

# New Astronomy With a Virtual Observatory

**S. G. Djorgovski**  
*(Caltech)*

With special thanks to  
Roy Williams, Alex Szalay,  
Bob Hanisch, Dave De Young,  
Jim Gray, and many other VO founders ...

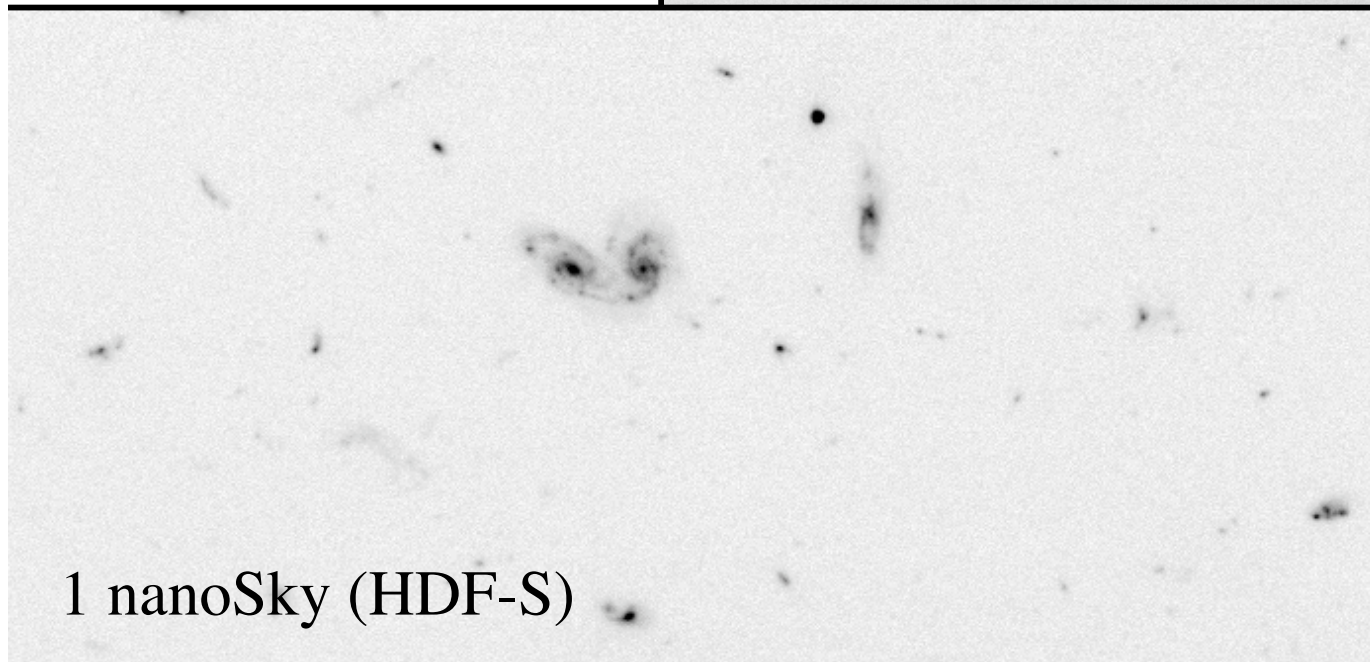
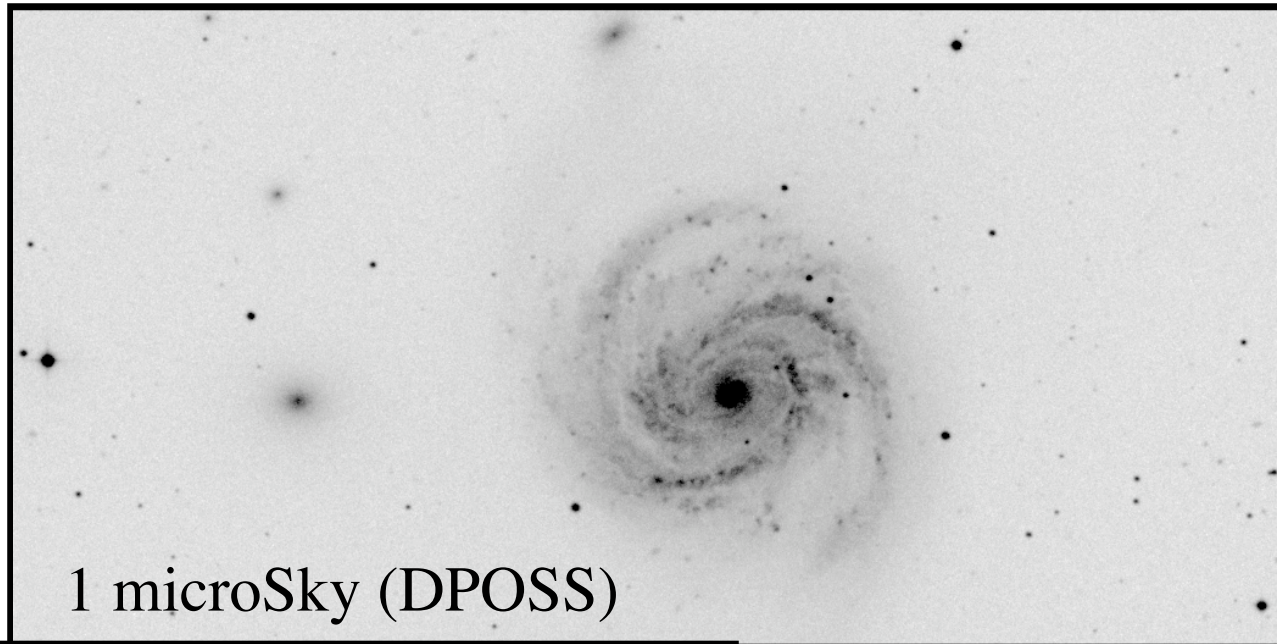


# An Overview:

- Astronomy in the era of information abundance
  - The IT revolution, challenges and opportunities
- The Virtual Observatory concept
  - What is it, how it all got started
- Virtual Observatory status
  - Where are we now, where are we going
- From technology to science (and back)
  - ... and some musings on cyber-science in general
- Concluding comments

# Astronomy is Facing a Major Data Avalanche:

Multi-Terabyte  
(soon: multi-PB)  
sky surveys and  
archives over a  
broad range of  
wavelengths ...

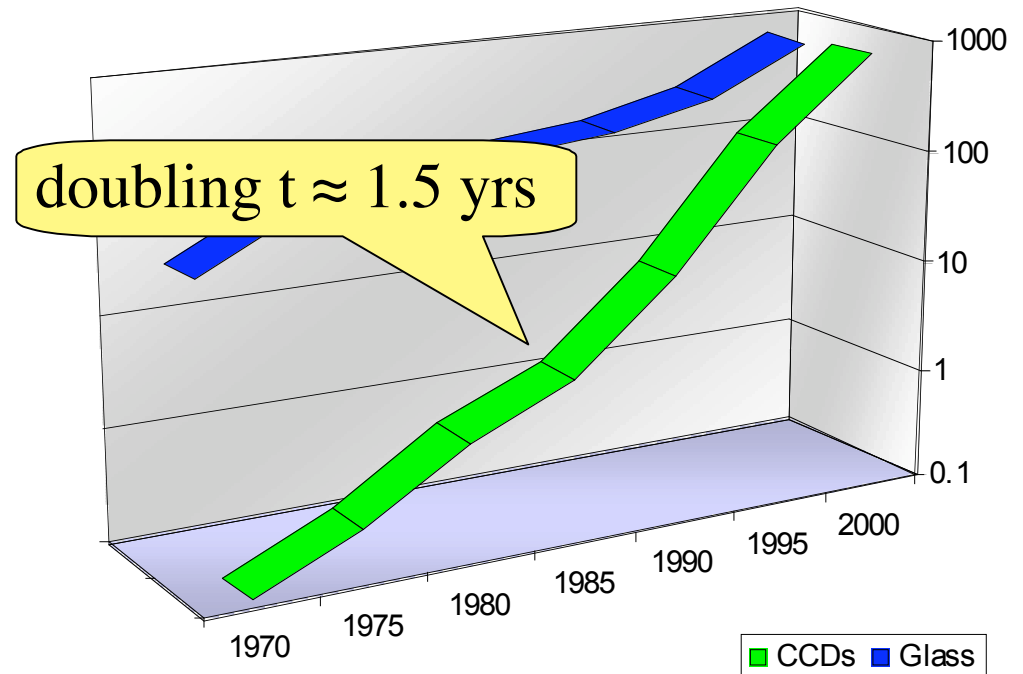


Billions of  
detected  
sources,  
hundreds of  
measured  
attributes  
per source ...

- **Large digital sky surveys** are becoming the dominant source of data in astronomy:  $\sim 10\text{-}100$  TB/survey (soon PB),  $\sim 10^6 - 10^9$  sources/survey, many wavelengths...
- **Data sets many orders of magnitude larger, more complex, and more homogeneous than in the past**

## Data $\rightarrow$ Knowledge ?

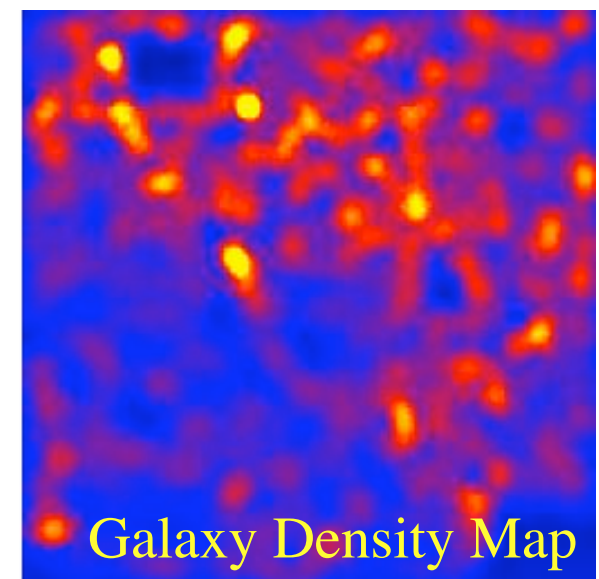
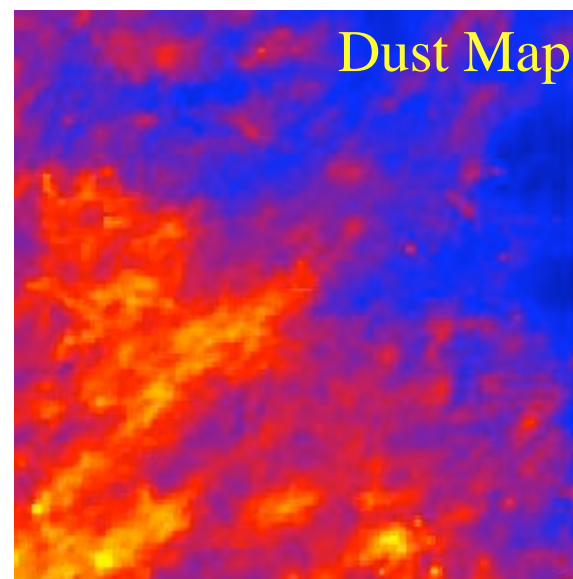
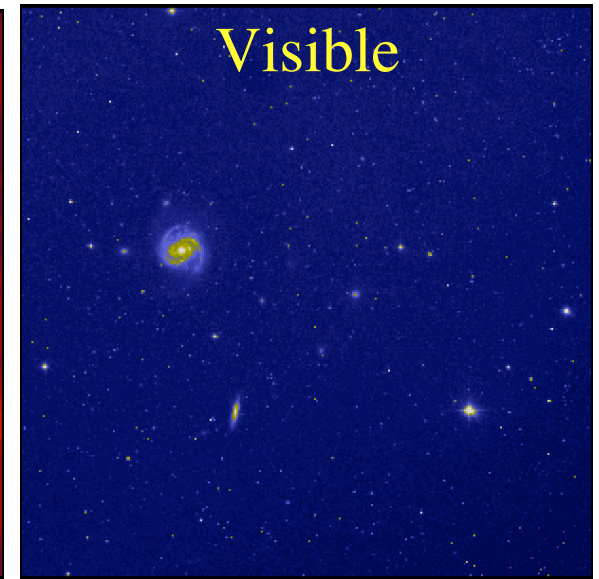
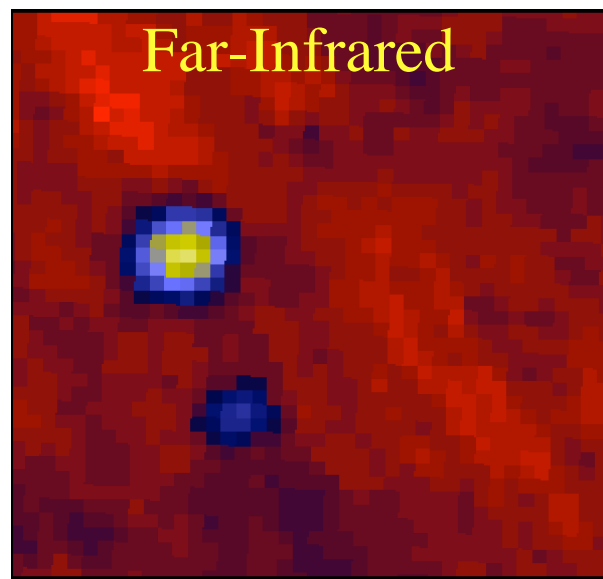
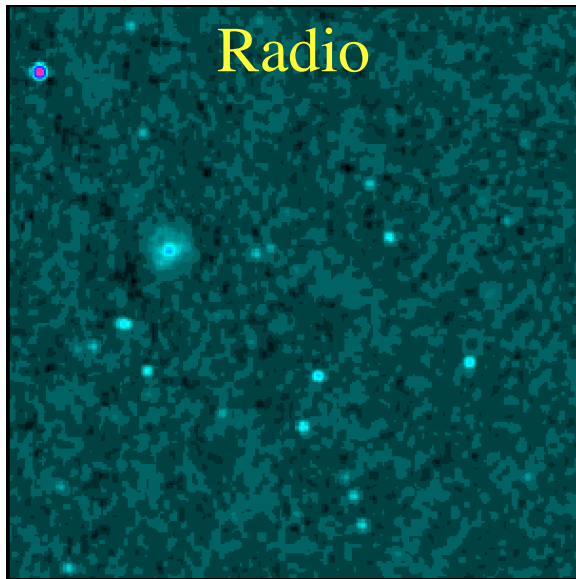
The exponential growth of data volume (and also complexity, quality) driven by the exponential growth in detector and computing technology



**... but our understanding of the universe increases much more slowly!**

# Panchromatic Views of the Universe:

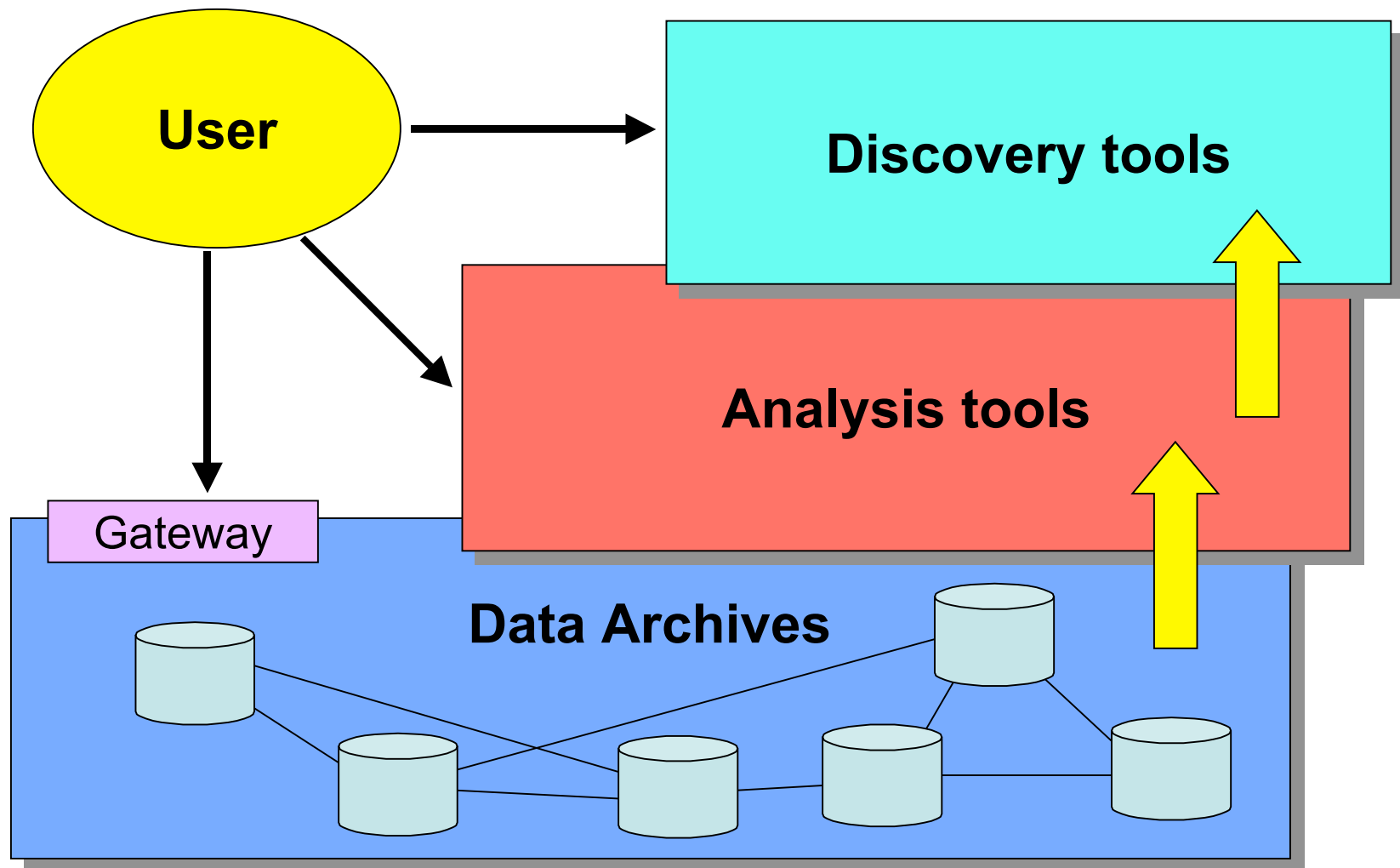
Data Fusion → A More Complete, Less Biased Picture



# The Virtual Observatory Concept

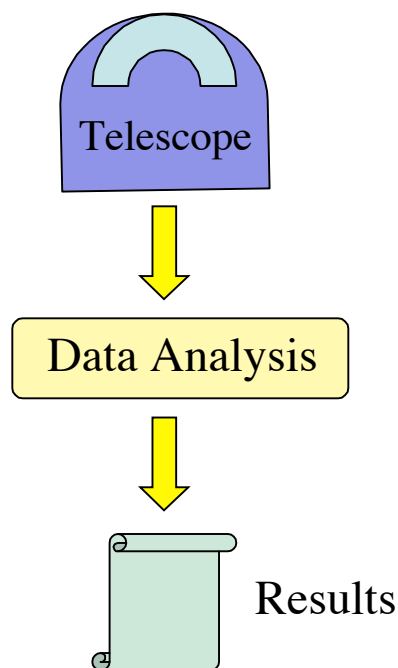
- Astronomy community response to the scientific and technological challenges posed by massive data sets
  - Highest recommendation of the NAS Decadal Astronomy and Astrophysics Survey Committee ⇒ **NVO**
  - International growth ⇒ **IVOA**
- A complete, dynamical, distributed, open *research environment for the new astronomy with massive and complex data sets*
  - Provide content (data, metadata) services, standards, and analysis/compute services
  - Federate the existing and forthcoming large digital sky surveys and archives, facilitate data inclusion and distribution
  - Develop and provide data exploration and discovery tools
  - *Technology-enabled, but science-driven*

# VO: Conceptual Architecture

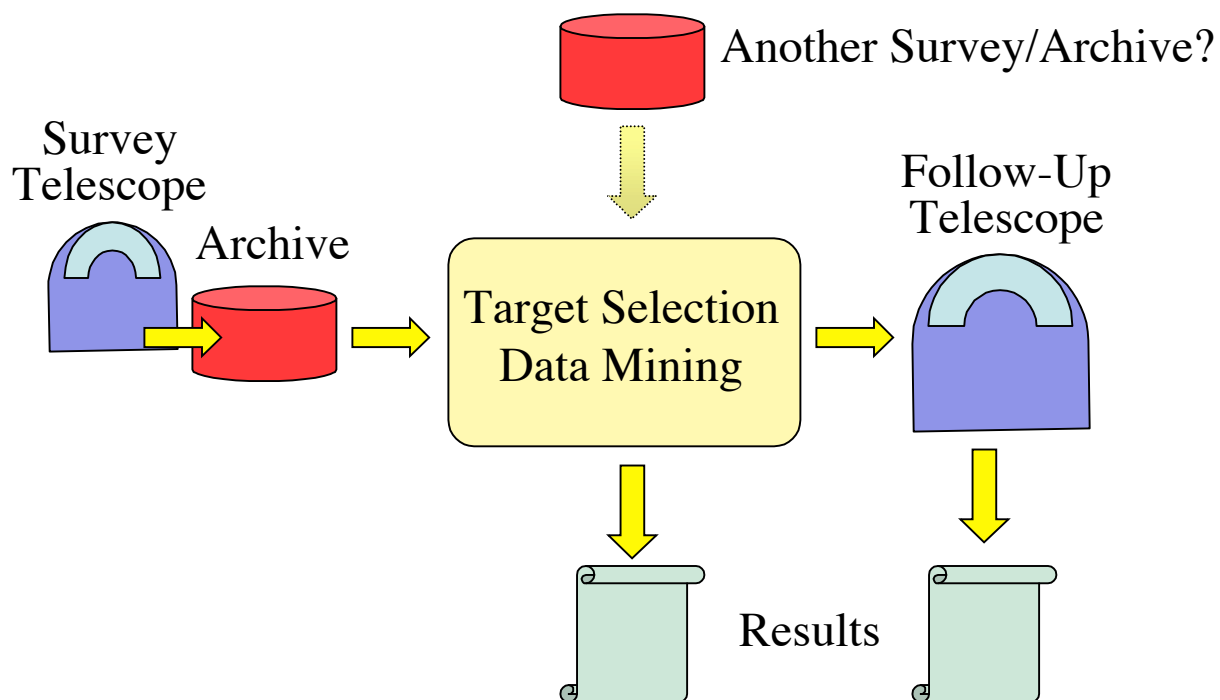


# From Traditional to Survey to VO-Based Science

Traditional:



Survey-Based:



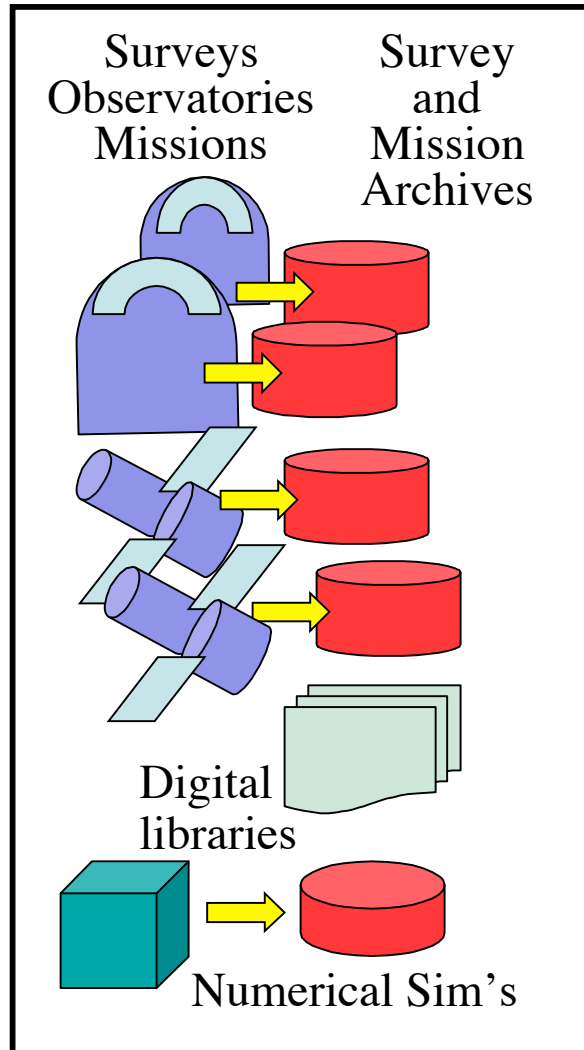
Highly successful and increasingly prominent, but inherently limited by the information content of individual surveys ...

*What comes next, beyond survey science is the VO science*

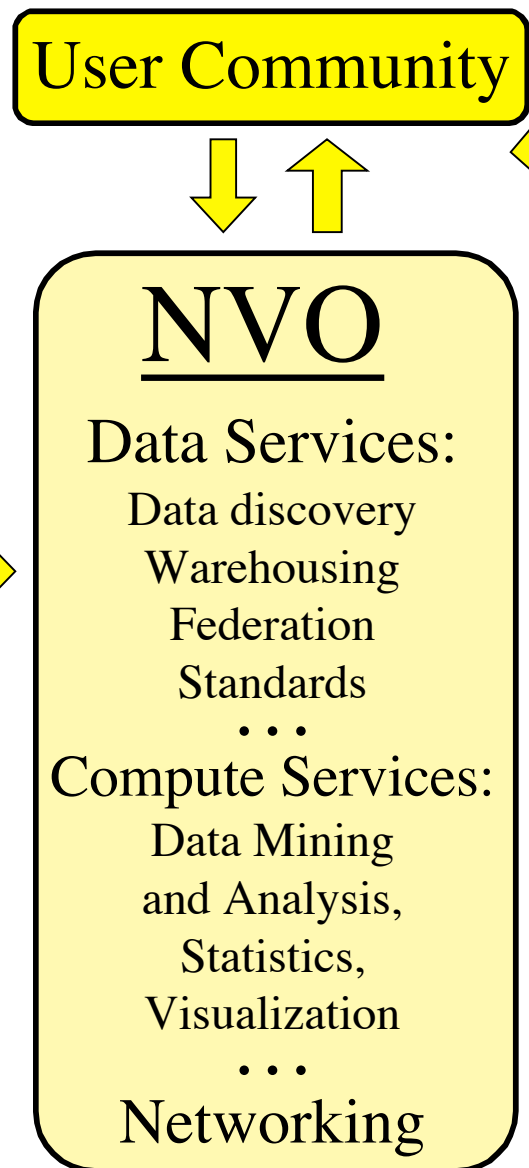


# A Systemic View of the NVO

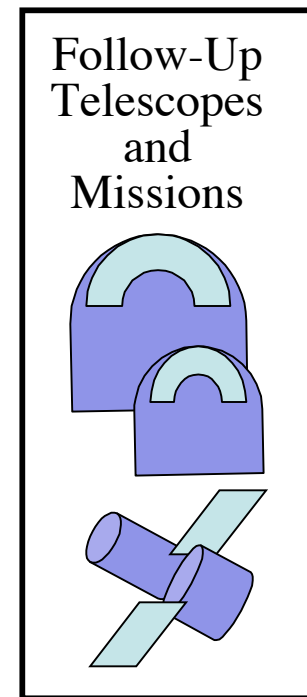
## Primary Data Providers



## User Community



## Secondary Data Providers



## International VO's



# Why is VO a Good Scientific Prospect?

- Technological revolutions as the drivers/enablers of the bursts of scientific growth
- Historical examples in astronomy:
  - 1960's: the advent of electronics and access to space  
*Quasars, CMBR, x-ray astronomy, pulsars, GRBs, ...*
  - 1980's - 1990's: computers, digital detectors (CCDs etc.)  
*Galaxy formation and evolution, extrasolar planets, CMBR fluctuations, dark matter and energy, GRBs, ...*
  - **2000's and beyond: information technology**

*The next golden age of discovery in astronomy?*



**VO is the mechanism to effect this process**

# How and Where are Discoveries Made?

- **Conceptual Discoveries:** e.g., Relativity, QM, Strings/Branes, Inflation ... *Theoretical, may be inspired by observations*
- **Phenomenological Discoveries:** e.g., Dark Matter, Dark Energy, QSOs, GRBs, CMBR, Extrasolar Planets, Obscured Universe ... *Empirical, inspire theories, can be motivated by them*

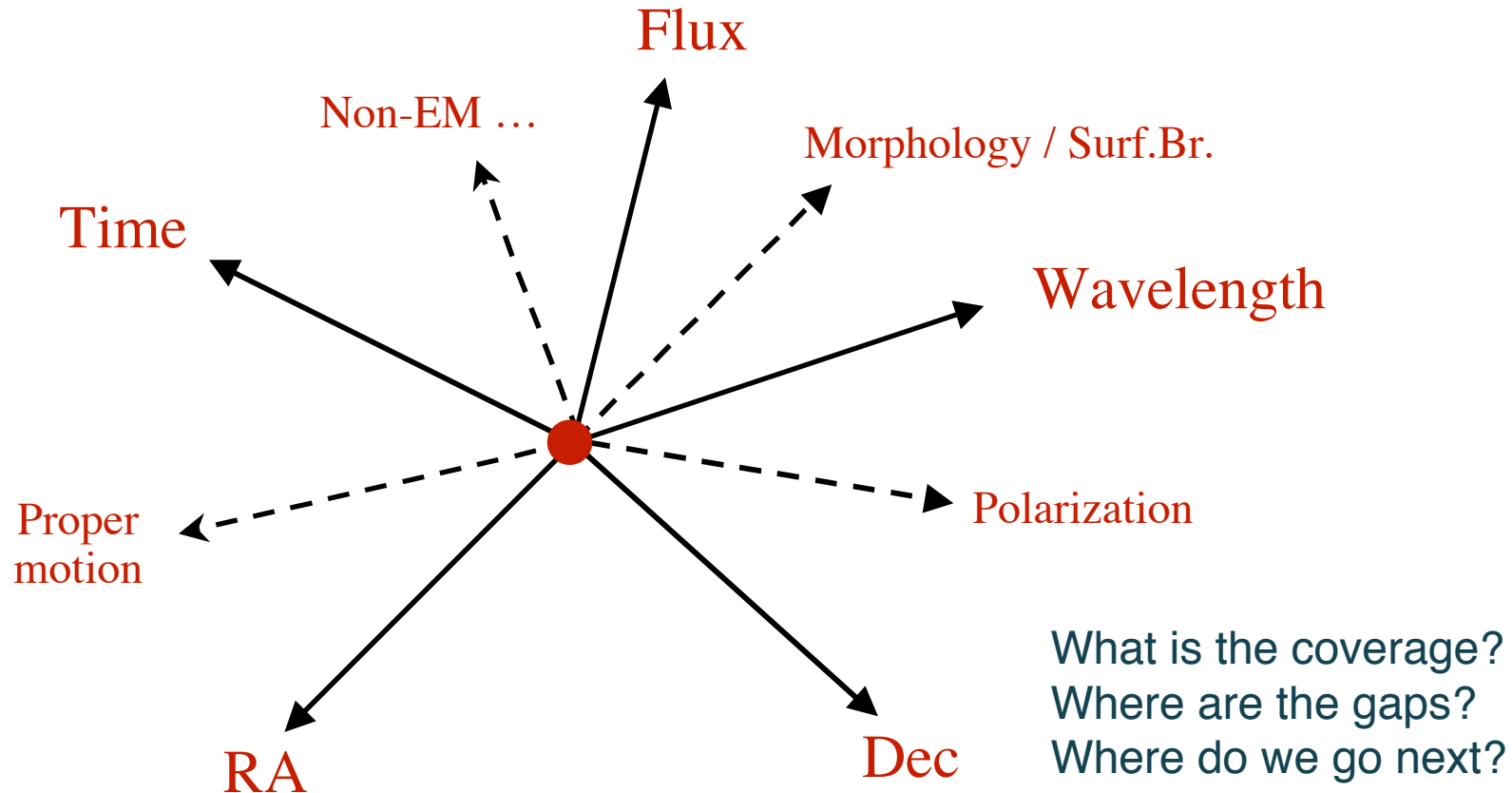


## **Phenomenological Discoveries:**

- Pushing along some parameter space axis  VO useful
- Making new connections (e.g., multi-□)  **VO critical!**

*Understanding of complex astrophysical phenomena requires complex, information-rich data (and simulations?)*

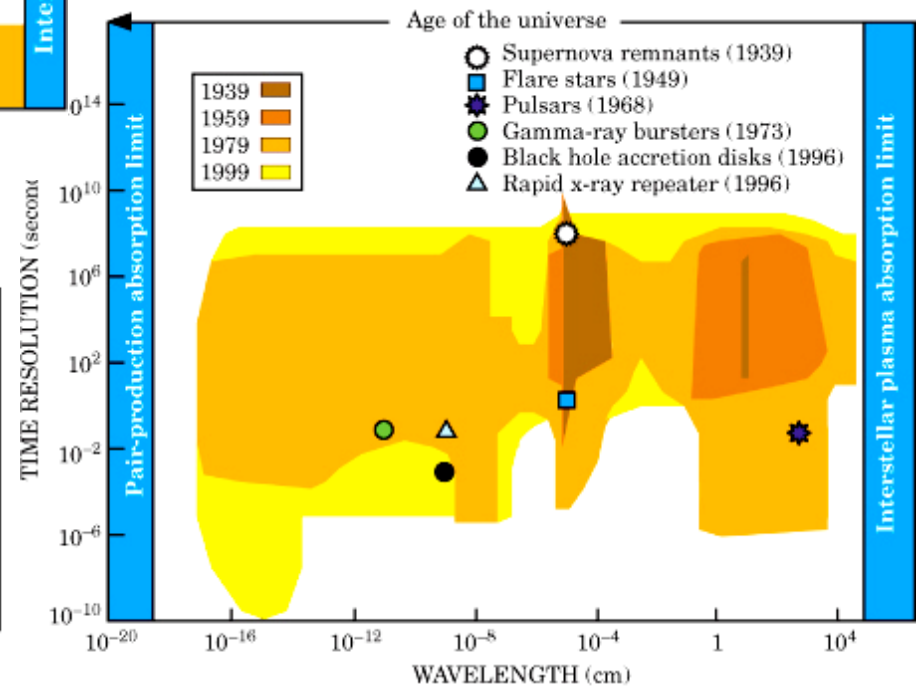
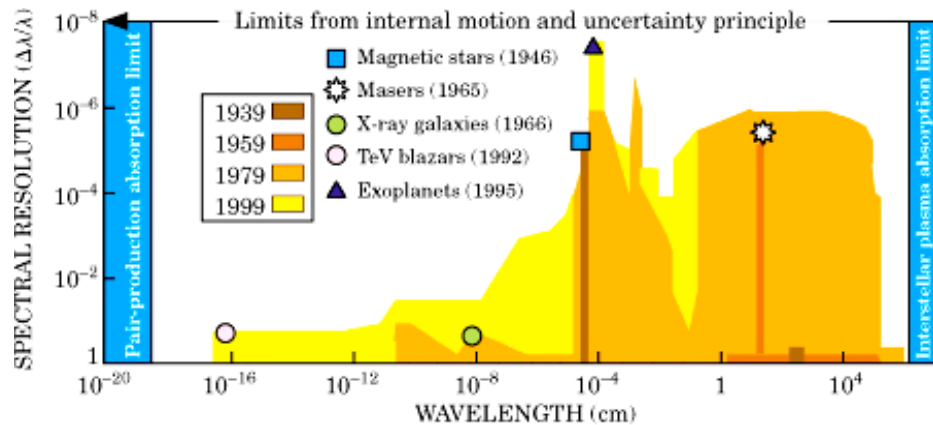
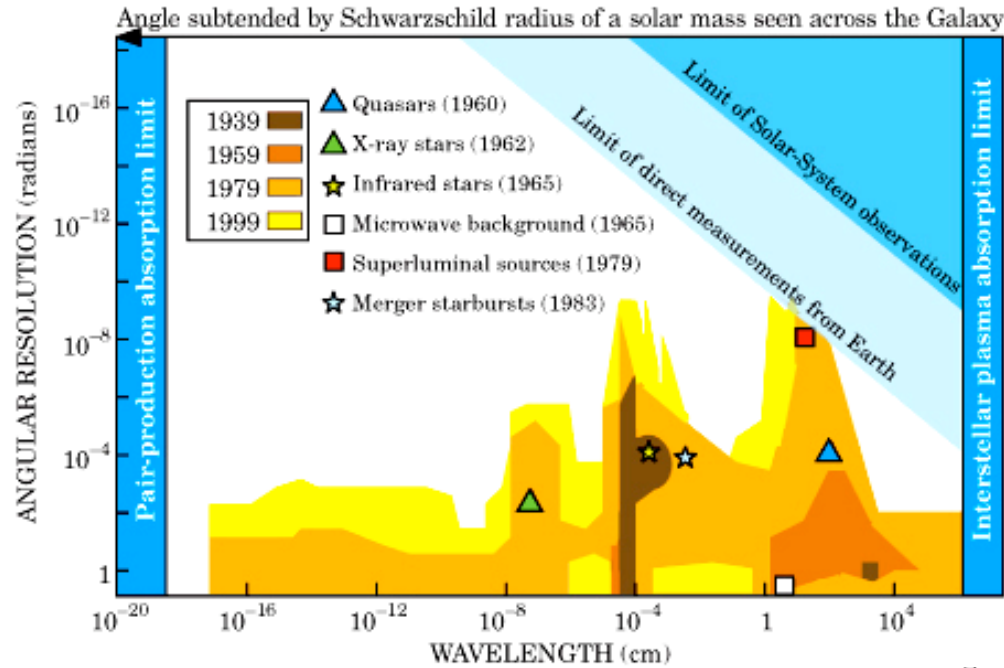
# Taking a Broader View: The Observable Parameter Space



Along each axis the measurements are characterized by the **position, extent, sampling and resolution**. All astronomical measurements span some volume in this parameter space.

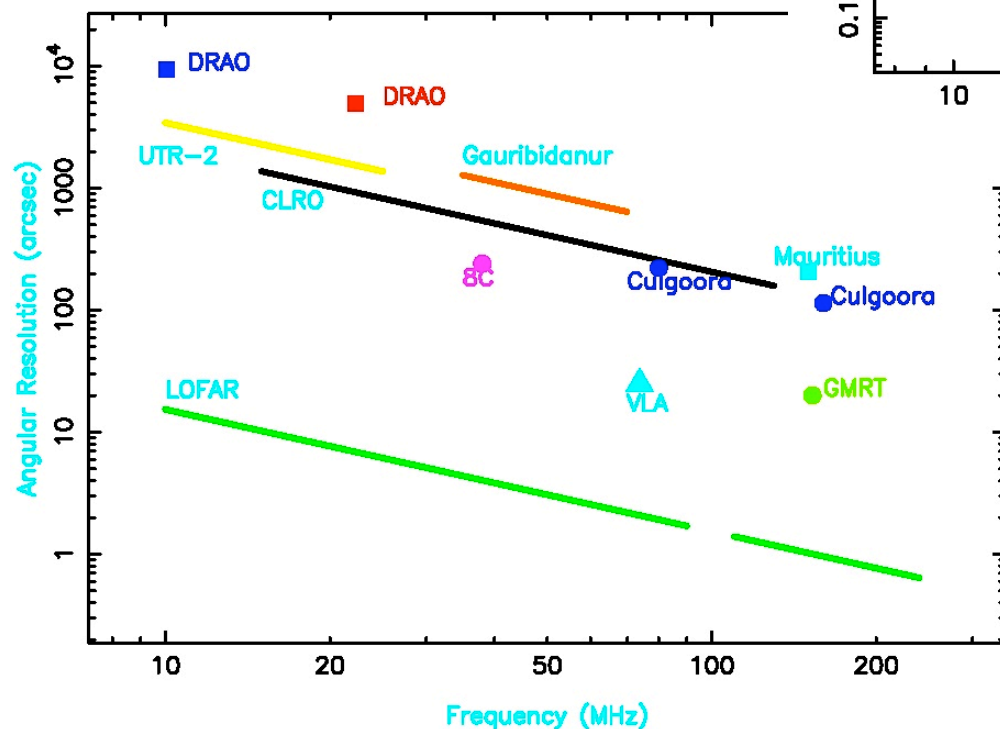
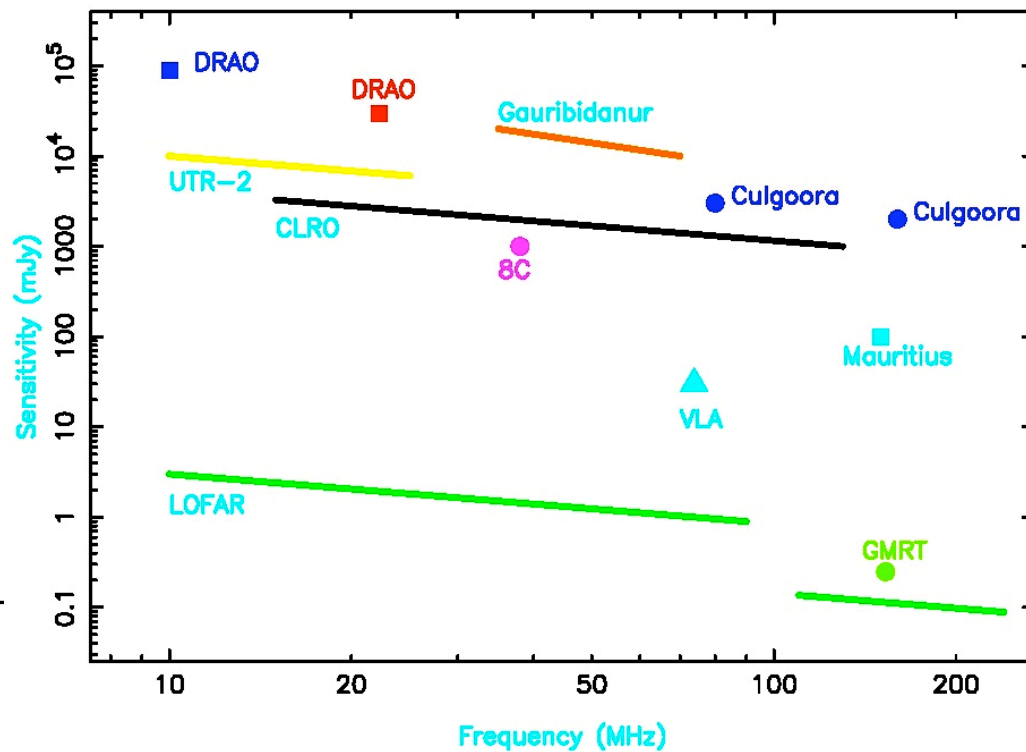
# Covering the Observable Parameter Space

(examples from M. Harwit)



# LOFAR/LWA

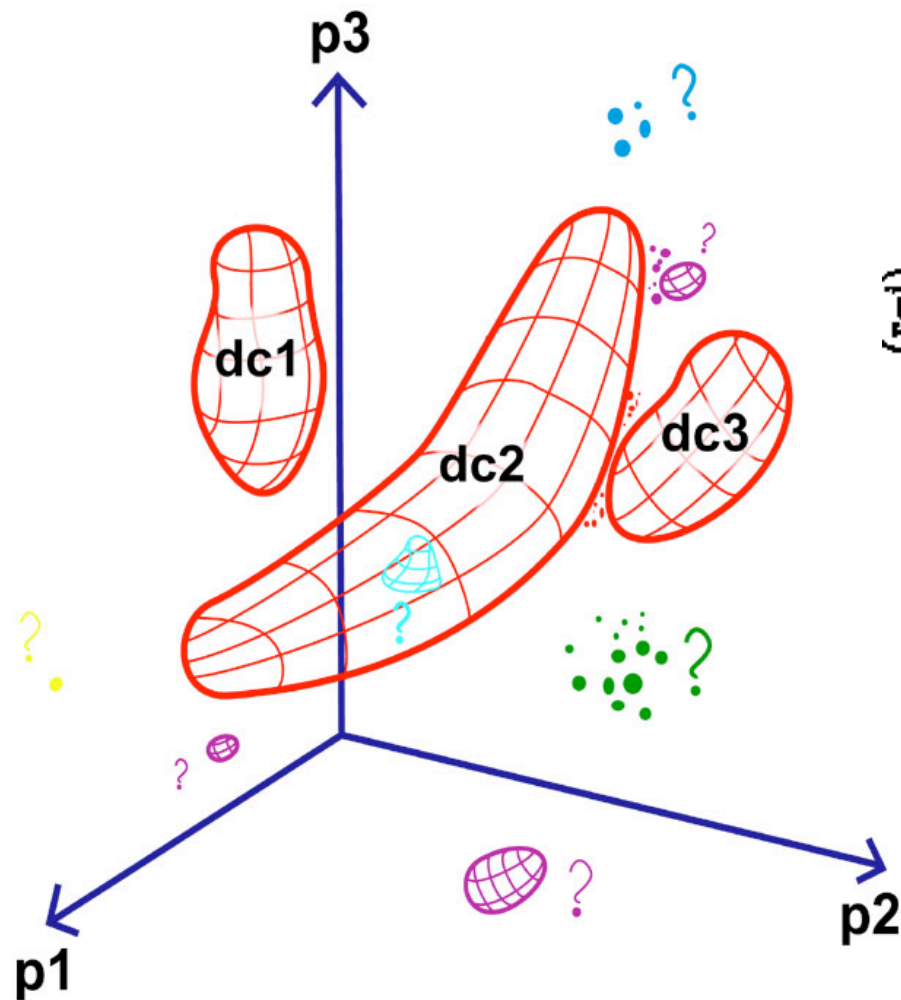
## Gains in the Parameter Space



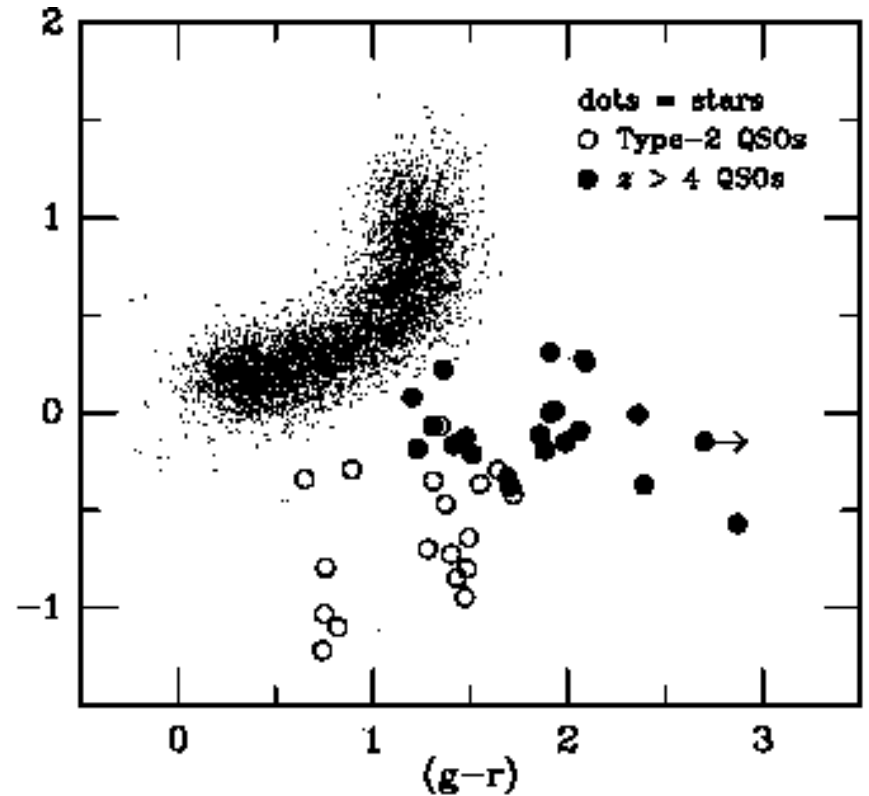
A factor of  $\sim 10^2 - 10^3$  over a broad range in frequency is not so bad ... and then, there is the time domain!

# Exploration of observable parameter spaces and searches for rare or new types of objects

A Generic Machine-Assisted Discovery Problem:  
Data Mapping and a Search for Outliers



A simple, real-life example:



Now consider  $\sim 10^9$  data vectors  
in  $\sim 10^2 - 10^3$  dimensions ...

# The Mixed Blessings of Data Richness

Modern digital sky surveys typically contain  $\sim 10 - 100$  TB, detect  $N_{\text{obj}} \sim 10^8 - 10^9$  sources, with  $D \sim 10^2 - 10^3$  parameters measured for each one -- and PB data sets are on the horizon

Potential for  
discovery

$\left\{ \begin{array}{l} N_{\text{obj}} \text{ or data volume} \rightarrow \text{Big surveys!} \\ N_{\text{surveys}}^2 \text{ (connections)} \rightarrow \text{Data federation} \end{array} \right.$

*Great!* However ...

It takes minutes to hours to search 1 TB (you'd like a few seconds to minutes); 1 PB will take a day to a few months!

We better do it right the first time ...

Or do something more clever (db structuring, statistics?)



## ... And Moreover ...

- **DM algorithms tend to scale very badly:**
  - Clustering  $\sim N \log N \rightarrow N^2, \sim D^2$
  - Correlations  $\sim N \log N \rightarrow N^2, \sim D^k$  ( $k \geq 1$ )
  - Likelihood, Bayesian  $\sim N^m$  ( $m \geq 3$ ),  $\sim D^k$  ( $k \geq 1$ )
- **Visualization fails for  $D > 3 - 5$** 
  - An inherent limitation of the human mind?
- We need better DM algorithms and some novel methods for dimensionality reduction (and some AI help?)
- Or, we learn to accept approximate results
  - Sometimes that is good enough, sometimes not

# Information Technology → New Science

- The information volume grows exponentially

*Most data will never be seen by humans!*

➔ The need for data storage, network, database-related technologies, standards, etc.


- Information complexity is also increasing greatly

*Most data (and data constructs) cannot be comprehended by humans directly!*

➔ The need for data mining, KDD, data understanding technologies, hyperdimensional visualization, AI/Machine-assisted discovery ...

- VO is the framework to effect this for astronomy

# Scientific Roles and Benefits of a VO

- **Facilitate science with massive data sets** (observations and theory/simulations)  **efficiency amplifier**
- Provide an **added value** from federated data sets (e.g., multi-wavelength, multi-scale, multi-epoch ...)
  - Discover the knowledge which is present in the data, but can be uncovered *only* through data fusion
- **Enable and stimulate some qualitatively new science** with massive data sets (not just old-but-bigger)
- **Optimize the use of expensive resources** (e.g., space missions, large ground-based telescopes, computing ...)
- Provide R&D drivers, application testbeds, and stimulus to the **partnering disciplines** (CS/IT, statistics ...)

# VO Developments and Status

- The concept originated in 1990's, developed and refined through several conferences and workshops
- Major blessing by the NAS Decadal Report
- **In the US:** National Virtual Observatory (NVO)
  - Concept developed by the NVO Science Definition Team (SDT). See the report at *<http://www.nvosdt.org>*
  - NSF/ITR funded project: *<http://us-vo.org>*
  - A number of other, smaller projects under way
- **Worldwide** efforts: International V.O. Alliance
- A good synergy of astronomy and CS/IT
- Good progress on data management issues, a little on data mining/analysis, first science demos forthcoming

  
search

## News

- [NVO Summer School](#)
- [Data Inventory Service](#)
- [Discovery by VO Demo](#)
- [VO Alliance Formed](#)
- [NVO News Archive](#)

## About

- [What is the NVO?](#)
- [Who is Involved?](#)
- [Science Objectives](#)

## Community

- [NVO Meetings](#)
- [International VO Alliance](#)

## Documents

Recent NVO Documents:

- [Conesearch definition](#)
- [Quarterly Report Q104](#)
- [Management Plan](#)
- [VO Resource Registry](#)
- [All NVO Documents](#)
- [IVOA Documents](#)



Supported by the  
[National Science Foundation](#)



Member of the  
International

## NVO - Facilitating Scientific Discovery

NVO's objective is to enable new science by greatly enhancing access to data and computing resources. The NVO is developing tools that make it easy to locate, retrieve, and analyze astronomical data from archives and catalogs worldwide, and to compare theoretical models and simulations with observations.

These tools are based upon international standards developed in collaboration with the [International Virtual Observatory Alliance](#).

We expect to deliver the first production quality services in early 2005. Some examples of existing prototypes:

- Use the [VO Spectrum Services](#) to analyze over 500,000 spectra.
- Cross-correlate objects from more than 15 surveys with [SkyQuery](#)
- Use [YourSky](#) to make custom infrared sky images based on DPOSS or 2MASS.

The NVO also provides software libraries and sample code of VO Services for people who want to write their own VO-enabled applications.

## NVO - Data Access

The NVO encourages astronomical research organizations to make their data collections and source catalogs available via the standard VO protocols. These include image access, spectrum access, and catalog search.

A number of [astronomical research facilities and survey projects](#) are already making use of NVO interfaces and protocols in support of data processing, analysis, and distribution.

Available collections and services can be located through the NVO Registries -- the Yellow Pages of astronomical resources, with regularly updated entries. Try the different interfaces at [NCSA](#), [STScI](#), or [Caltech](#) to the NVO registries already containing more than 6000 entries!

## NVO - Education and Public Outreach

Astronomical images are treasured by the public for their beauty, and thus are an excellent vehicle for science education at all levels. We seek partnerships with educational organizations, museums, and planetariums to help them use our tools to incorporate NVO-ready data into their programs and curricula. Sample projects:

- [Project LITE](#) is an interactive environment to study astronomical spectra
- [AnyOne](#), a next-generation web-browser providing encyclopedic access to science information.

This site is a community-maintained collection with content control by the NVO Executive Committee. It is included in the NVO Registry by the extent to which it: (a) reflects an aspect of the Virtual Observatory, such as astronomy data, (b) uses VO standards or software, or (c) exemplifies grid-based astronomical computing. If you would like a description of your project, data, or software included here, please write to web at us-vo.org with a short description of your work.

## Summer School



Aspen CO, Sep 13-17.  
[More Information](#)

## Interop Meeting



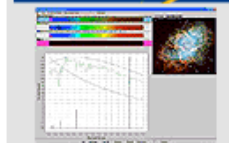
Sep 27-29, 2004, IUCAA  
Pune, India [More Information](#)

## Data Inventory



Find images and catalog objects around a given sky position with the [Data Inventory Service](#).

## Project LITE

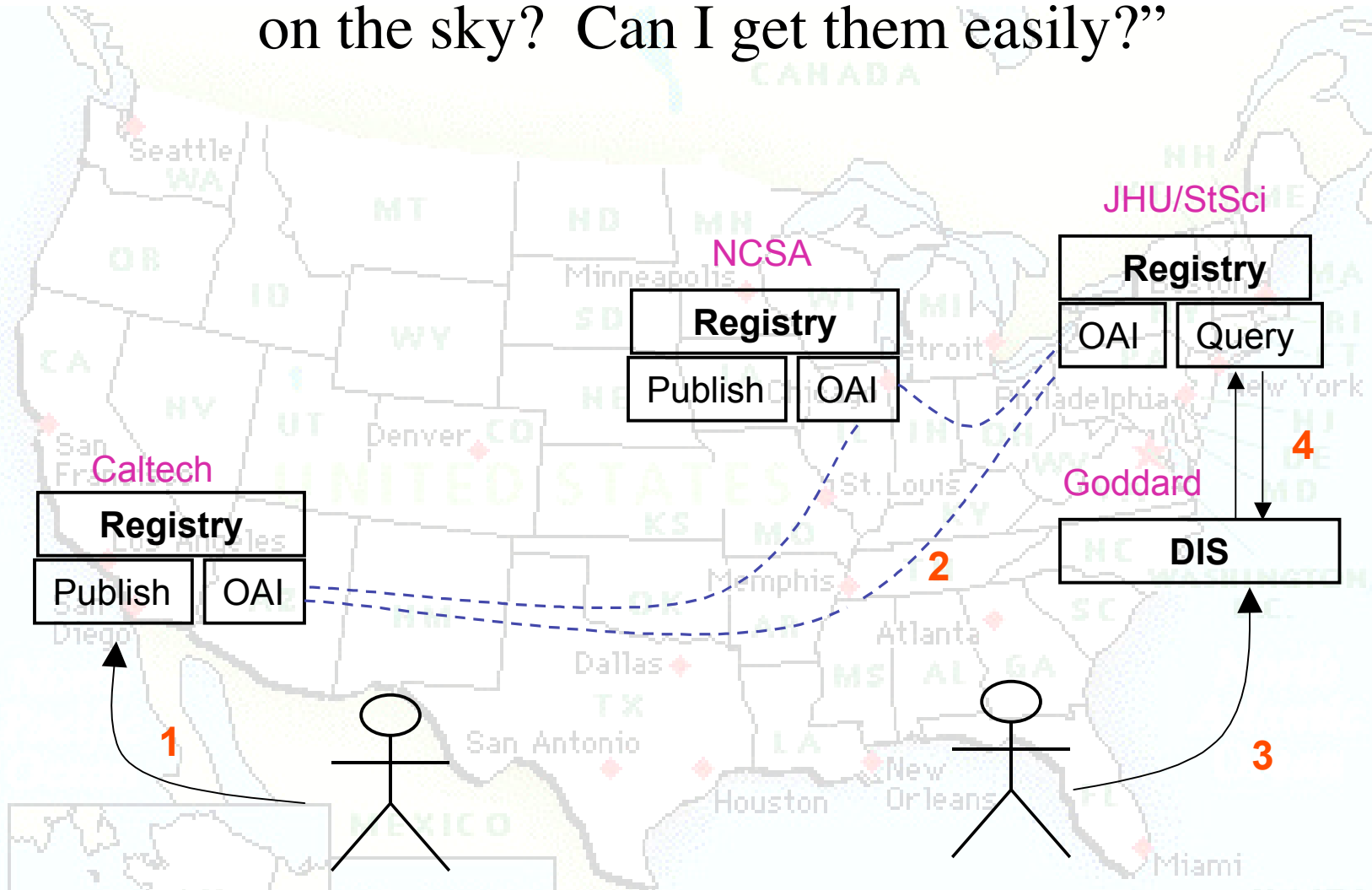


[Project LITE](#) is an instructional environment for astronomical spectra

<http://us-vo.org>

# NVO: A Prototype Data Inventory Service

“What data are available for some object or some region on the sky? Can I get them easily?”





What do we know about regions of sky?

Using new Virtual Observatory protocols we can gather and organize information efficiently on a given region of sky.

Enter a position(or name) and the maximum size of the region of sky you're interested in.

Object Position or Name:  (degrees or sexagesimal)

Size:  (in decimal degrees)

Ignore cache! The DIS will reprocess an identical request rather than linking to the existing cache results.

#### Example Inputs for the Object Position or Name

- 13.29, -18.47 [Object Position: Decimal degrees]
- 6 45 10.8, -16 41 58 [Object Position: Sexagesimal format; RA in hours]
- 3c273 [Object name]
- Use a comma to delimit J2000 RA and Dec pair.

#### About Data Inventory Service

1. A user request is broadcast to sites scattered all over the world using two simple common protocols.
2. Catalog data and lists of available images are returned using the new VOTable XML standard.
3. Image, observation and catalog data from these sites are collected and organized for immediate viewing.
4. Data may be analyzed or visualized in Aladin or OASIS

Participating sites currently include: NRAO, NOAO, JHU, ST ScI, HEASARC, NCS, IRSA, CDS, NED, ESO, SDSS, CXC.



Note: Inventory request completed

RA	Dec	Size
13 25 27.62	-43 01 08.8	0.25

Check All

Images (FITS/GIF)

<b>Optical</b>	<input type="checkbox"/> <a href="#">DSS1 SV</a>	<input checked="" type="checkbox"/> <a href="#">DSS2</a>	<input type="checkbox"/> <a href="#">DSS2B</a>	<input type="checkbox"/> <a href="#">DSS2IR</a>	<input type="checkbox"/> <a href="#">DSS2R</a>
<b>Infrared</b>	<input type="checkbox"/> <a href="#">2MASS-H</a>	<input type="checkbox"/> <a href="#">2MASS-J</a>	<input type="checkbox"/> <a href="#">2MASS-K</a>		
<b>Radio</b>	<input checked="" type="checkbox"/> <a href="#">SUMSS</a>				
<b>X-ray</b>	<input type="checkbox"/> <a href="#">RASS B</a>	<input checked="" type="checkbox"/> <a href="#">Chandra(6)</a>			

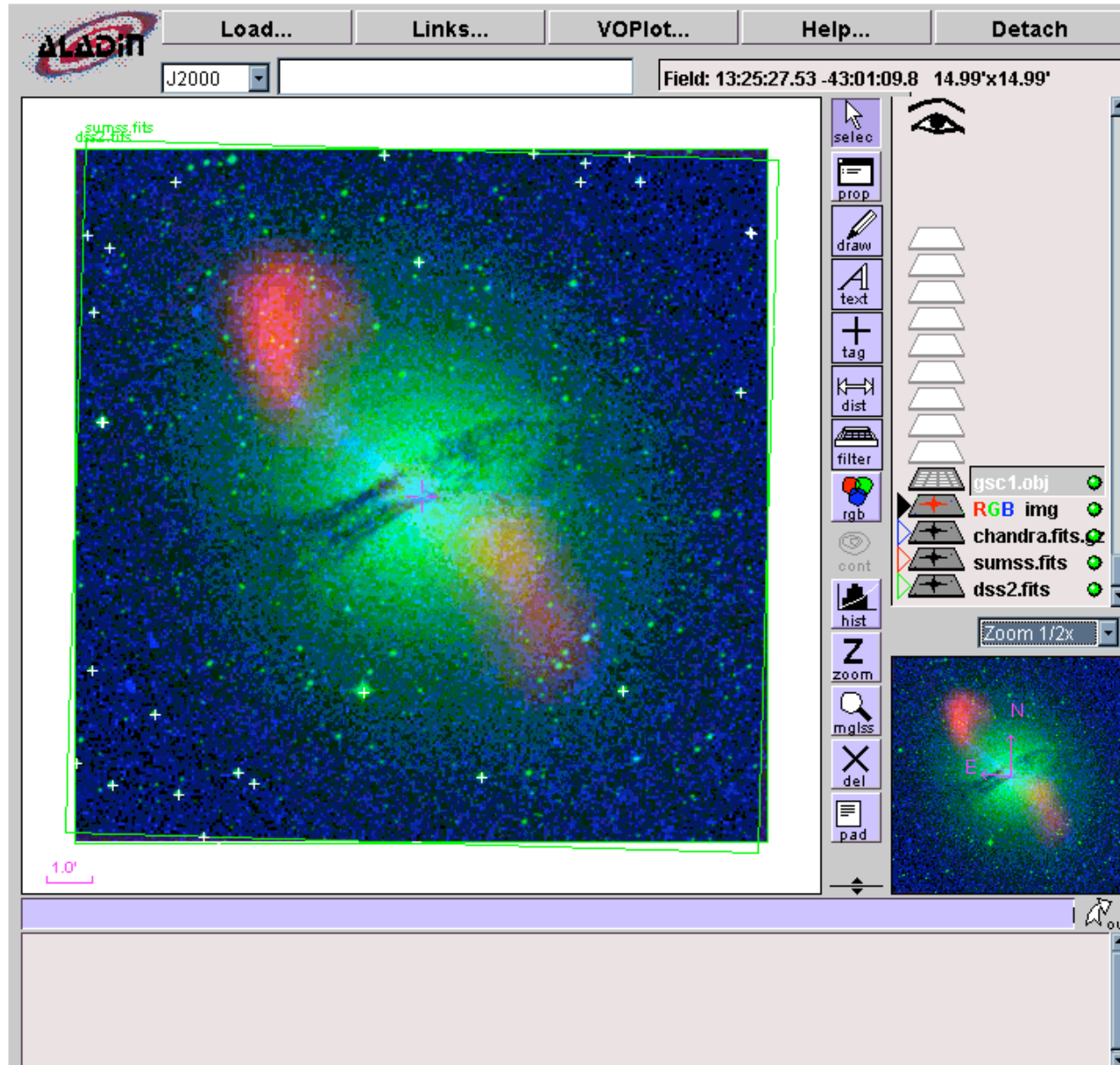
Observations (VOTable)

<b>Optical</b>	<input type="checkbox"/> <a href="#">HST(100)</a>	<input type="checkbox"/> <a href="#">STIS(100)</a>	<input type="checkbox"/> <a href="#">WFPC2(100)</a>	<input type="checkbox"/> <a href="#">WFPC1(22)</a>	<input type="checkbox"/> <a href="#">HSTG(394)</a>
<b>Infrared</b>	<input type="checkbox"/> <a href="#">NICMOS(100)</a>				
<b>X-ray</b>	<input type="checkbox"/> <a href="#">ASCA(3)</a>	<input type="checkbox"/> <a href="#">ROSAT(9)</a>	<input type="checkbox"/> <a href="#">ROSPUBLIC(10)</a>	<input type="checkbox"/> <a href="#">RXTE(23)</a>	<input type="checkbox"/> <a href="#">EXOSAT(12)</a>
	<input type="checkbox"/> <a href="#">CHANMAST(10)</a>	<input type="checkbox"/> <a href="#">Einstein(5)</a>	<input type="checkbox"/> <a href="#">XMMMAST(3)</a>	<input type="checkbox"/> <a href="#">ASCMAST(3)</a>	<input type="checkbox"/> <a href="#">XTEINDEX(5)</a>
<b>Gamma-ray</b>	<input type="checkbox"/> <a href="#">OSSE(29)</a>				
<b>UV</b>	<input type="checkbox"/> <a href="#">FUSE(1)</a>	<input type="checkbox"/> <a href="#">FOC(20)</a>	<input type="checkbox"/> <a href="#">HUT(2)</a>	<input type="checkbox"/> <a href="#">IUE(41)</a>	<input type="checkbox"/> <a href="#">UIT(7)</a>
	<input type="checkbox"/> <a href="#">WUPPE(1)</a>				

Objects (VOTable)

<b>Surveys</b>	<input type="checkbox"/> <a href="#">USNO-A2.0(1197)</a>	<input type="checkbox"/> <a href="#">USNO-SA2.0(1197)</a>	<input checked="" type="checkbox"/> <a href="#">GSC1(289)</a>	<input type="checkbox"/> <a href="#">GSC2.2(2259)</a>	<input type="checkbox"/> <a href="#">UCAC1(305)</a>
	<input type="checkbox"/> <a href="#">USNO-A2.0 CDS(999)</a>				
<b>Galaxies</b>	<input type="checkbox"/> <a href="#">SGC(1)</a>	<input type="checkbox"/> <a href="#">PGC(1)</a>	<input type="checkbox"/> <a href="#">NBG(1)</a>	<input type="checkbox"/> <a href="#">RC3(1)</a>	<input type="checkbox"/> <a href="#">RNGC(1)</a>
	<input type="checkbox"/> <a href="#">PSCz(3)</a>				
<b>Stars</b>	<input type="checkbox"/> <a href="#">HIP(1)</a>	<input type="checkbox"/> <a href="#">SAO(2)</a>	<input type="checkbox"/> <a href="#">WDS(1)</a>	<input type="checkbox"/> <a href="#">AC2000.2(30)</a>	<input type="checkbox"/> <a href="#">ASCC-2.5(21)</a>
	<input type="checkbox"/> <a href="#">HD(4)</a>				
<b>Misc.</b>	<input type="checkbox"/> <a href="#">EGRET3(45)</a>	<input type="checkbox"/> <a href="#">WGACAT(35)</a>	<input type="checkbox"/> <a href="#">Radio Catalogs(69)</a>	<input type="checkbox"/> <a href="#">2MASS-PSC(CDS)(999)</a>	<input type="checkbox"/> <a href="#">Veron-Veron(1)</a>
	<input type="checkbox"/> <a href="#">TYCHO-2(22)</a>				





The screenshot displays the Aladin sky atlas interface. At the top, there are menu buttons: "Load...", "Links...", "VOPlot...", "Help...", and "Detach". Below these is a search bar with "J2000" selected and a field coordinate "Field: 13:25:27.53 -43:01:09.8 14.99'x14.99'". The main window shows a large astronomical image with a green border, labeled "sumss.fits" and "dss2.fits" in the top left. The image is a multi-color representation of a galaxy field. To the right of the main image is a vertical toolbar with icons for "selec", "prop", "draw", "text", "tag", "dist", "filter", "rgb", "cont", "hist", "zoom", "mgls", "del", and "pad". Below the toolbar is a list of loaded data layers: "gsc1.obj", "RGB img", "chandra.fits", "sumss.fits", and "dss2.fits", each with a green checkmark. A "Zoom 1/2x" dropdown is also present. At the bottom right, there is a small thumbnail of the image with a coordinate system (N, E) and a "1.0''" scale bar. The interface is styled with a light blue and white color scheme.

# SkyQuery: NVO Prototype Catalog Cross-Matching Service


The screenshot displays the SkyQuery.net web interface. On the left, a vertical list of surveys is shown with radio buttons, including ROSAT, GALEX, INTWFS, RC3, DLS, USNOB, TWODF, TWOQZ, SDSS, HSTEXP, HDFN, HDF5, GOODS, UDF, TWOMASS, PSCZ, IRAS, NVSS, FIRST, and AGC. Below this list, it says 'Surveys ordered by wavelength. Click on a survey to see WS test.' The main content area features the 'SkyQuery.net' logo, a 'Table Info' section with a 'Tables' dropdown menu and an 'Argument(if required):' input field, and a 'Search' button. Below the search form is a 'SkyQL query' section containing a SQL query: 

```
SELECT o.objId, o.ra, o.r, o.type, t.objId
FROM SDSS:PhotoPrimary o,
      TWOMASS:PhotoPrimary t
WHERE XMATCH(o,t)<3.5
      AND AREA(181.3,-0.76,6.5)
      AND o.type=3
```

 At the bottom of the interface, there is a 'Submit' button, a set of four numbered buttons (1, 2, 3, 4), and a link to 'Sample queries from tutorial'. On the right side of the interface, there is a large image of a blue, irregularly shaped galaxy or nebula against a dark background with other stars.

... and  
much  
more is  
coming!

# Broader and Societal Benefits of a VO

- **Professional Empowerment:** Scientists and students anywhere with an internet connection would be able to do a first-rate science  A broadening of the talent pool in astronomy, democratization of the field
- **Interdisciplinary Exchanges:**
  - The challenges facing the VO are common to most sciences and other fields of the modern human endeavor
  - Intellectual cross-fertilization, feedback to IT/CS
- **Education and Public Outreach:**
  - Unprecedented opportunities in terms of the content, broad geographical and societal range, at all levels
  - Astronomy as a magnet for the CS/IT education

*“Weapons of Mass Instruction”*

<http://virtualsky.org> (R. Williams et al.)

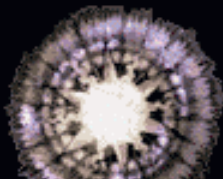
[Increase map size](#)  
 [Decrease map size](#)  
 [Show Marker](#)  
 **Clicking on sky will:**  
 Zoom In  
 Zoom Out  
 Recenter  
 Galaxy lookup  
 Star lookup  
 **You are at:**  
**Right Ascension**  
  
**Declination**



30 arcmin

Image courtesy of [Digital Palomar Observatory Sky Survey](#)

Image center is RA,Dec=14.679072 60.933364  
Image width is 53.715627 minutes  
[Get VS jpeg](#) [Get VS FITS](#)




# International Virtual Observatory Alliance

## Member Organizations

[http:// ivoa.net](http://ivoa.net)



# Do We Know How to Run a VO?

- The VO is *not* yet another data center, archive, mission, or a traditional project  *It does not fit into any of the usual structures today*
  - It is inherently *distributed*, and web-centric
  - It is fundamentally based on a *rapidly developing technology* (IT/CS)
  - *It transcends the traditional boundaries* between different wavelength regimes, agency domains
  - It has an *unusually broad range of constituents* and interfaces
  - It is inherently *multidisciplinary*
- The VO represents *a novel type of a scientific organization* for the era of information abundance

**The rate of the overall computing power has been amazingly growing for more than one hundred years**

### Computing efficiency in ops/s/\$ had 3 growth curves:

Combination of Hans Moravac + Larry Roberts + Gordon Bell  
WordSize\*ops/s/sysprice

**1890-1945**

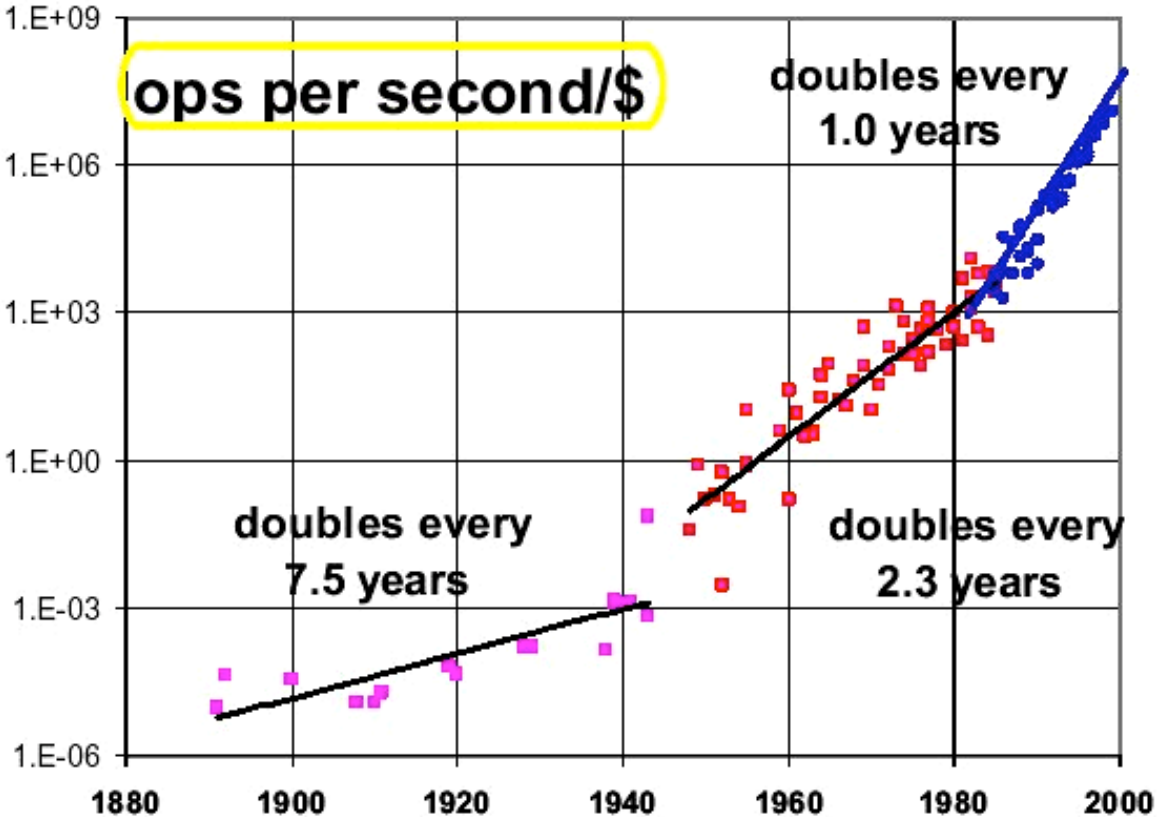
Mechanical  
Relay  
7-year doubling

**1945-1985**

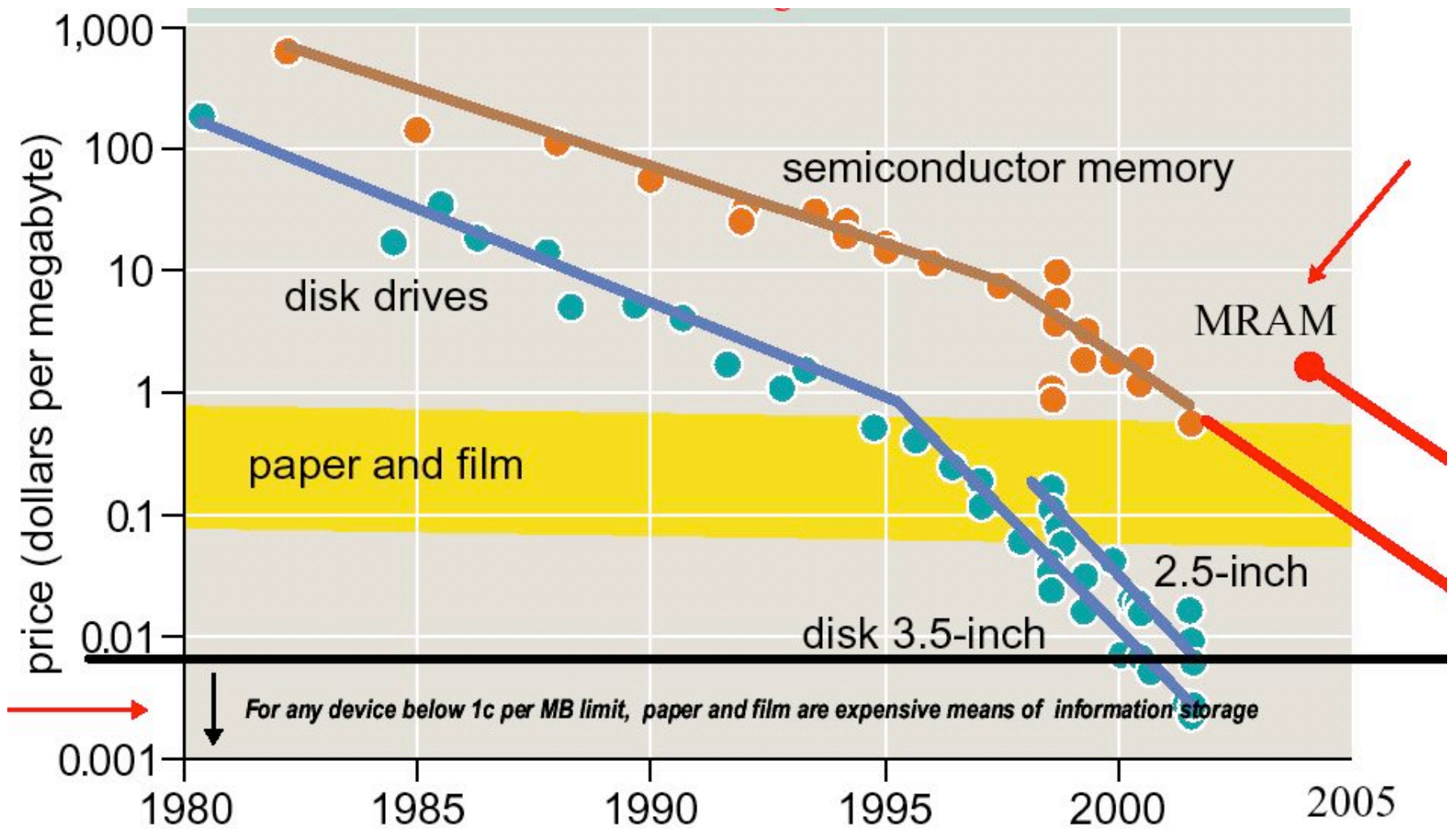
Tube, transistor,...  
2.3 year doubling

**1985-2000**

Microprocessor  
1.0 year doubling



# Exponentially Declining Cost of Data Storage



Hayes, Grochowski: American Scientist, 2002



# Computing is Cheap ...

Today (~2004), 1 \$ buys:

- 1 day of CPU time
- 4 GB (fast) RAM for a day
- 1 GB of network bandwidth
- 1 GB of disk storage for 3 years
- 10 M database accesses
- 10 TB of disk access (sequential)
- 10 TB of LAN bandwidth (bulk)
- 10 KWh = 4 days of computer time

... Yet somehow computer companies make billions: you do want some toys, about \$  $10^5$  worth  $\approx$  1 postdoc year

**... But People are Expensive!**

People ~ Software, maintenance, expertise, creativity ...

# Moving Data is Slow!

How long does it take to move a Terabyte?

(how about a Petabyte?)

Context	Speed Mbps	Rent \$/month	\$/Mbps	\$/TB Sent	Time/TB
Home phone	0.04	40	1,000	3,086	6 years
Home DSL	0.6	50	117	360	5 months
T1	1.5	1,200	800	2,469	2 months
T3	43	28,000	651	2,010	2 days
OC3	155	49,000	316	976	14 hours
OC 192	9600	1,920,000	200	617	14 minutes
100 Mbps	100				1 day
Gbps	1000				2.2 hours

Source: TeraScale Sneakernet, Microsoft Research, Jim Gray et al.

# Disks are Cheap and Efficient

- Price/performance of disks is improving faster than the computing (Moore's law): a factor of  $\sim 100$  over 10 years!
  - Disks are now already cheaper than paper
- Network bandwidth used to grow even faster, but no longer does
  - And most telcos are bankrupt ...
  - Sneakernet is faster than any network
- Disks make data preservation easier as the storage technology evolves
  - Can you still read your 10 (5?) year old tapes?

# An Early Disk for Information Storage

- Phaistos Disk:  
Minoan, 1700 BC



- No one can read it 😊

(From Jim Gray)

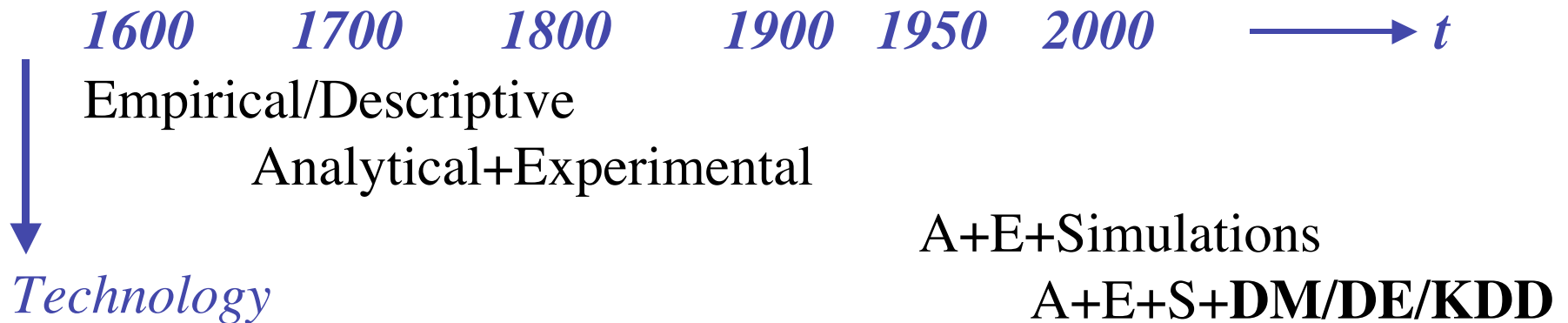
# The Gospel According to Jim Gray:

- Store everything on disks, with a high redundancy (cheaper than the maintenance/recovery)
  - Curate data where the expertise is
- Do not move data over the network: **bring the computation to data!**
  - The Beowulf paradigm: Datawulf clusters, smart disks ...
  - The Grid paradigm (done right): move only the questions and answers, and the flow control
- You *will* learn to use databases!
- And we need a better fusion of databases and data mining and exploration

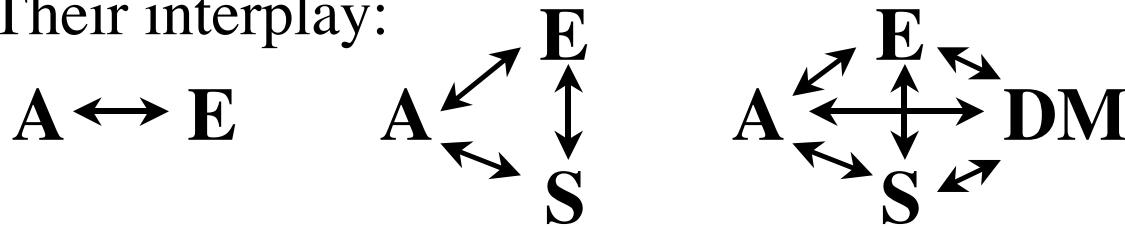
# We are Not Alone in Our Predicament

- Astronomical data volume *ca.* 2004: **a few  $\square 10^2$  TB**  
(but PB's are coming soon!)
- All recorded information in the world: **a few  $\square 10^7$  TB**  
(but most of it is video, *i.e.*, junk)
- The data volume everywhere is growing exponentially, with *e*-folding times  $\sim 1.5$  yrs (Moore's law)
  - NB: the data rate is also growing exponentially!
- So, ***everybody*** needs efficient db techniques, DM (searches, trends & correlations, anomaly/outlier detection, clustering/classification, summarization, visualization, etc.)
- What others discover will help us, and maybe we can also help change the world!

# The Evolution of Science



Their interplay:



Computational science rises with the advent of computers

Data-intensive science is a more recent phenomenon

## The Evolving Role of Computing:

Number crunching  $\rightarrow$  Data intensive (data farming, data mining)

# Some Musings on CyberScience

- Enables a broad spectrum of users/contributors
  - From large teams to small teams to individuals
  - Data volume  $\sim$  Team size
  - Scientific returns  $\neq f(\text{team size})$
  - Human talent is distributed very broadly geographically
- Transition from data-poor to data-rich science
  - Chaotic  $\rightarrow$  Organized ... However, *some* chaos (or the lack of excessive regulation) is good, as it correlates with the creative freedom (recall the WWW)
- Computer science as the “new mathematics”
  - It plays the role in relation to other sciences which mathematics did in  $\sim 17^{\text{th}}$  -  $20^{\text{th}}$  century  
(The frontiers of mathematics are now elsewhere...)



# Summary

- National/International Virtual Observatory is an *emerging framework* to harness the power of IT for astronomy with massive and complex data sets
  - Facilitate inclusion of major new data providers, surveys
  - Enable data archiving, fusion, exploration, discovery
  - Cross the traditional boundaries ( $\square$ , ground/space, th/obs)
  - An excellent synergy with the applied CS/IT, statistics...
  - Broad professional empowerment via the WWW
  - Great for E/PO at all educational levels
- Extracting knowledge from massive/complex data sets is *a universal problem facing all sciences today*:  
Quantitative changes in data volumes + IT advances:  
→ *Qualitative changes in the way we do science*