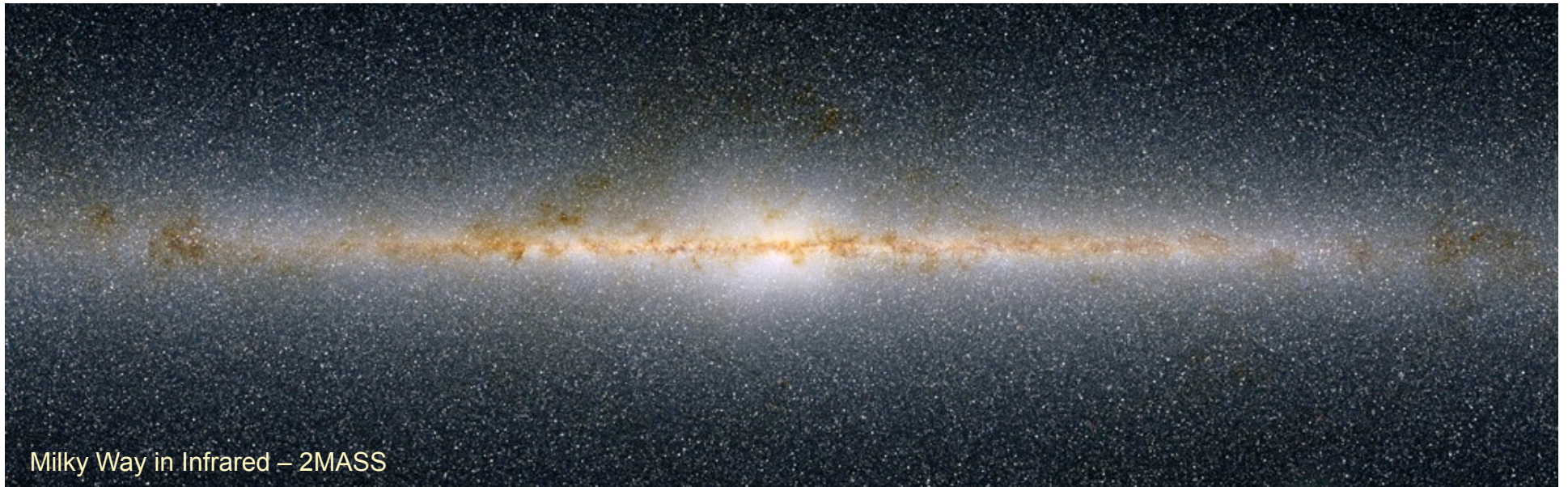


# **Astronomy 1: The Evolving Universe**

## **Spring 2014**

**Prof. S. G. Djorgovski**



**[http:// www.astro.caltech.edu/~george/ay1/](http://www.astro.caltech.edu/~george/ay1/)**

# Some Class Logistics

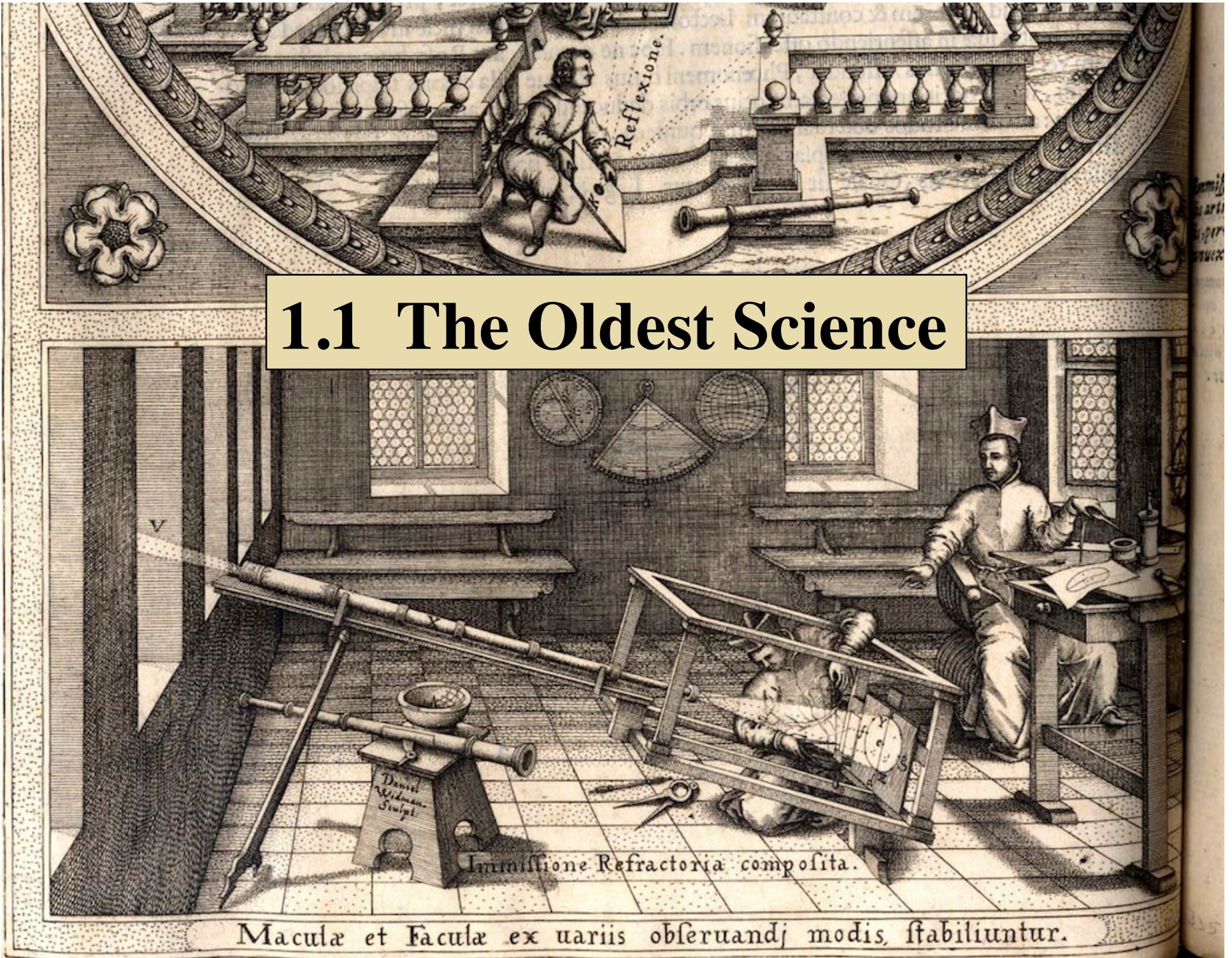
- Class website: [http:// www.astro.caltech.edu/~george/ay1/](http://www.astro.caltech.edu/~george/ay1/)
  - *Everything* will be there
- Lectures MW 2-3 pm
  - Lectures are video recorded. Videos and slides will be posted on-line the same day. Along with posted readings and links, that is your “textbook”
  - These will also be offered as a MOOC with edX, with a 3 week delay
- Sections (mandatory!): various times & locations, sign up on line
- Weekly homework, except for the midterm and final weeks
- Class trip(s) to be arranged
- **Ask questions!**

# Use the *WorldWide Telescope*

- Windows: download from <http://www.worldwidetelescope.org>
- Others: use <http://www.worldwidetelescope.org/webclient/>



# 1.1 The Oldest Science



Immissione Refractoria composita.

Macula et Facula ex uariis obseruandj modis, stabiliuntur.

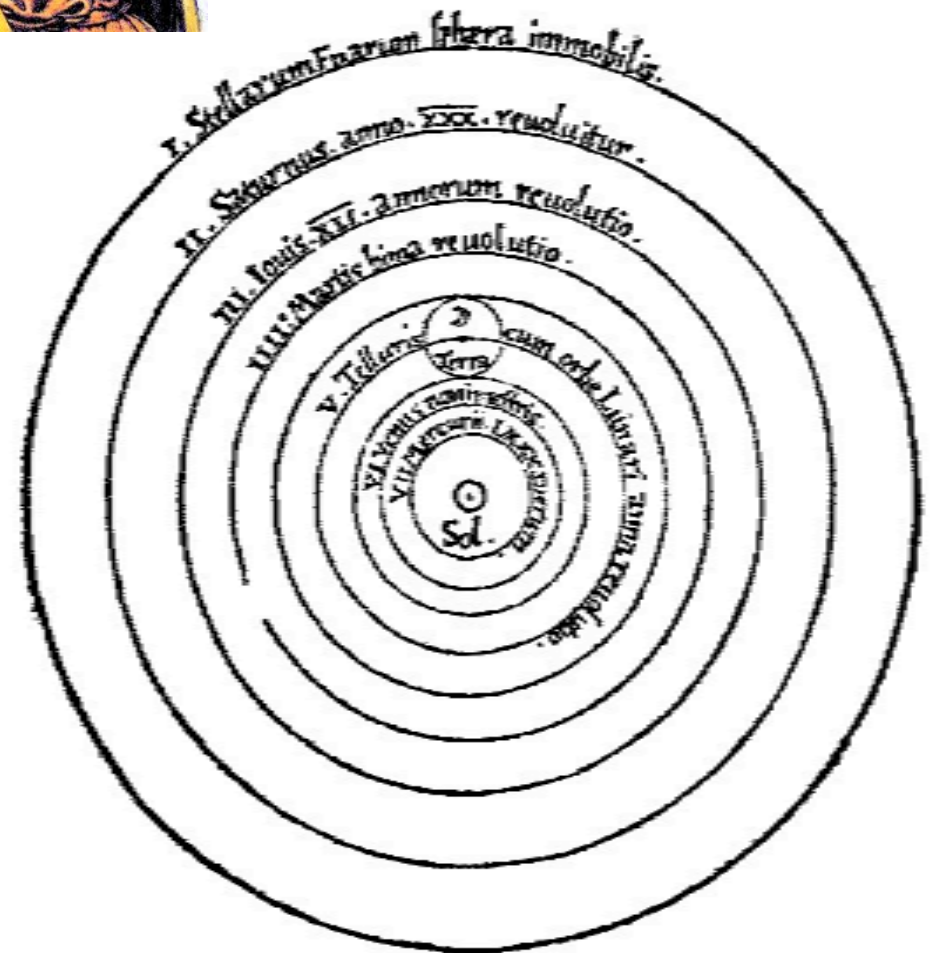
# An Early 30-m Telescope



# Ptolemaic System

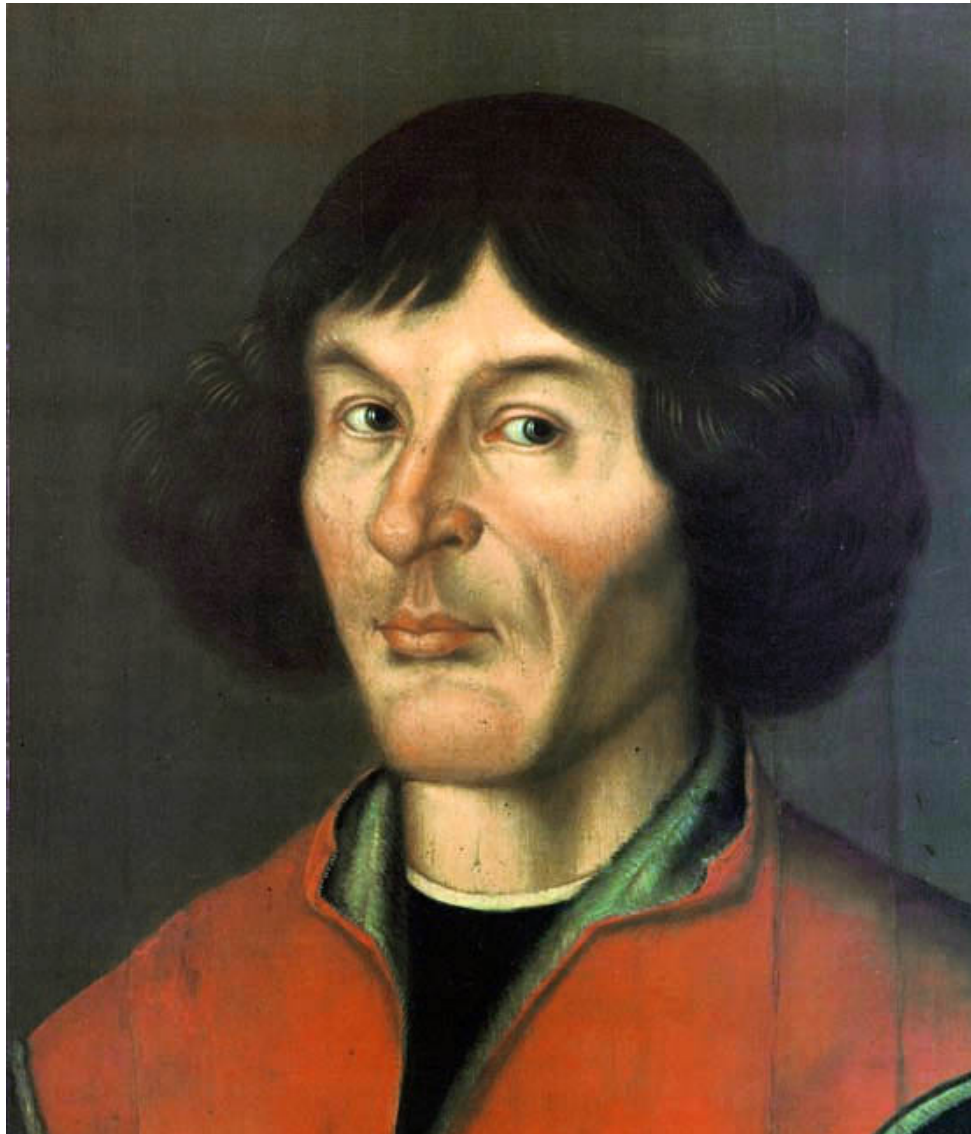


~ 1<sup>st</sup> century



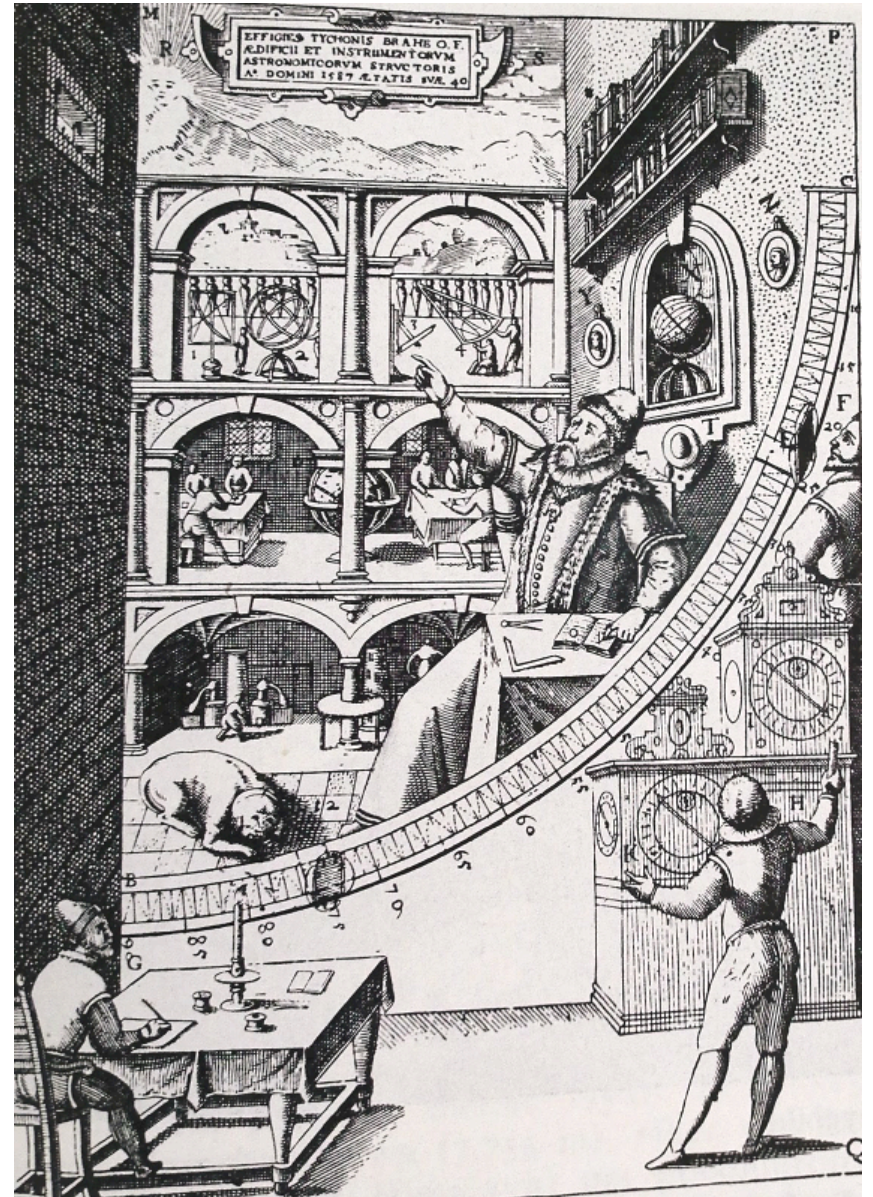


# Copernicus: De Revolutionibus Orbium Coelestium (1543)

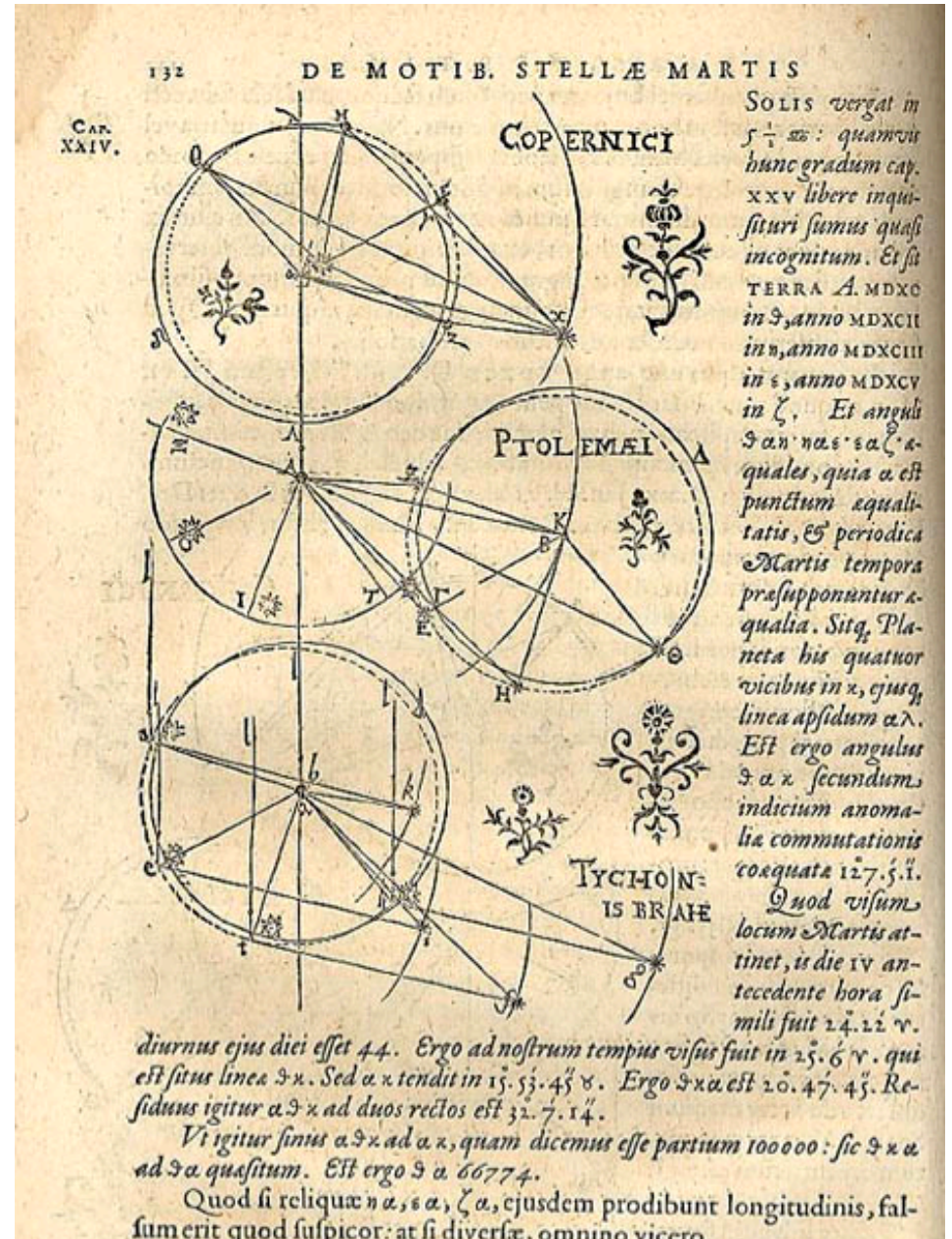




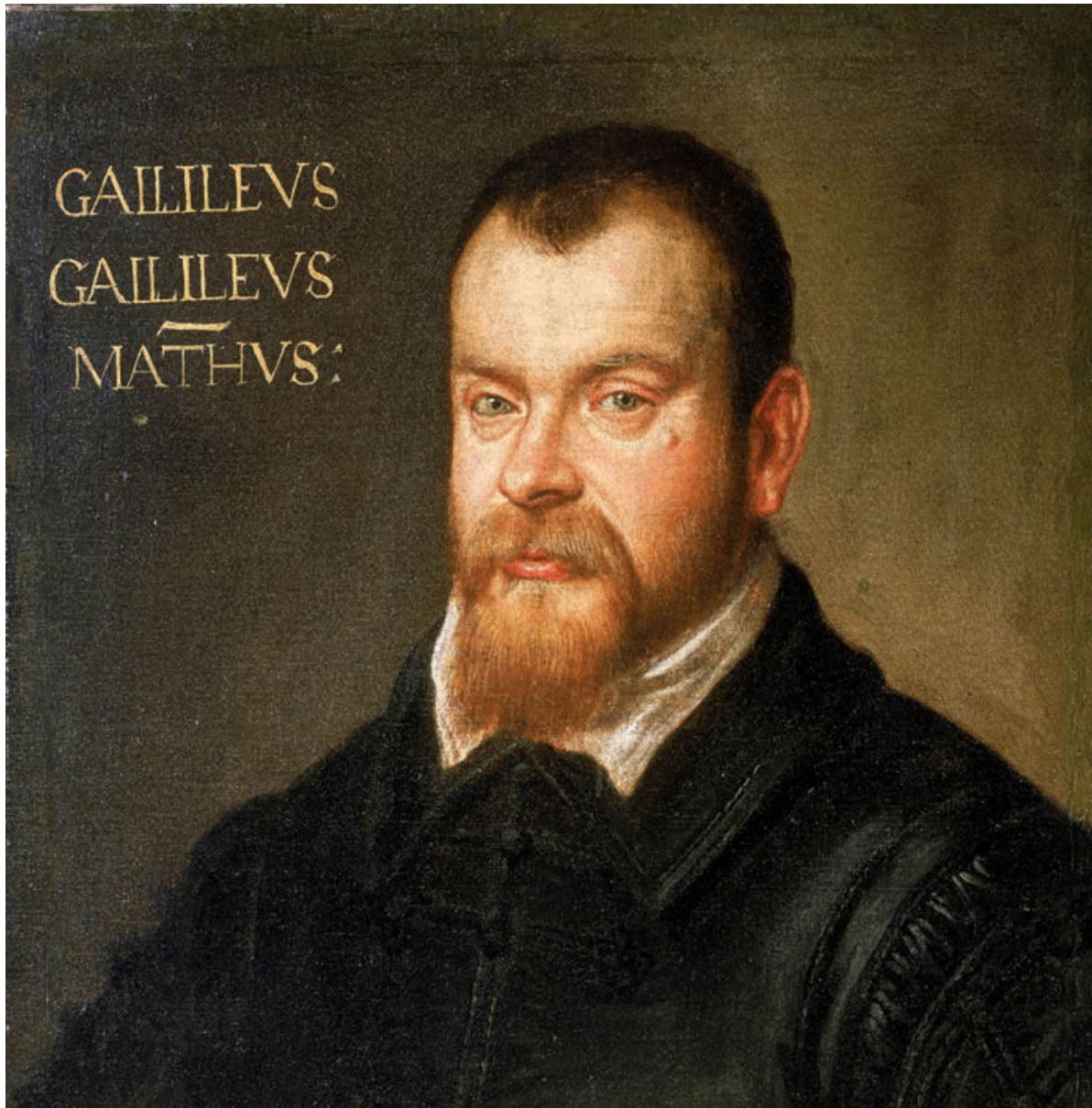
# Tycho Brahe

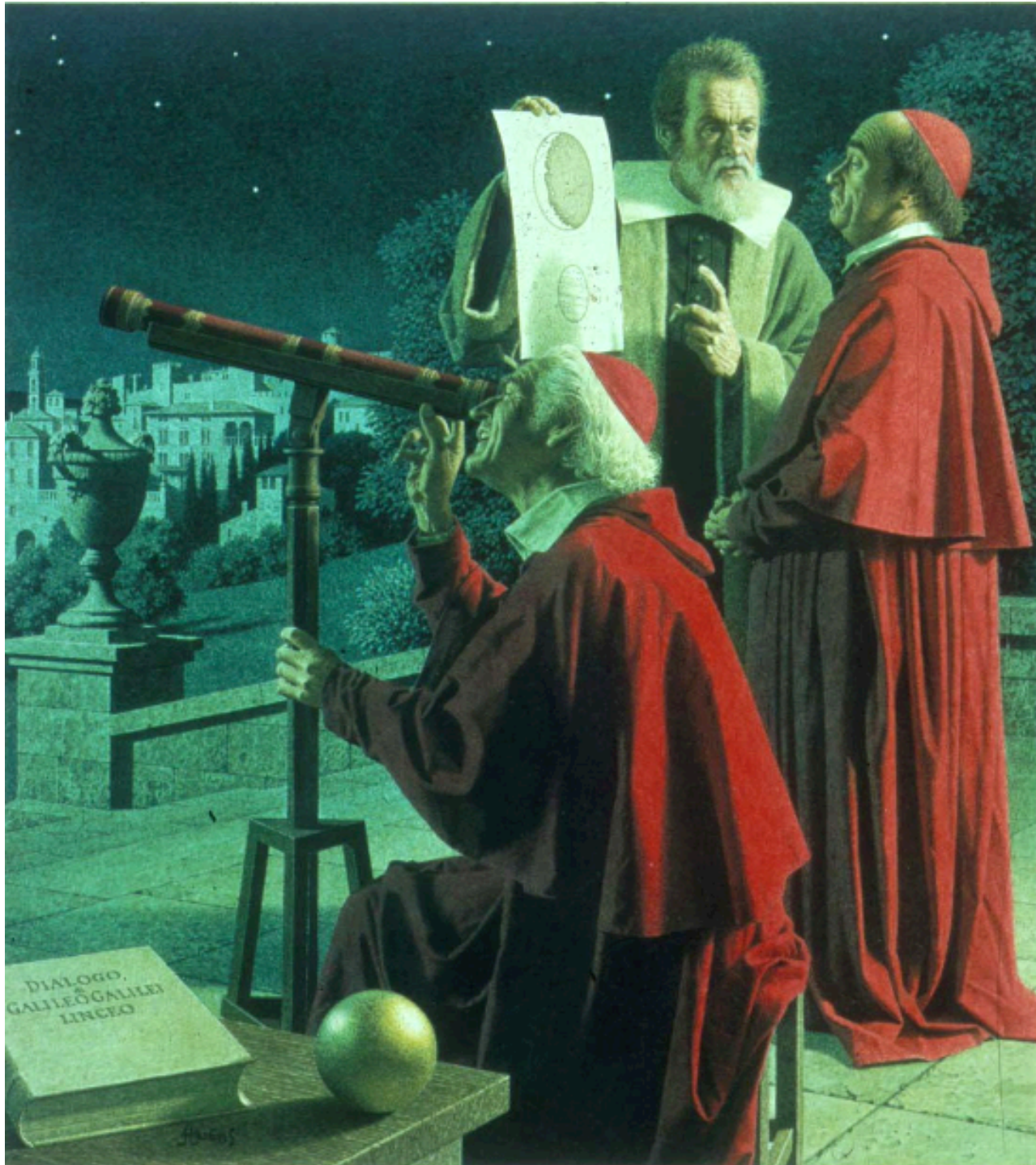


# Kepler: Astronomia Nova (1609)



# Galileo: Starry Messenger (1610), Dialogue Concerning the Two Chief World Systems(1632)





# 1.2 Astronomy as a Science



# The Vast Scale of the Universe

**If the Earth was the size of a grain of sand**

... The Sun would be **5 feet away**

... The nearest star would be **250 miles away**

... Our Galaxy would be **10 million miles across**

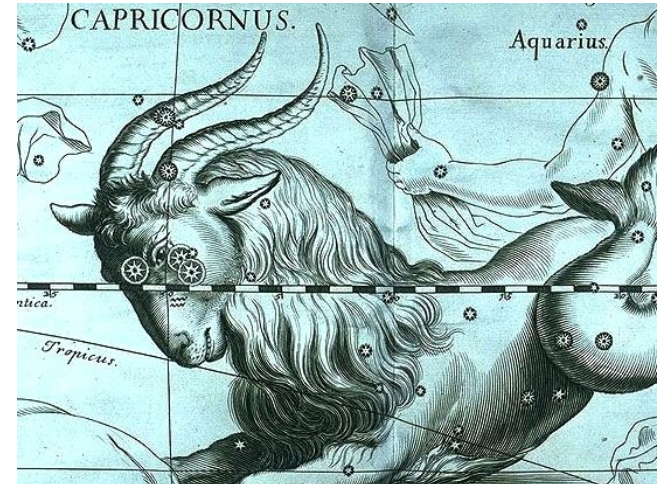
... The nearest other big galaxy (Andromeda) would be **130 million miles away**

**If our Galaxy were the size of a frisbee ...**

... The most distant objects currently known would be about **100 miles away**

# The Evolution of Astronomy

- From astrology to classical astronomy (~ positional astronomy and celestial mechanics) to astrophysics
- A strong and growing connection with physics, starting with Newton ... Today astronomy is one of the most exciting branches of physics
- Many important developments happened in Pasadena (Hale, Hubble, Zwicky, Baade, Minkowski, Sandage, ...)



# How is Astronomy Possible?

The universe is really, really big,  
and we cannot experiment in the  
lab with any of the objects in it

But we can use...

Data



+

Logic (~ math)



= Scientific method

... and use physics as an  
interpretative framework



# The Nature of the Astronomical Inquiry

- The peculiar nature of astronomy as a science
  - Is it like history? Geology? Paleontology? (are there extinct species of astronomical objects?)
  - Observing vs. experiments, and repeatability
  - A single object of study: universe as a whole, CMBR...  
But the experiments are repeatable
  - Non-repeatable phenomena, e.g., SNe, GRBs, microlensing events... But there are *classes* of them
- Observing a narrow time-slice of the past light cone
  - Using “symmetry” principles (e.g., Copernican, cosmological) as a substitute for unobtainable information
  - $t(\text{astronomy}) \ll t(\text{universe}) \rightarrow$  inevitable biases
- Observing the past, or deducing it from the “fossil” information (e.g., galaxy formation and evolution)

# Astronomy as a Branch of Physics

- Using the apparatus of physics to gather and interpret the data: assume that our physics is universal (and we can test that!)
- Astronomical phenomena as a “cosmic laboratory”
  - Relativistic physics (black holes, high  $\Gamma$ , ...)
  - Cosmic accelerators (HECR) and the early universe
  - Matter in extreme conditions (e.g., neutron/quark stars, GRBs, high & low density plasmas ...)
- Astronomical discoveries as a gateway to the new physics (e.g., dark matter and dark energy; neutrino mixing; inflation; etc.)
- Progress driven by technology (telescopes, detectors, computing...)

# Fundamental Limits to Measurements and Selection Effects

- S/N Poissonian and quantum limits of detection
- Geometrical optics limits of angular resolution
- Opacity of the Earth's atmosphere and the Galactic ISM (example: soft X-rays and the missing baryons)
- Obscuration by dust in galaxies
- Turbulence of the atmosphere/ISM: erasing the spatial information
- Convolved backgrounds and foregrounds (examples: CMBR, CIRBs)
- And the “un-natural” limits: politics, funding, social psychology ...

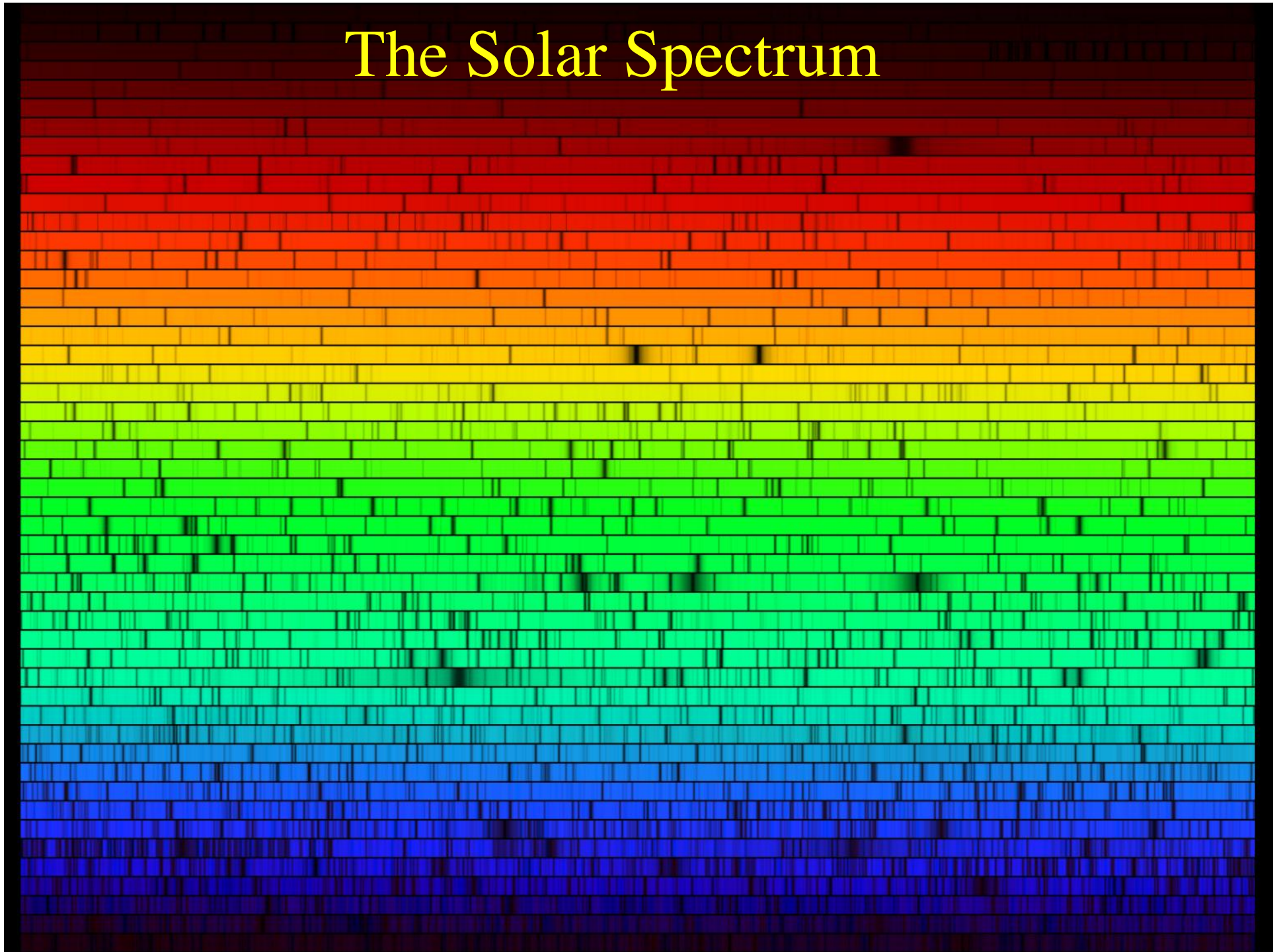
# 1.3 Messengers from the Universe



# Information Flows in the Universe

- Physical parameters → Observables (but possibly in a very convoluted manner - complex phenomena)
- Unresolved imagery/photometry: a very low information content; resolved imagery: morphology
- Spectroscopy is where most of the physics is!
- Primary continuum spectra (thermal, synchrotron...) : a low information content; abs./em. lines encode most of the interesting information
- Thermalization by dust erases information from the original energy flux (e.g., the power sources of ULIRGS)
- Different phenomena → different signals (some spectrum regions may be favored)

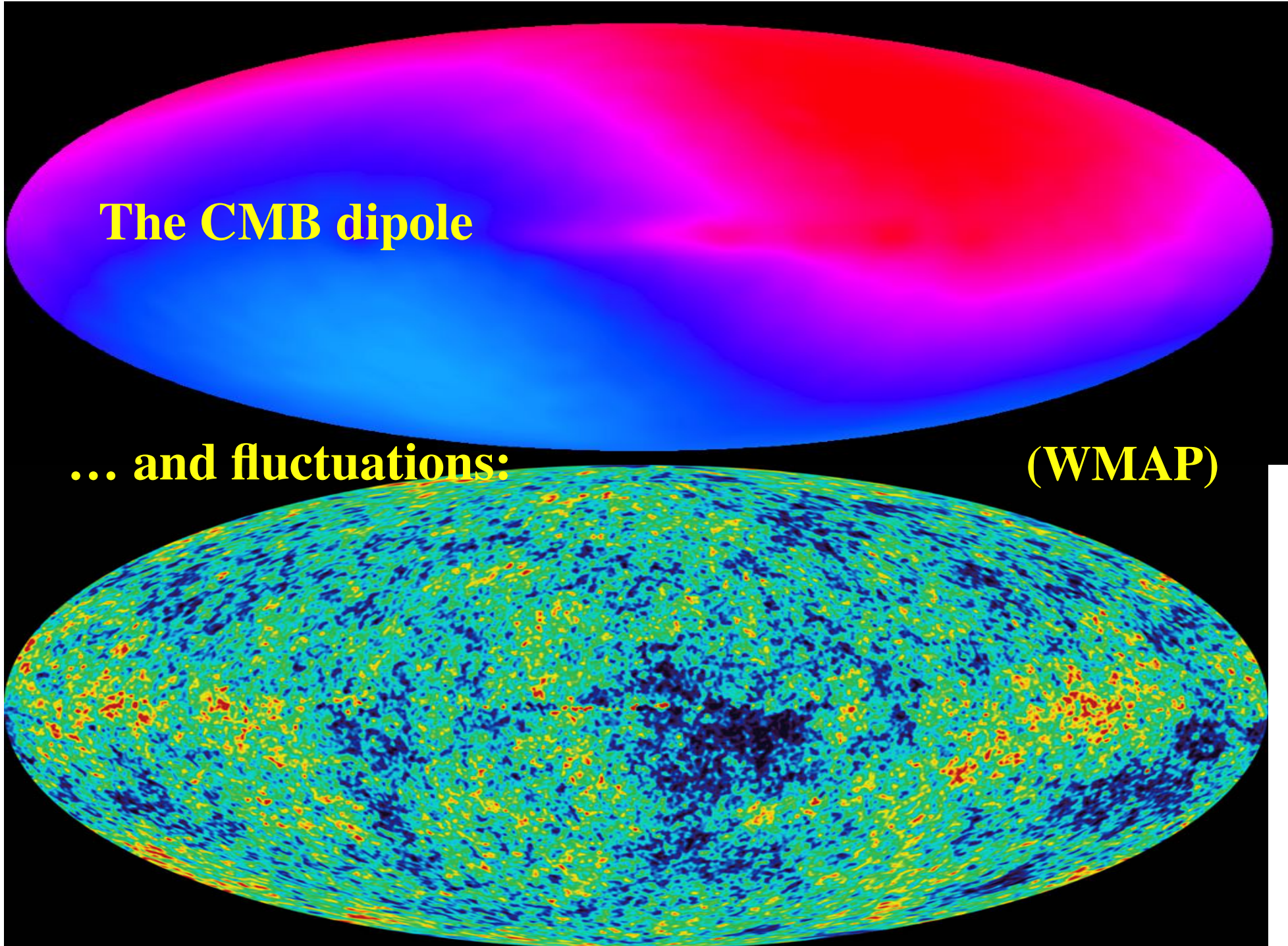
# The Solar Spectrum



**The CMB dipole**

**... and fluctuations:**

**(WMAP)**

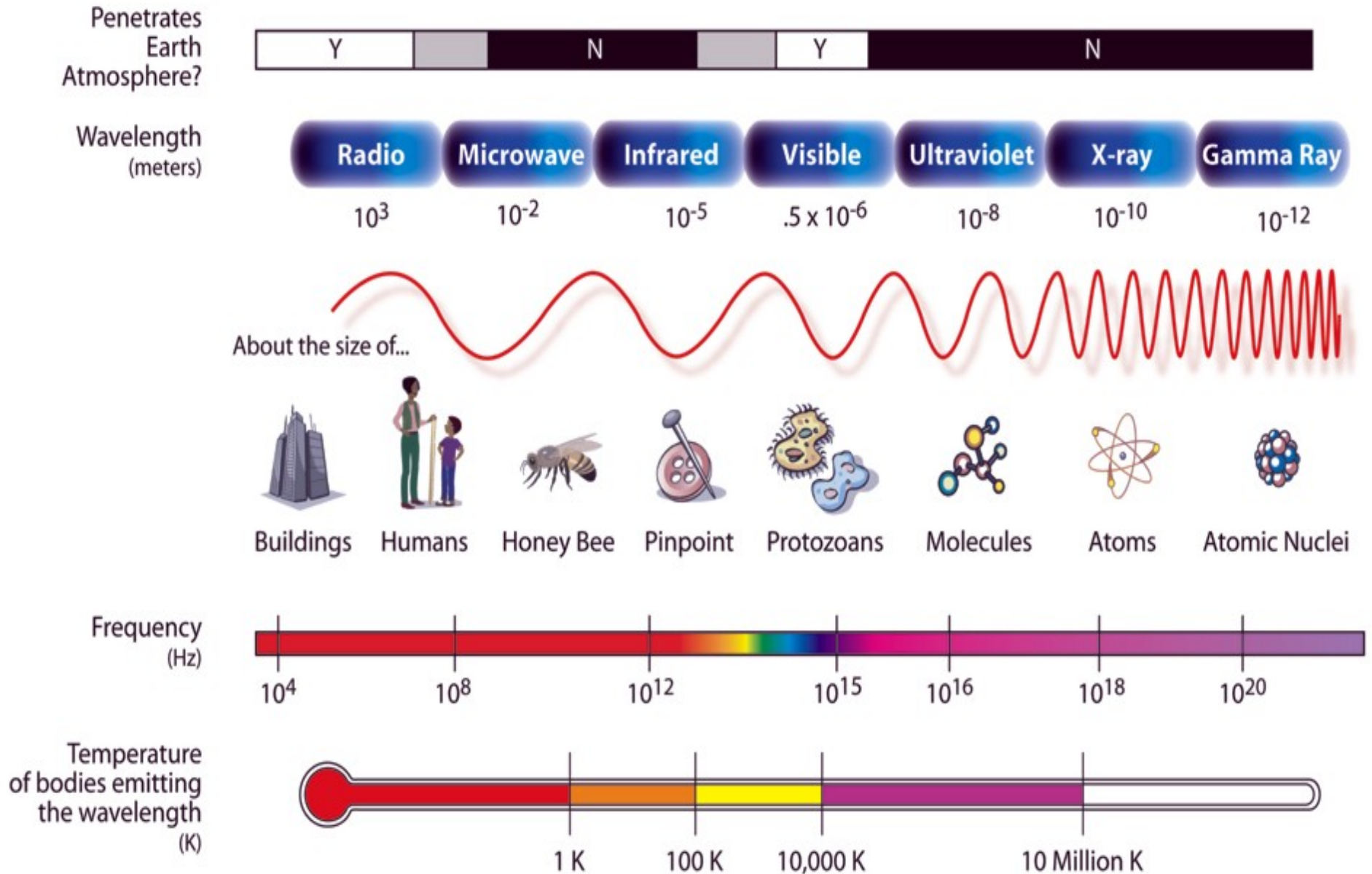


# Information Channels in Astronomy

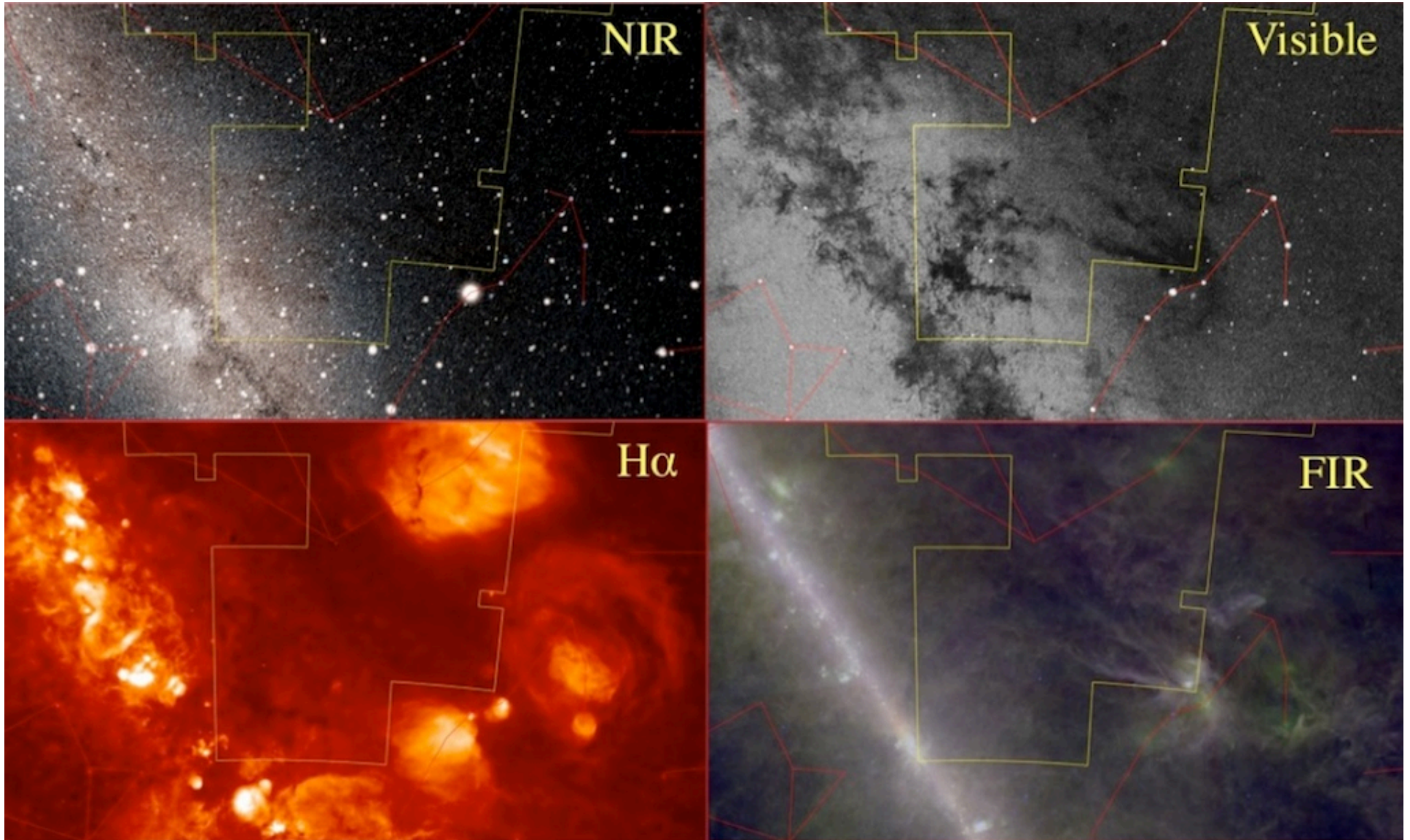
- Mostly electromagnetic! Methodologies:
  - Single-channel photometry
  - 2D imaging (photometry, morphology, positions/motions)
  - 1D spectroscopy
  - 2D (long-slit) spectroscopy
  - 3D data cubes (2 spatial + 1 spectro)
  - All can include polarimetry
  - All can be time-resolved (synoptic) or not
  - All can be single-dish, some (all?) can be interferometric
- Particles:
  - Cosmic rays: Cherenkov, particle detectors, geochemistry
  - Neutrinos: big tanks of something ...
- Gravity Waves: LIGO/LISA type interferometers
- Dark Matter: lab detectors, gravitational lensing



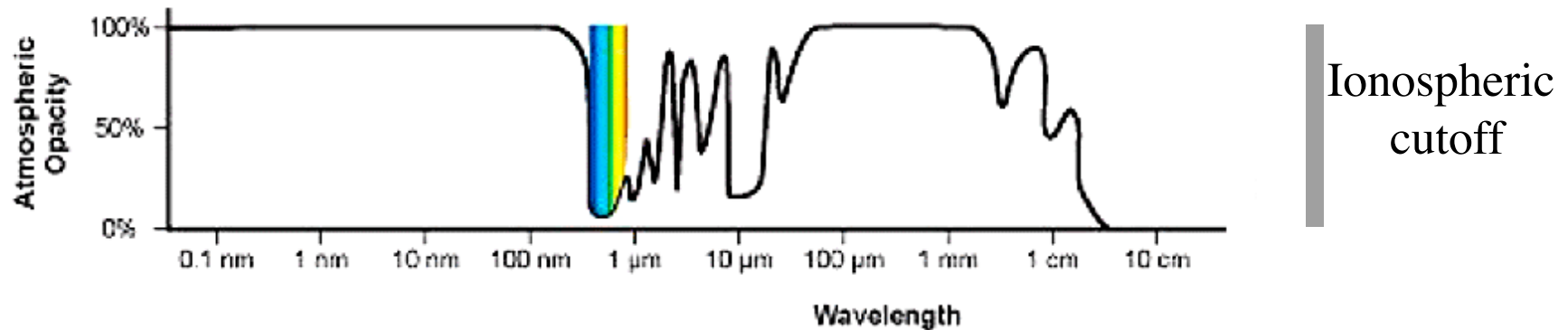
# The Electromagnetic Spectrum



# The Panchromatic Universe



# Atmospheric Transmission Windows



And that is why we need space observatories!

**But there is an even more profound limitation:**

The Galactic “atmosphere” - the interstellar medium - also absorbs very long wavelengths, and hard UV / soft X-rays (the interstellar fog); and of course the dust absorbs the blue/UV light (the interstellar smog).

This may be very important: perhaps 90% of the baryons in the universe are in the form of a “warm” ( $T \sim 10^5$  K) gas, which emits mostly soft X-rays

# How Are Discoveries Made?

- **Conceptual Discoveries:** e.g., Relativity, QM, Inflation ...  
*Theoretical, may be inspired by observations*
- **Phenomenological Discoveries:** e.g., Dark Matter, QSOs, GRBs, CMBR, Extrasolar Planets, Obscured Universe ...  
*Empirical, inspire theories, can be motivated by them*



## **Phenomenological discoveries are made by:**

- Pushing along some parameter space axis
- Making new connections (e.g., multi- $\lambda$ )

Different astrophysical phenomena populate different parts of the observable parameter space, and require different observables and measurement methodologies - and vice versa.