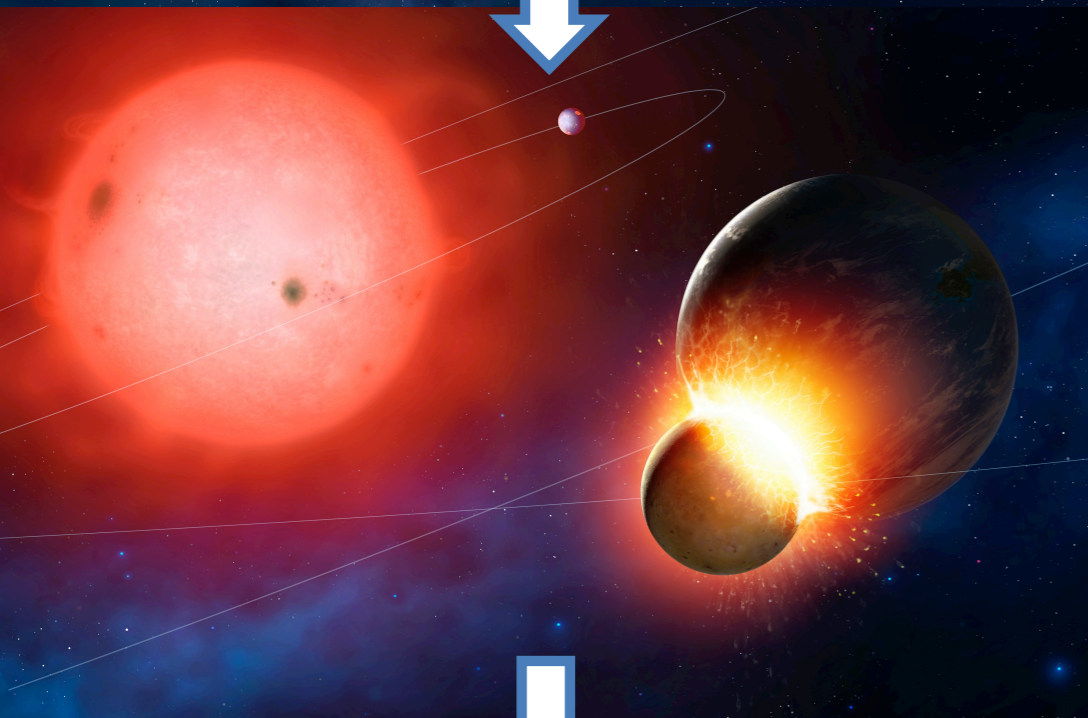
McDonald Observatory  
The University of Texas at Austin



The University of Texas at Austin  
Department of Astronomy  
College of Natural Sciences







1. We are observing the future of planetary systems around 2-3 solar-mass ZAMS stars.

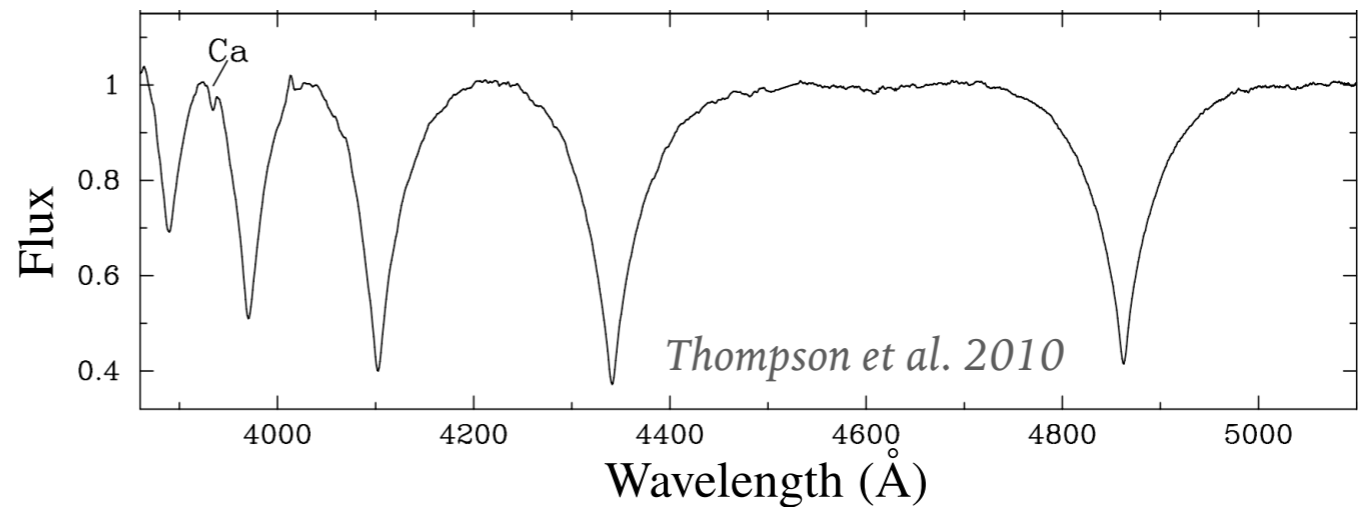
2. As the host evolves, the **orbits** of surviving planets **expand**. Objects destabilize, some **scatter** in.

3. Ancient solar systems have leftover debris. We can see it if it pollutes a pristine **white dwarf**.

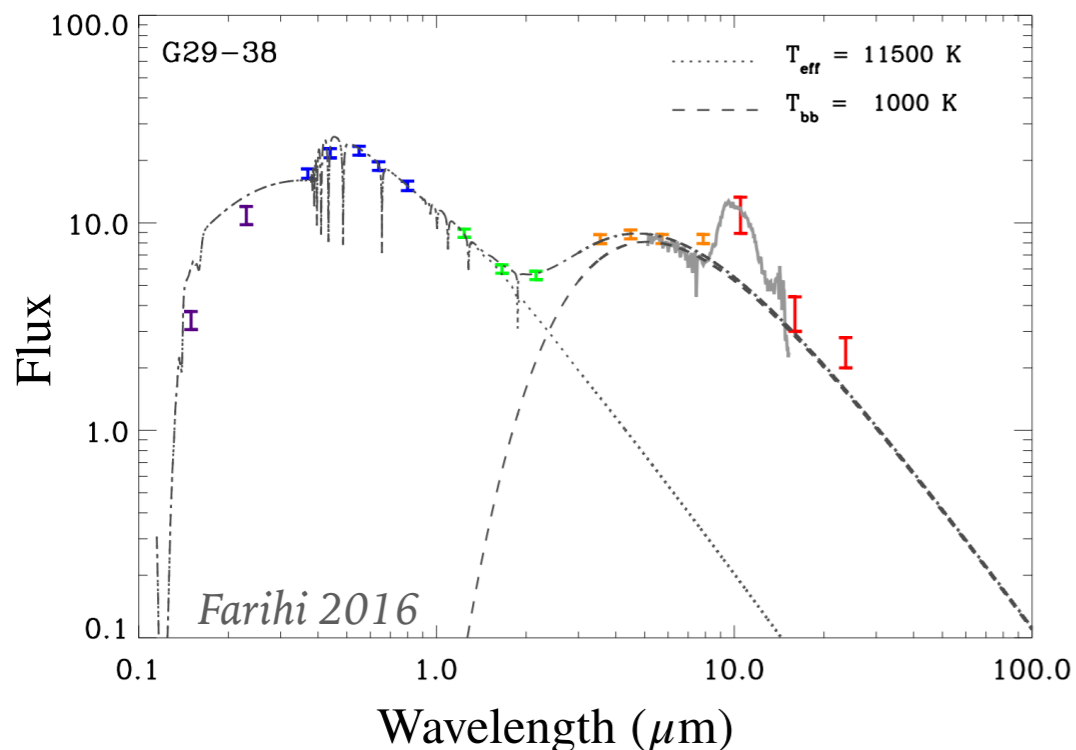
# DETECTING PLANETARY DEBRIS AROUND WHITE DWARFS

**DAZ = Metal Polluted  
H-atmosphere White Dwarf**

≈30% of white dwarfs below  
20,000 K exhibit metal pollution

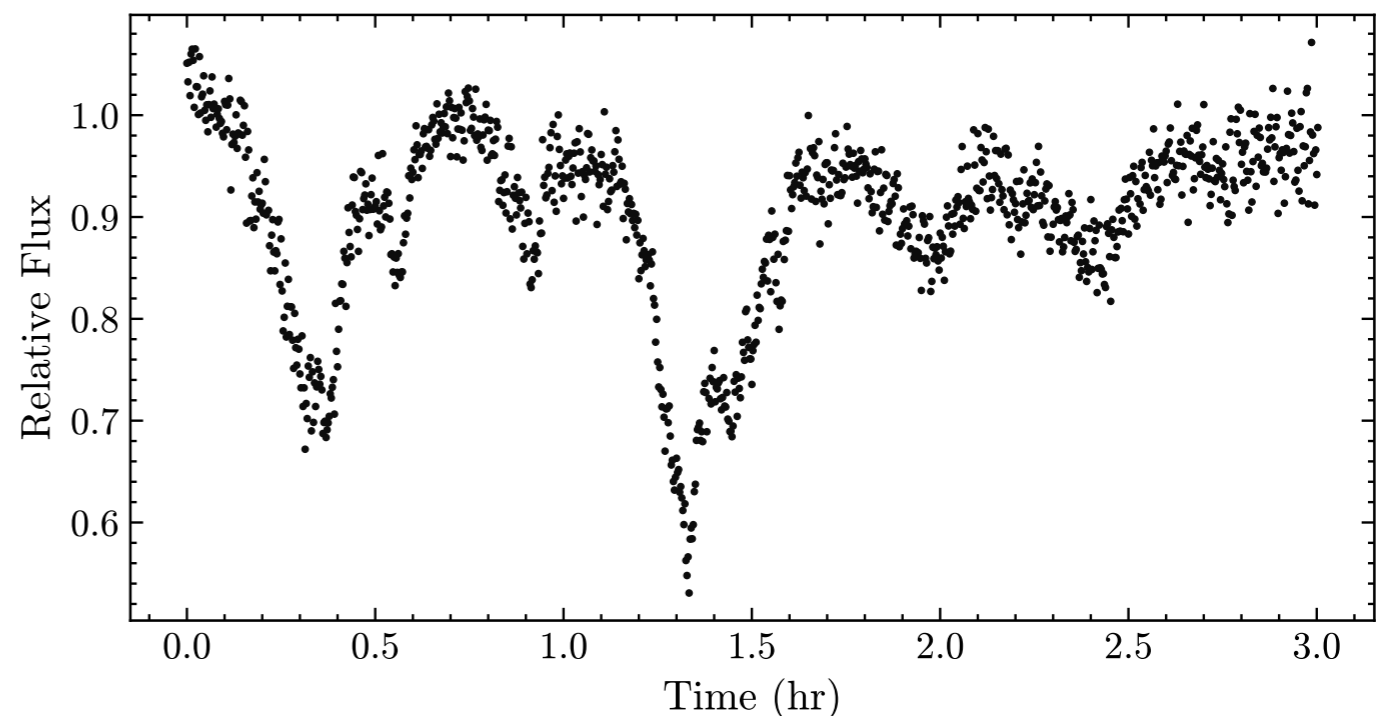


IR-excess due to warm  
circumstellar dust seen in  
1-2% of white dwarfs



Transiting Planetary Debris offers the most  
direct evidence, only two objects known

**WD 1145+017, 4.5–4.9 hour orbital periods**



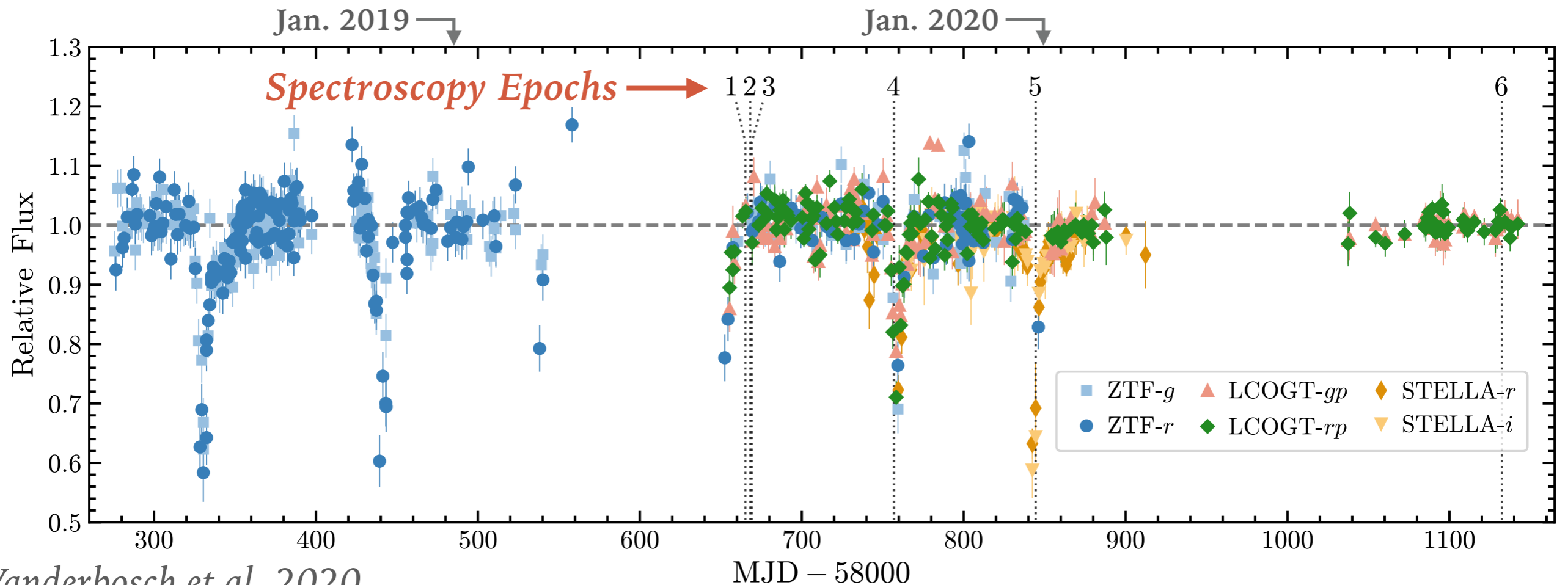


# ZTF J0139+5245: TRANSITING DEBRIS IN A 107-DAY ORBIT

Orbital period **Two Orders of Magnitude** longer than **WD 1145+017!**

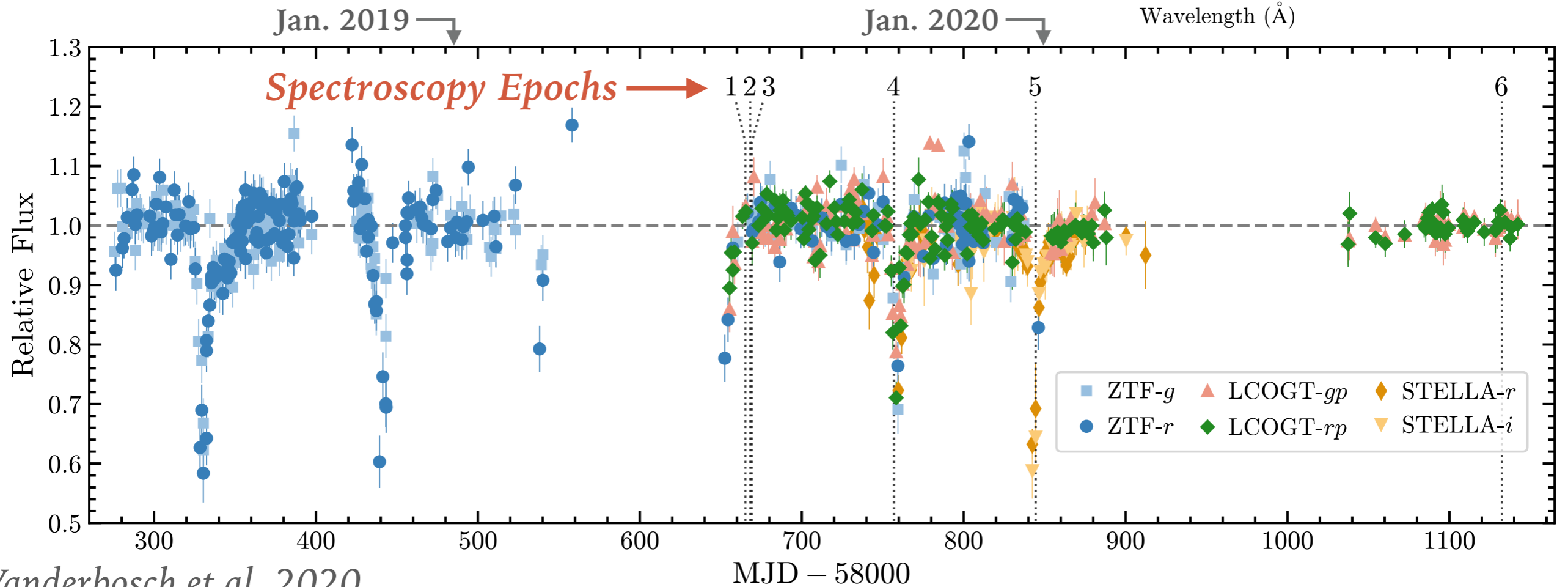
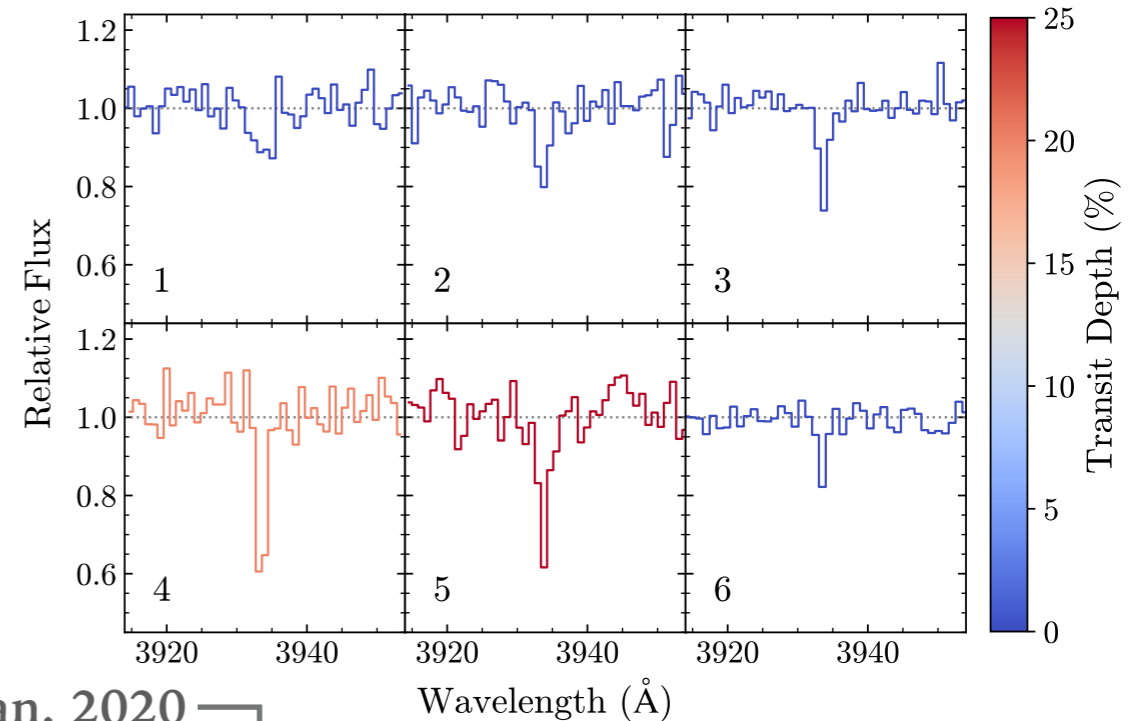
## Key Properties

$T_{\text{eff}} = 10,530 \text{ K}$	$P_{\text{orb}} \approx 107 \text{ days}$
$M_* = 0.52 M_{\odot}$	Transit Durations $\approx 15 - 25 \text{ days}$
1 – 10 year Ca Diffusion Timescale	Minimum Debris Mass comparable to solar system objects ( $> 10^{16} \text{ g}$ )

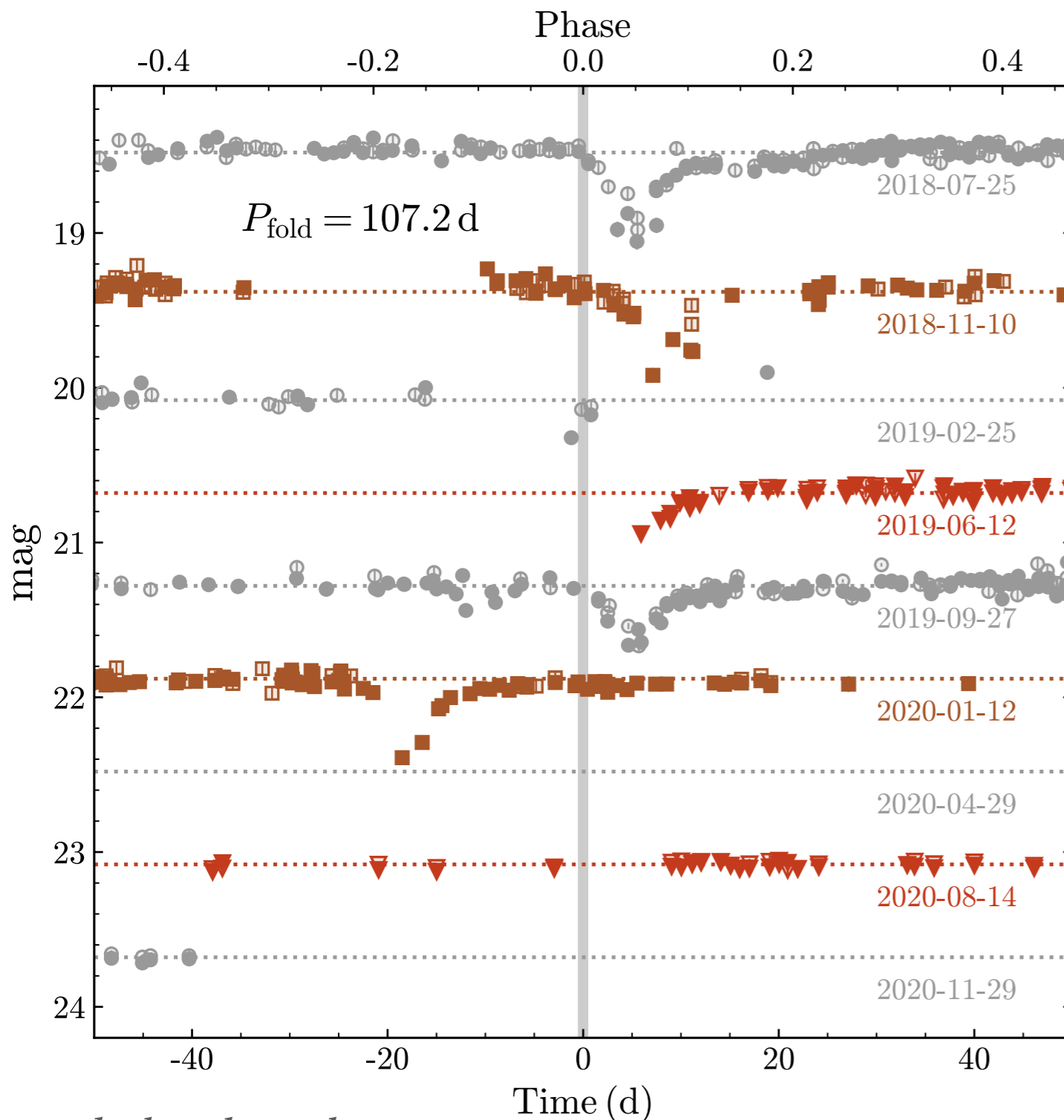


# ZTF J0139+5245: DEPTH-DEPENDENT CALCIUM ABSORPTION

- 1-10 year diffusion timescale suggests recent accretion events are an unlikely source.
- Absorption from circumstellar debris is likely responsible.
- Active disruption or collision events are required to maintain supply of metallic gas.



# ZTF J0139+5245: A DYNAMIC SYSTEM



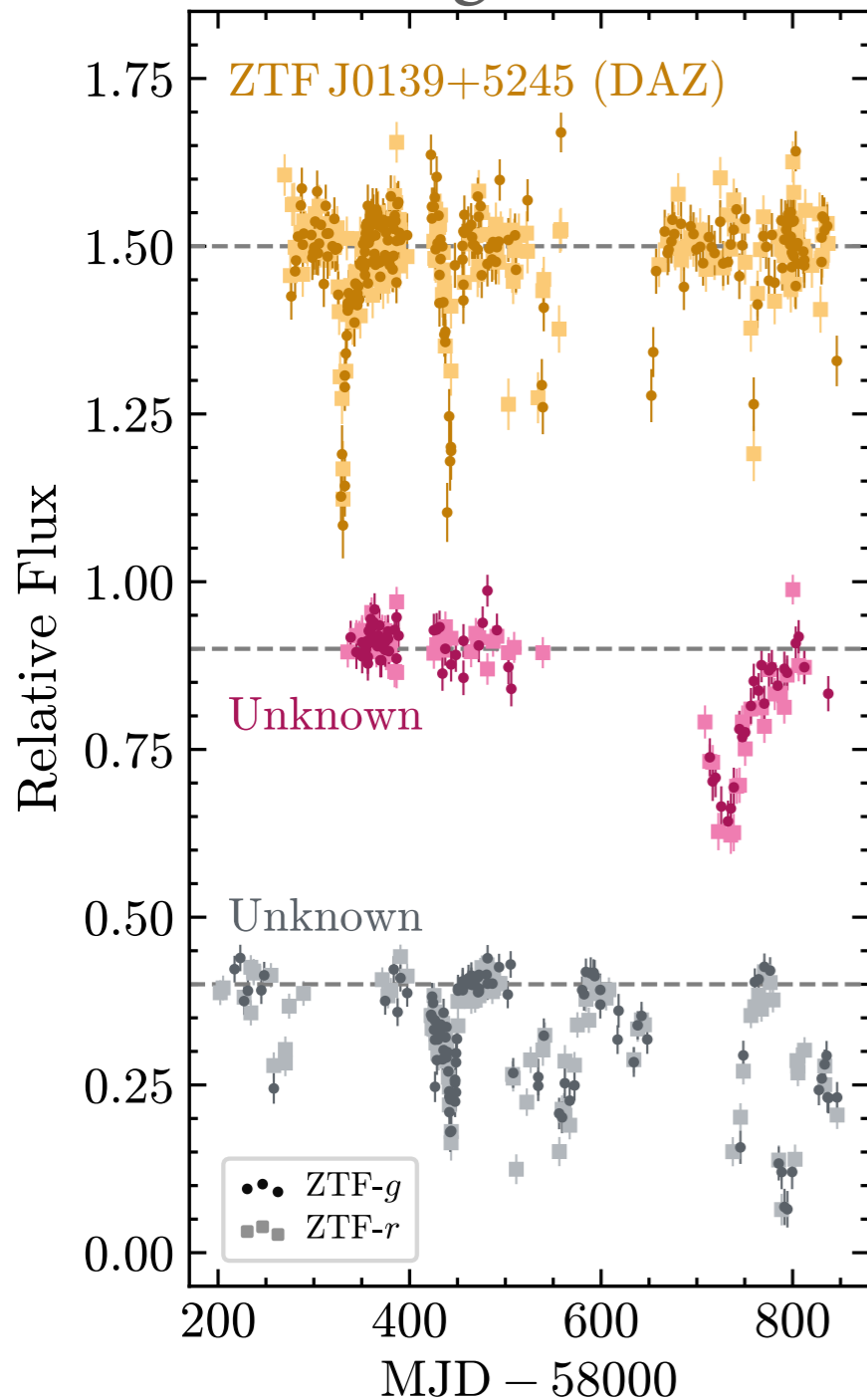
*Vanderbosch et al. 2020*

## Transit Recurrence Times Observed to Vary up to 20-days

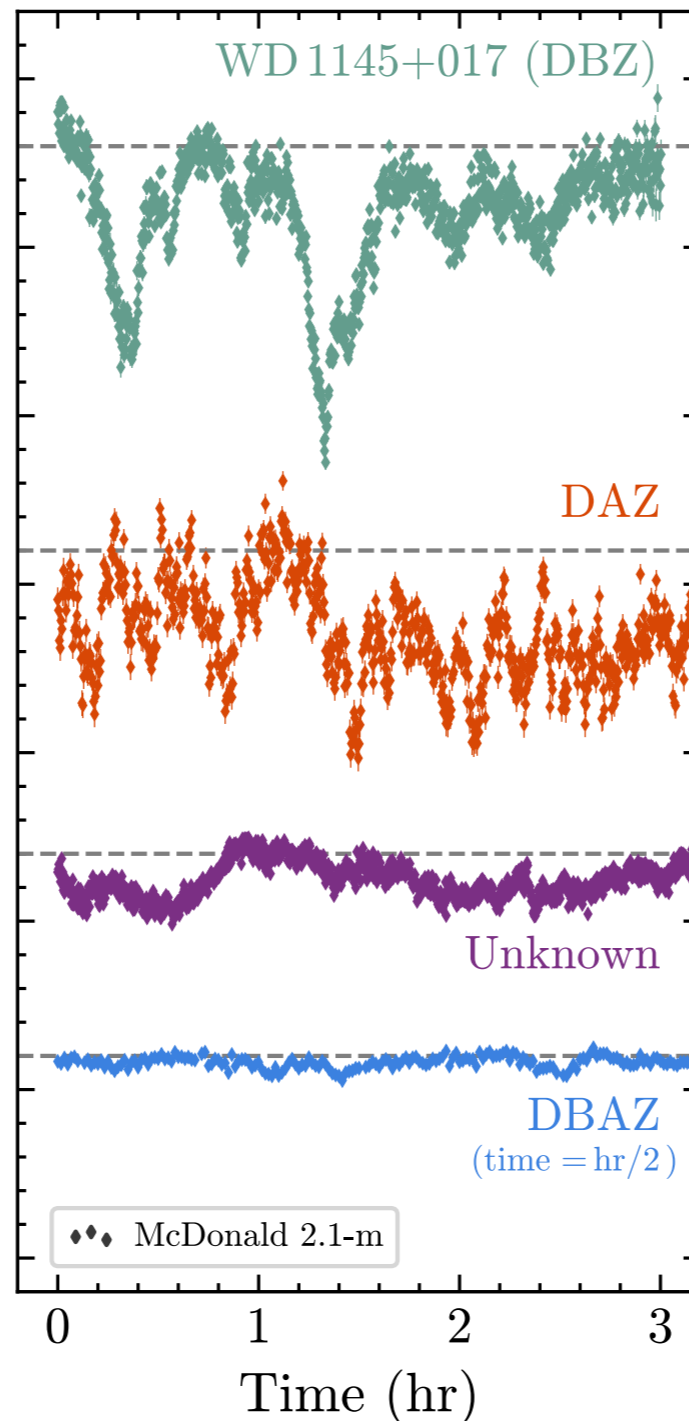
- Is the planetary debris evolving onto a closer orbit?
- Is the system actively generating new debris?
- **Ongoing photometric monitoring is required to determine whether the transit recurrence time has shrunk permanently, or if variations of this magnitude are commonplace.**
- Finding more transiting debris systems is necessary

# KNOWN AND CANDIDATE PLANETARY DEBRIS SYSTEMS

*Long Period*



*Short Period*



Transiting Debris Candidates Identified using **ZTF + Gaia Variability Metrics**

- ▶ Excess ZTF Light Curve Scatter, excess Gaia flux uncertainty, and ZTF Alerts
- ▶ No J-band IR-excess
- ▶ Spectroscopic Metal pollution
- ▶ Flux dips on short or long timescales
- ▶ Gemini FT proposal accepted to get spectral IDs for two objects

***DAZ = Metal polluted H-atmosphere White Dwarf***