

Hyperfine structure:

Fine-structure is due to interaction of the magnetic field due to orbit of proton with the magnetic momentum of the electron.

We now consider the interaction of the magnetic field of the proton (which is magnetized) with the magnetic moment of the electron.

The proton itself is magnetized

$$\vec{\mu}_p = \gamma_p \frac{e\hbar}{2m_p} \frac{\vec{I}}{\hbar} \quad I: \text{spin of proton}$$

$$\gamma_p = 2.675 \dots \mu_N = e\hbar/2m_p \quad (\text{nuclear magneton})$$

Incidentally $\vec{\mu}_n \equiv \gamma_n \frac{e\hbar}{2m_p} \frac{\vec{I}}{\hbar}$, $\gamma_n = -1.832 \dots$

This magnetic moment creates a B-field

$$\vec{B} = \frac{3\vec{r}(\vec{r} \cdot \vec{\mu}) - r^2 \vec{\mu}}{r^5} + \frac{8\pi}{3} \mu \delta^3(\vec{r})$$

The resulting "hyper" fine structure is $\frac{m_e}{m_p}$ smaller than fine structure.

We focus on $l=0$.

$$H_{fs} = \frac{8\pi}{3} \frac{\mu_B e \hbar^{-1} \cdot \mu_N e \hbar^{-1}}{2m_p \cdot 2m_e} \mathbf{I} \cdot \mathbf{S}$$

$$\langle H_{fs} \rangle = H_{fs} |\Psi_{1s, \frac{1}{2}}(0)|^2$$

As before, let $\mathbf{F} = \mathbf{I} + \mathbf{S}$

$$\begin{aligned} \frac{\mathbf{I} \cdot \mathbf{S}}{\hbar^2} &= \frac{1}{2} (\overline{\mathbf{F} \cdot \mathbf{F}} - \overline{\mathbf{S} \cdot \mathbf{S}} - \overline{\mathbf{I} \cdot \mathbf{I}}) \\ &= \frac{1}{2} (F(F+1) - S(S+1) - I(I+1)) \end{aligned}$$

$$F=0, F=1$$

1 Rydberg is $\frac{1}{2} m_e c^2 (Z\alpha)^2$

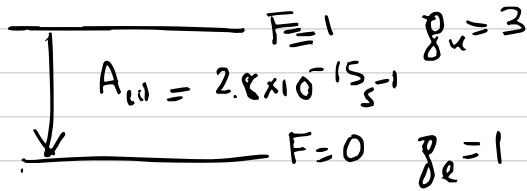
Bohr levels: $\frac{Ry}{n^2}$

fine structure: $Ry \left(\frac{Z\alpha}{n}\right)^2$

hyperfine structure: $Ry \frac{(Z\alpha)^2}{n} \times \frac{m_e}{m_N} \text{ nucleus}$

21-cm line: $\nu = 1420.4 \text{ MHz}$

$$\frac{h\nu}{k_B} = 68 \text{ mK}$$



$$\frac{n_u/g_u}{n_l/g_l} = \exp\left(-\frac{h\nu_{ul}}{kT}\right)$$

for any reasonable T $\exp\left(-\frac{h\nu}{kT}\right) = 1 - \frac{h\nu}{kT} \approx 1$

$$n_u \approx \frac{3}{4} n_{\text{HI}}, \quad n_l \approx \frac{1}{4} n_{\text{HI}}$$

The emissivity is

$$j_\nu = \frac{n_u A_{ul}}{4\pi} h\nu_{ul} \phi(\nu) = \frac{3}{4} \frac{n_{\text{HI}} A_{ul} h\nu_{ul} \phi_\nu}{4\pi}$$

The absorption coefficient is

$$\begin{aligned} k_\nu &= n_l \sigma_{lu} - n_u \sigma_{ul} \\ &= n_l \frac{g_u}{g_l} \frac{A_{ul}}{8\pi} \lambda_{ul}^2 \phi_\nu \left[1 - \frac{n_u/g_u}{n_l/g_l} \right] \\ &\approx \frac{3}{32\pi} A_{ul} \frac{hc \lambda_{ul}}{kT} n_{\text{HI}} \phi_\nu \end{aligned}$$

$T \rightarrow$ "spin" temperature (Salpeter)

Thus $K_v \propto \frac{1}{T_{\text{spin}}}$

$$\tau_v = 2.19 \left(\frac{N_{\text{HI}}}{10^{21} \text{ cm}^{-2}} \right) \left(\frac{100 \text{ K}}{T_{\text{spin}}} \right) \left(\frac{\text{km s}^{-1}}{\sigma_v} \right) e^{-\frac{u^2}{2\sigma_v^2}}$$