

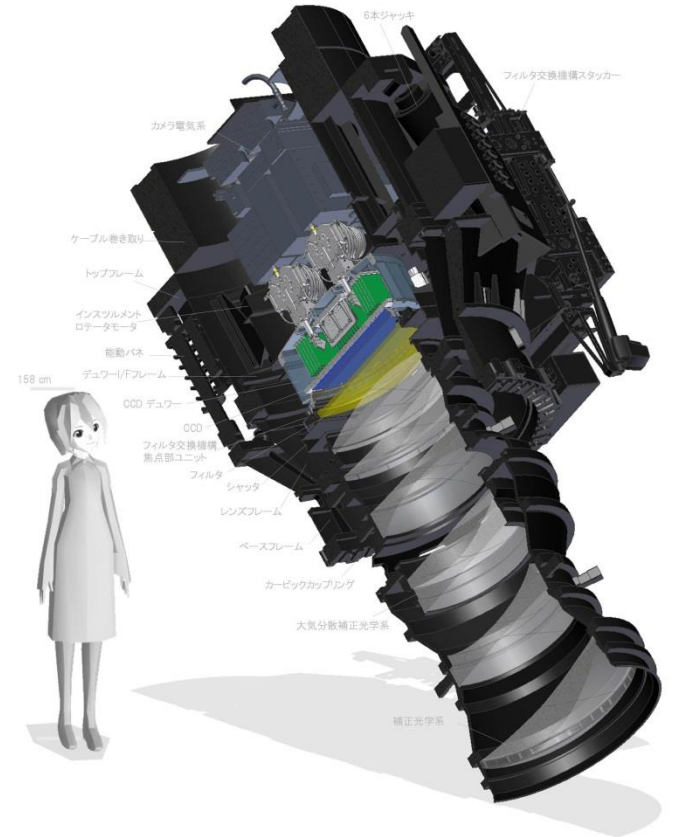
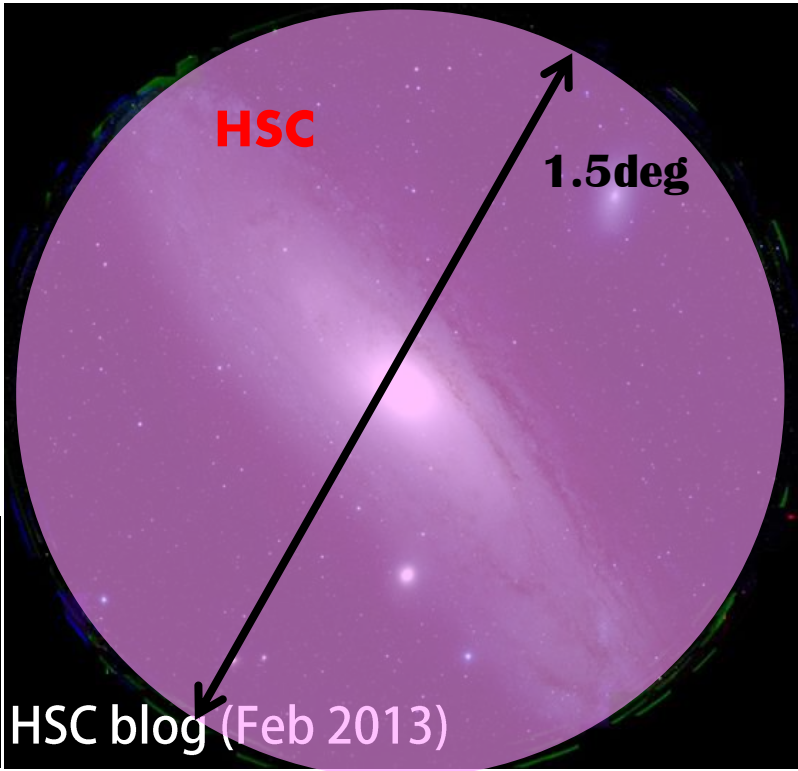
Subaru transient survey

Nozomu Tominaga
(Konan University)



Subaru/Hyper Suprime Cam

- Hyper Suprime-Cam (HSC)
 - Diameter: 8.2m, FoV: 1.77deg², ~900M pixels



HSC instrument parameters

- Number of science CCDs:
104 chips (4 unavailable chips)
- Overhead time: ~35 sec per exposure
- Number of filters: 6
- Filter exchange: ~30 min

5 σ lim. mag.
w/ 1min exp.

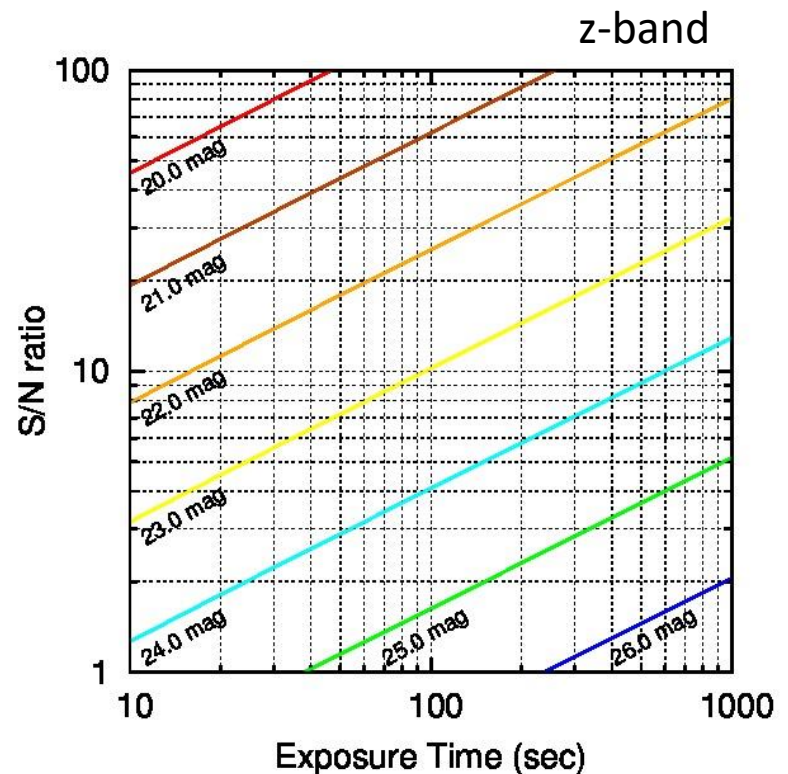
g 25.5

r 23.9

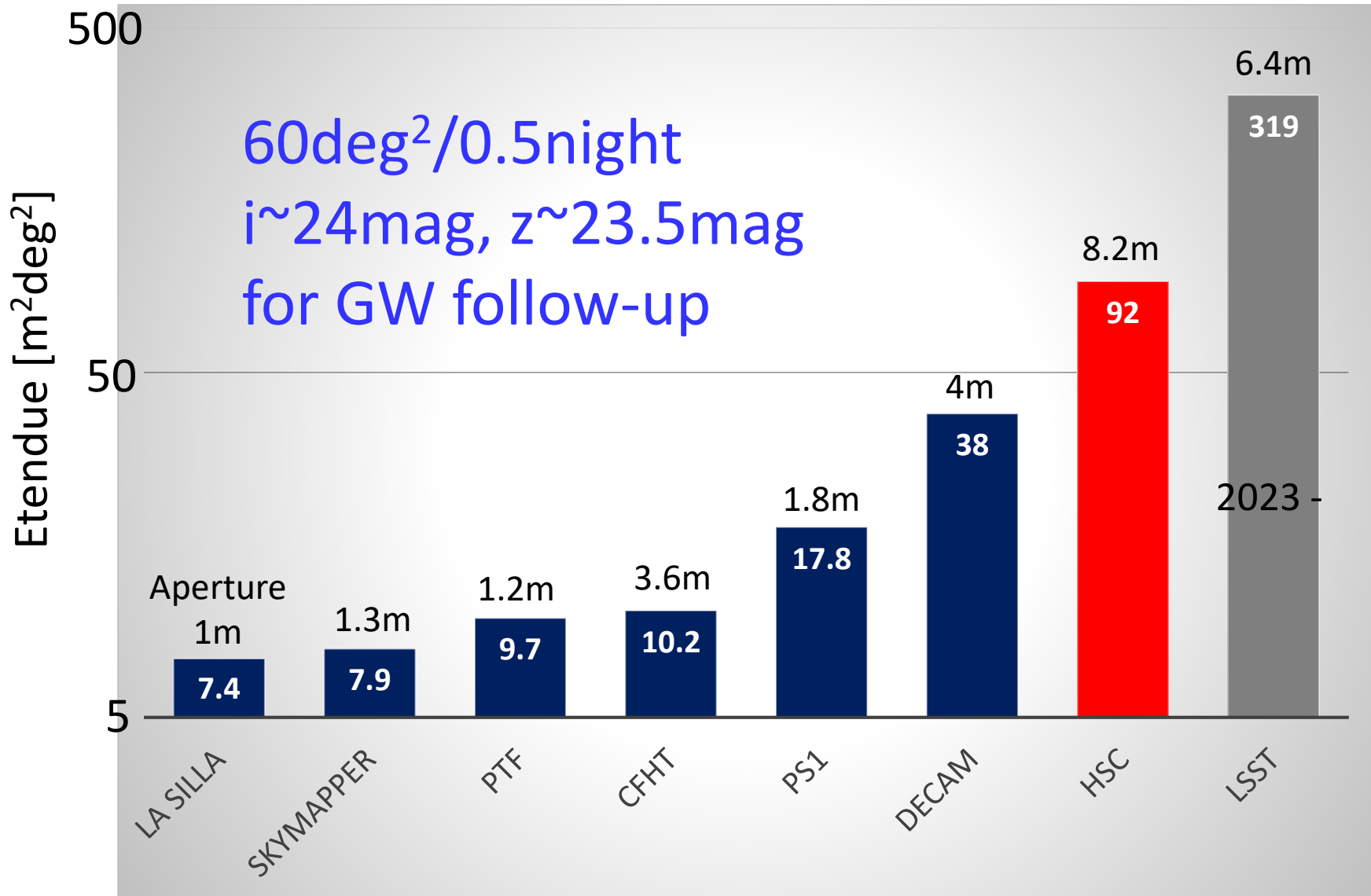
i 24.2

z 23.6

Y 22.8



Etendue of optical telescopes



Available only on gray/dark nights

Schedule for January 2018

Sun	Mon	Tue	Wed	Thu	Fri	Sat
	Jan 01 ○	Jan 02	Jan 03	Jan 04	Jan 05	Jan 06
	Obs FOCAS	S17B-002 Kodama MOIRCS		S17B-002 Kodama MOIRCS	S17B-130 Kotani CHARIS+SCExAO	UH-07B Hodapp CHARIS+SCExAO
	Obs FOCAS				CHARIS+SCExAO	S17B-130 Kotani CHARIS+SCExAO
Jan 07	Jan 08 ◐	Jan 09	Jan 10	Jan 11	Jan 12	Jan 13
UH-07B Hodapp CHARIS+SCExAO	SSP HSC	SSP HSC	Queue HSC	Queue HSC	Queue HSC	SSP HSC
Taiken Kikaku (1hr)/Obs HDS	Eng/Queue HSC	S17B-055I Suzuki HSC			S17B-055I Suzuki HSC	
Jan 14	Jan 15	Jan 16 ●	Jan 17	Jan 18	Jan 19	Jan 20
UH-18B Tholen HSC	UH-18B Tholen HSC	S17B-116 [ToO] Y. Tanaka HSC	S16B-001I Inoue HSC	S16B-001I Inoue HSC	S16B-001I Inoue HSC	SSP HSC
		SSP HSC				
Jan 21	Jan 22	Jan 23 ◐	Jan 24	Jan 25	Jan 26	Jan 27
S17B-044 Yoshida HSC	Keck Prochaska HSC	Eng/Queue HSC	UH-28A Goebel CHARIS+SCExAO	S17B-093 Currie CHARIS+SCExAO	S16A-119I Aoki HDS	S16A-119I Aoki HDS
Queue HSC	S17B-055I Suzuki HSC	SSP HSC	S16A-119I Aoki HDS	S16A-119I Aoki HDS		Obs IRCS+AO188(LGS)
Jan 28	Jan 29	Jan 30 ○	Jan 31			
Keck Melis COMICS	S17B-092 Takagi IRCS+AO188(LGS)					

Need to submit proposals

- Deadline: early Sep/Mar for A/B semesters

Semester		Proposals			Nights		
		Submitted	Accepted	Ratio	Requested	Awarded	Ratio
S17A	2017/02 - 2017/07	166	42	25%	418.3	82	20%
S17B	2017/08 - 2018/01	135	37	27%	294	69.5	24%
S18A	2018/02 - 2018/07	155	45	29%	347.3	94	27%
S18B	2018/08 - 2019/01	156	50	32%	415.7	84.5	20%
S19A	2019/02 - 2019/07	133	46	35%	354.9	89.5	25%

- HSC is more competitive than the average.

There are 3 kinds of proposals.

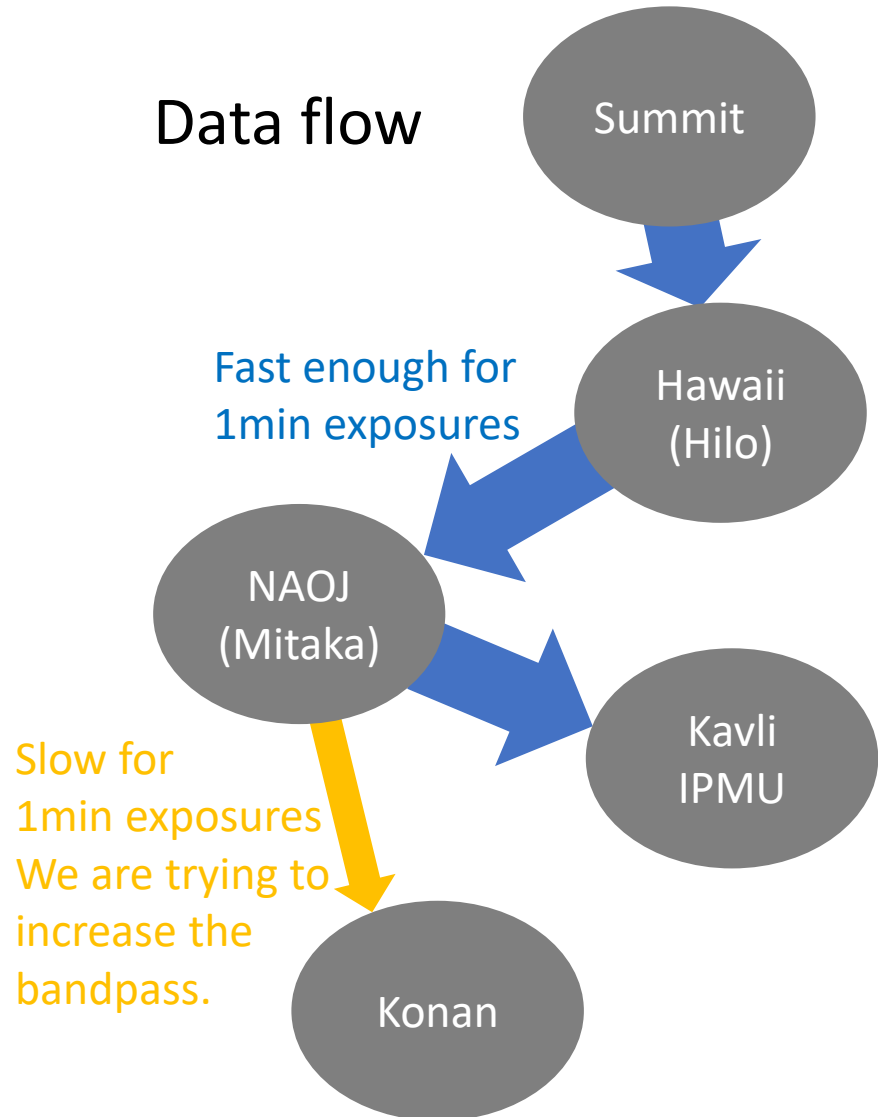
- Openuse: ≤ 5 nights, 1 semester
- Intensive: ≤ 40 nights, ≤ 6 semesters
- SSP: ≤ 300 nights, ≤ 10 semesters

Data analysis for transient surveys

Transient finding system

- Naoki Yasuda and NT are working on it.
- Hawaii observatory
 - CPU: 176 cores
 - Storage: 20TB
- Kavli IPMU
 - CPU: 1200 cores
 - Storage: 3.5PB
- Konan University
 - CPU: 488 (+320) cores
 - Storage: 500TB

Data flow

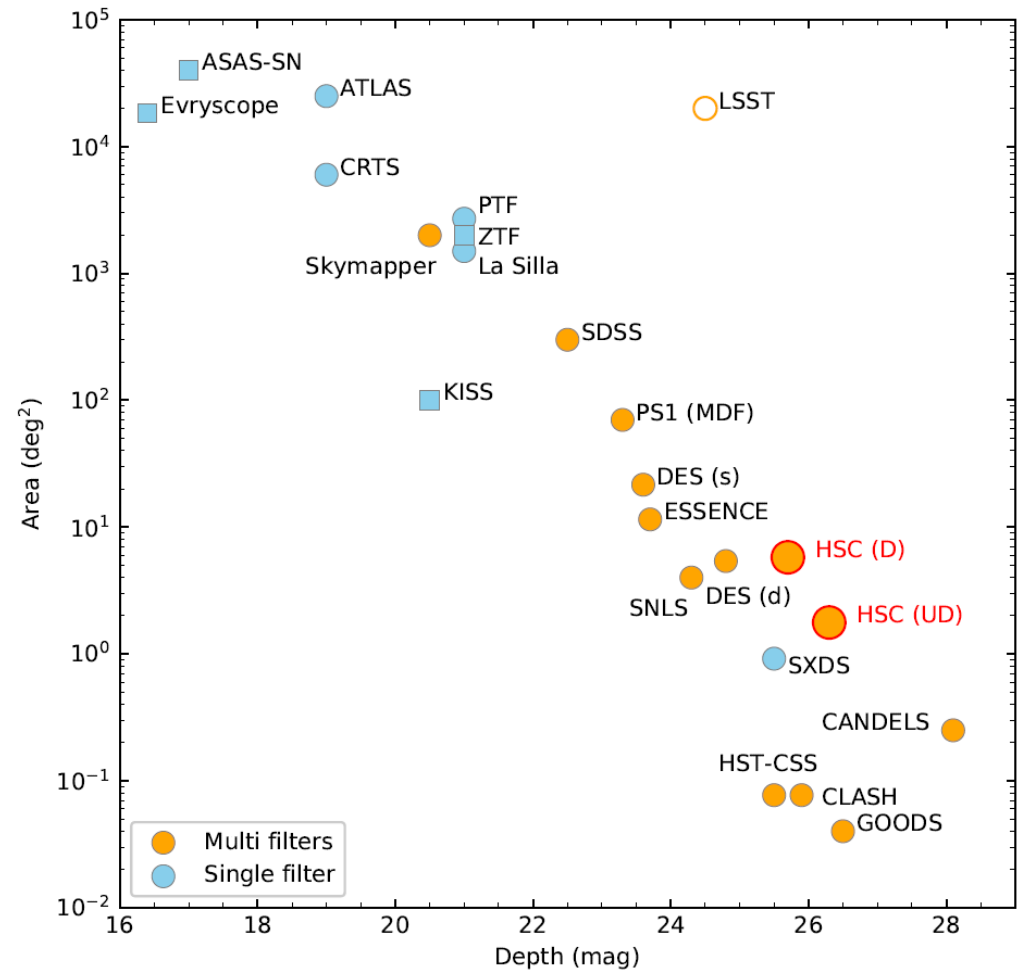
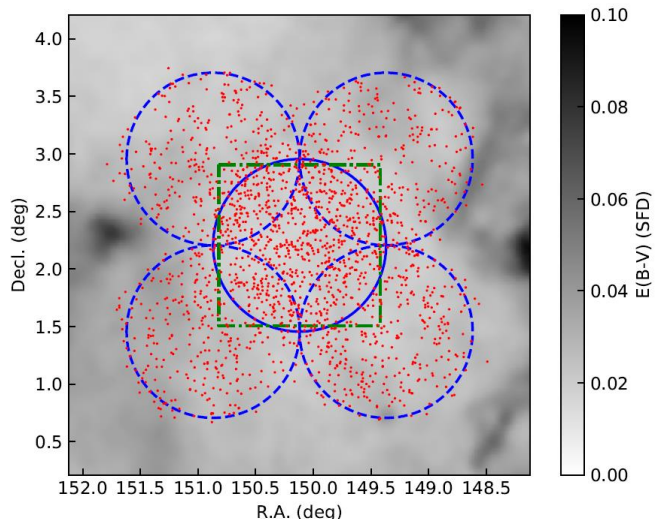


Subaru strategic program (SSP)

COSMOS (Nov 2016 – Apr 2017): several days cadence

Two layers:

- Deep: $\Omega=5.8\text{deg}^2$
 $m_{\text{lim}} \sim 25.7$, 4 months
- Ultradeep: $\Omega=1.8\text{deg}^2$
 $m_{\text{lim}} \sim 26.3$, 6 months

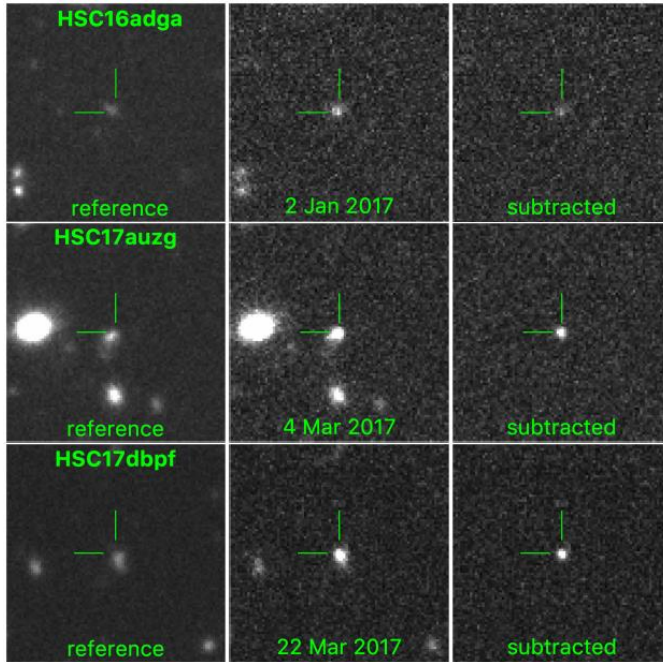


Yasuda + submitted

Pick-up results -SLSNe-

Subaru High-Z sUpernova CAmpaign (SHIZUCA)

Moriya+18 (arXiv:1801.08240)
Curtin+18 (arXiv:1801.08241)



SHIZUCA discovered SLSNe at $z=1.9-2.4$.

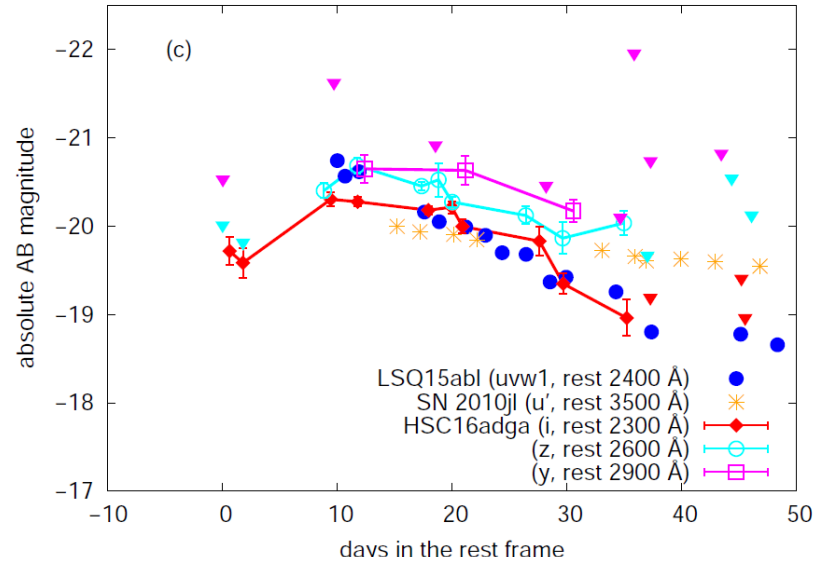


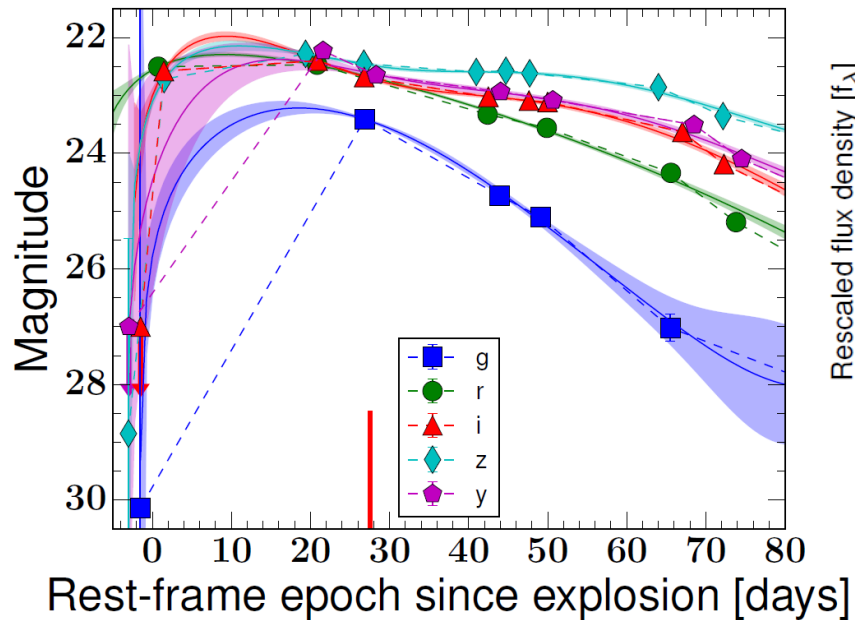
Table 1. List of SNe and SN candidates.

HSC name	IAU name	redshift	host galaxy magnitudes in the HSC filters					Section
			g	r	i	z	y	
HSC16adga	SN 2016jhm	2.399 ± 0.004^a	24.55 ± 0.03	24.42 ± 0.04	24.48 ± 0.06	24.29 ± 0.07	24.20 ± 0.13	3.1
HSC17auzg	SN 2016jhn	1.965 ± 0.004^a	23.88 ± 0.02	23.77 ± 0.02	23.54 ± 0.02	23.41 ± 0.03	23.58 ± 0.06	3.2
HSC17dbpf	SN 2017fei	1.851 ± 0.004^a	24.11 ± 0.02	23.91 ± 0.02	23.67 ± 0.03	23.63 ± 0.04	23.60 ± 0.08	3.3
HSC16apuo	AT 2016jho	$2.8225^{+0.4727}_-0.7032^b$	27.00 ± 0.75	25.31 ± 0.19	25.50 ± 0.35	24.92 ± 0.29	26.10 ± 0.29	4.1
HSC17dsid	AT 2017fej	$4.1974^{+0.0908}_-0.126^b$	27.74 ± 0.34	25.07 ± 0.04	24.83 ± 0.04	24.68 ± 0.05	25.23 ± 0.18	4.2

^aSpectroscopically confirmed (Curtin et al. 2018).

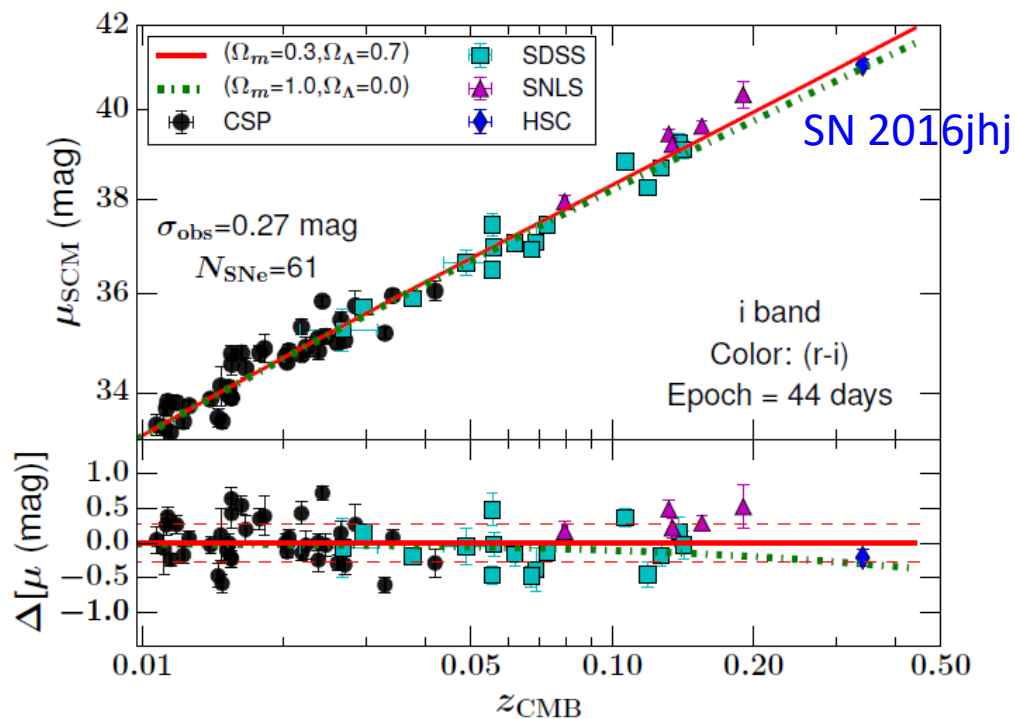
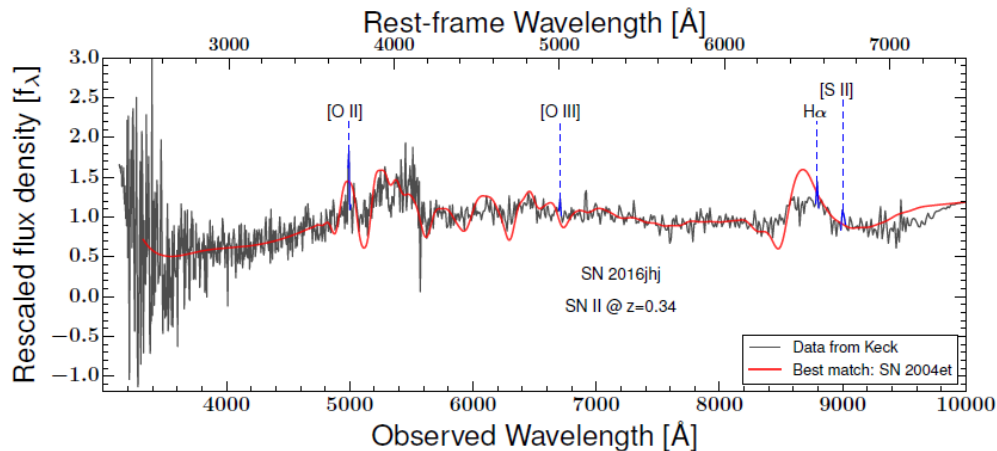
^bCOSMOS2015 photometric redshift (Laigle et al. 2016).

Pick-up results -IIP cosmology-



SN 2016jhj at $z=0.34$ extended the Type II supernova Hubble diagram.

de Jaeger+18 (arXiv:1709.01513)

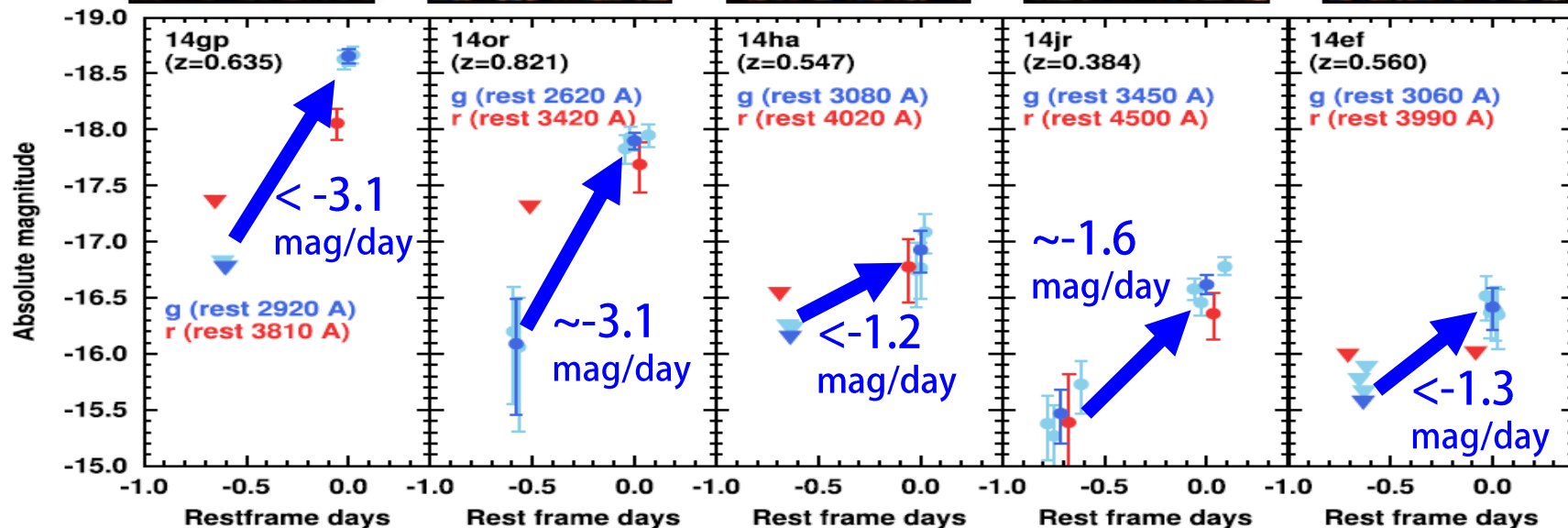
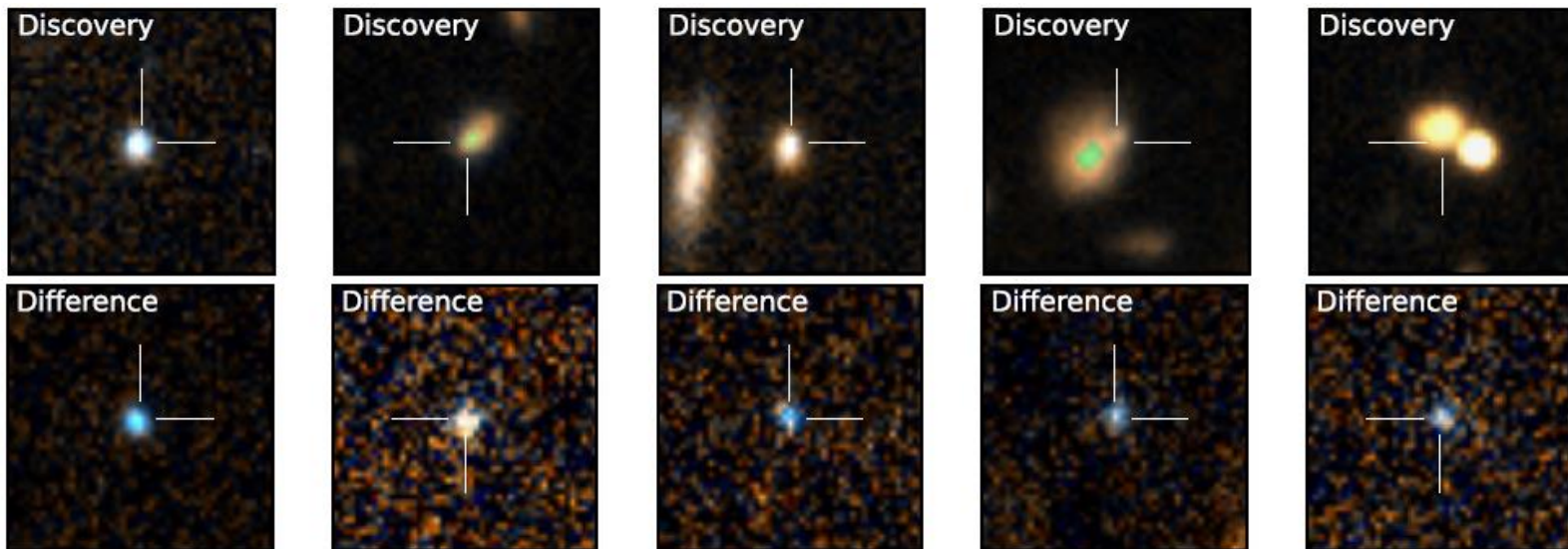


Openuse program

- Subaru HSC survey optimized for optical transients
 - PI: Nozomu Tominaga
 - Jul 2014 (2 nights): g,r 10min exp., 7 fields, ~1hr cadence
follow-up at ~1month later
 - Nov 2014 (2 nights): g,r 36min exp., 2 fields, ~1hr cadence
no follow-up
 - May 2015 (1 night): g,r 6min exp., 13 fields, ~1.5hr cadence
follow-up cancelled
 - Aug 2015 (1 night): poor weather
 - Mar 2016 (2 x 0.5 nights): poor weather
 - Jun 2018 (4 x 0.5 nights): cancelled
- Multi-band Subaru Survey for Early-phase SNe Ia
 - PI: Jian Jiang
 - Apr 2016 (1.5 nights): g,r 1.5-2min exp., 35 fields, ~1-2hr
cadence

Pick-up results -rapid rising transients-

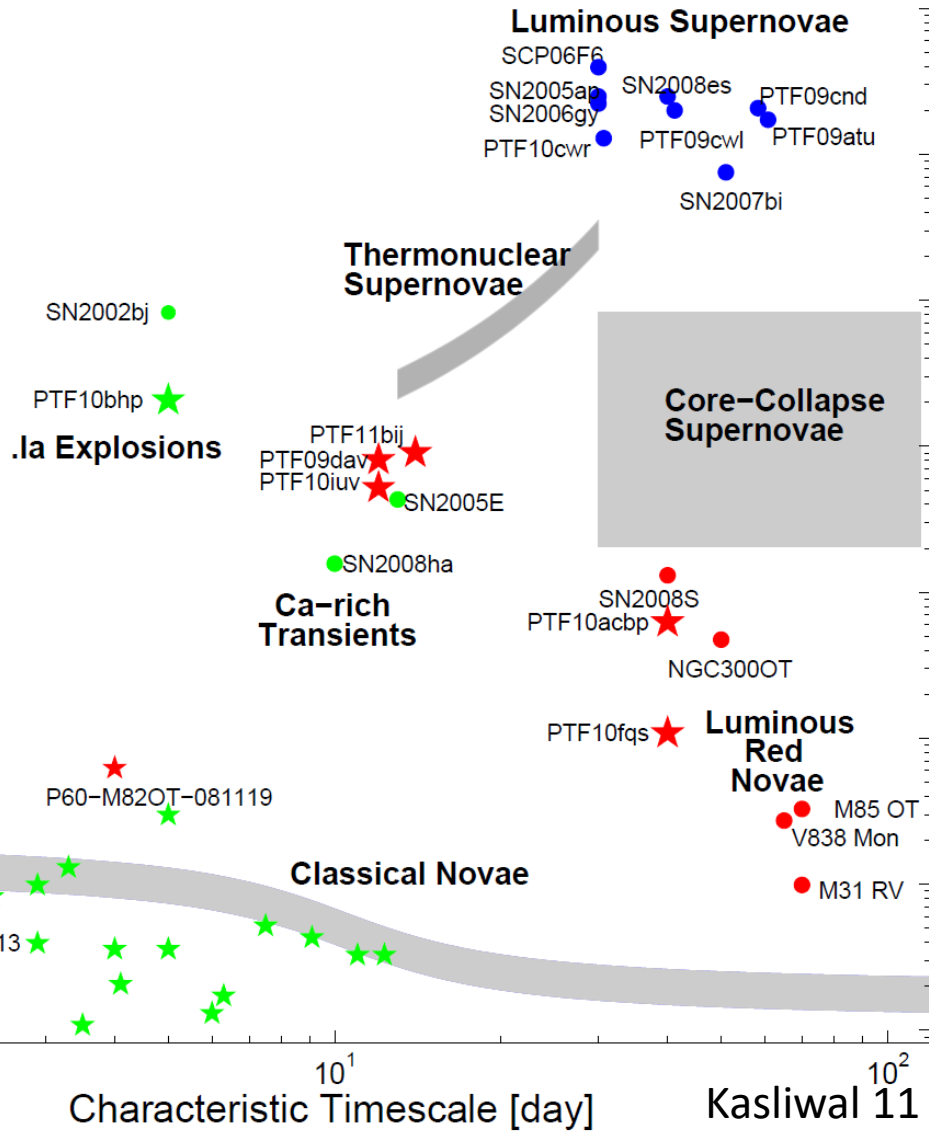
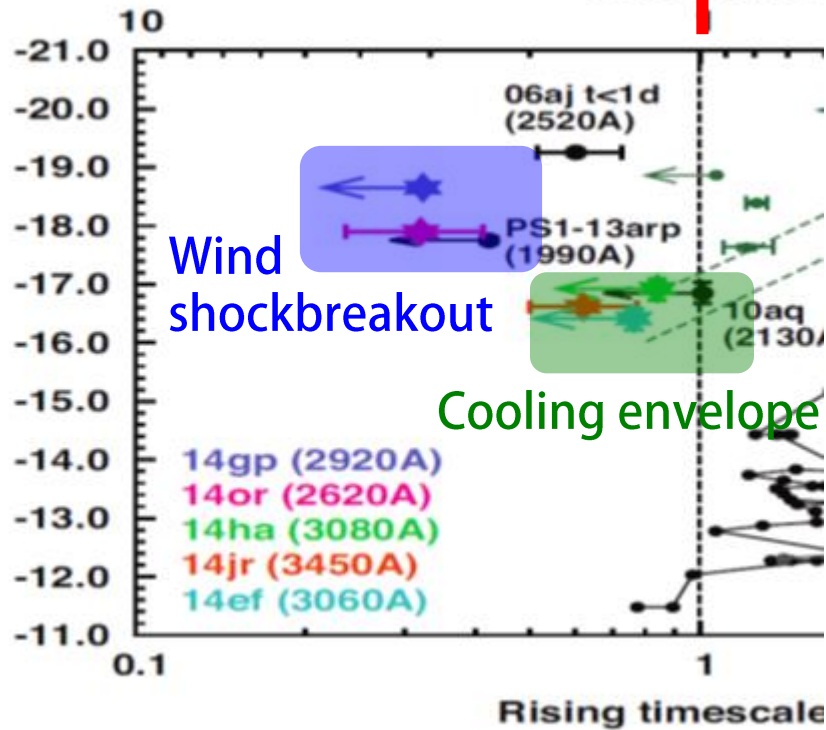
5 rapid-rising transients were discovered in 2 nights.



Pick-up results -rapid rising transients-

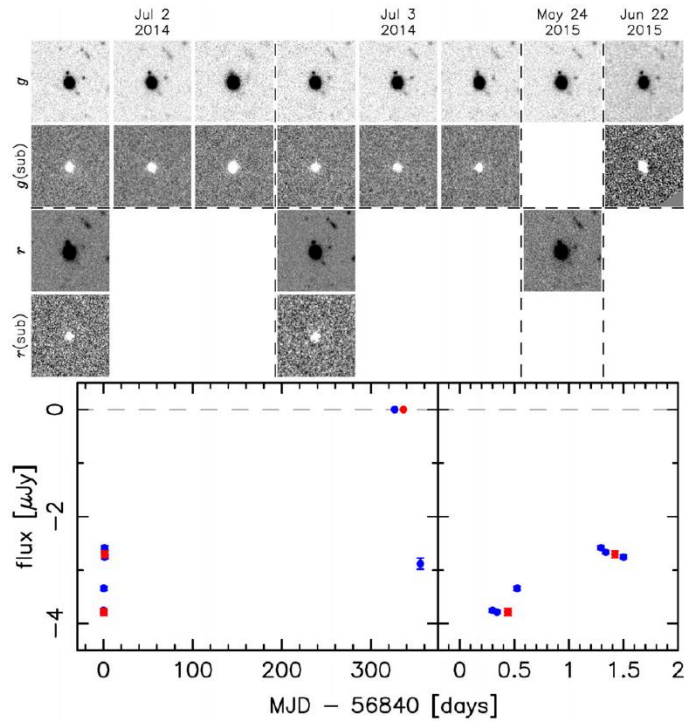
< 1 day

$|\Delta m / \Delta t|$ (mag)



The event rate of rapid-rising transients is $> \sim 9\%$ of the core-collapse supernova rate.

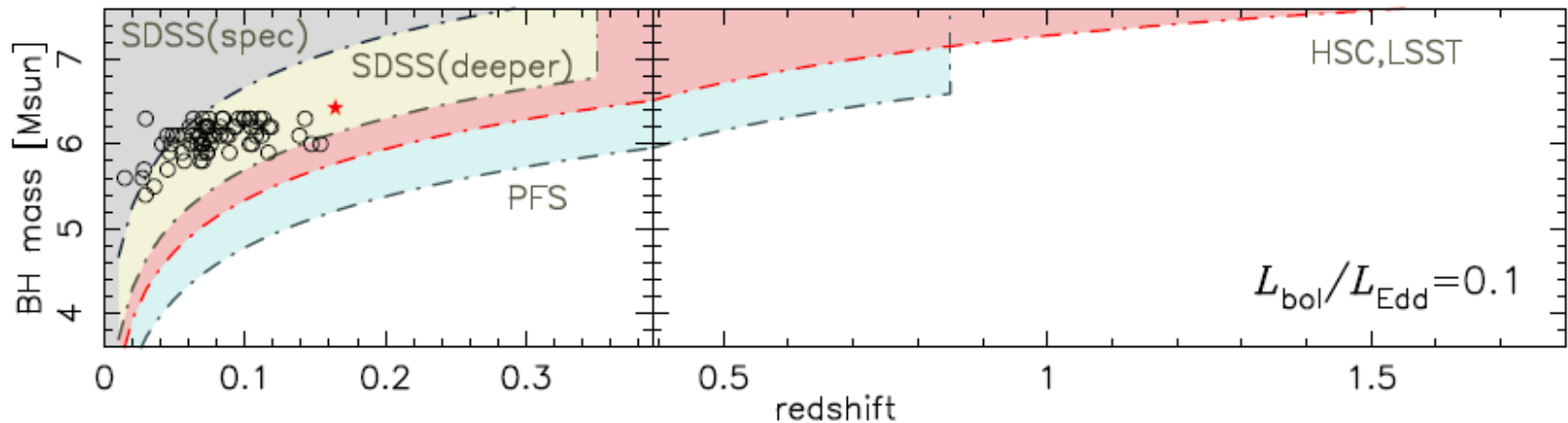
Pick-up results -low-mass AGN-



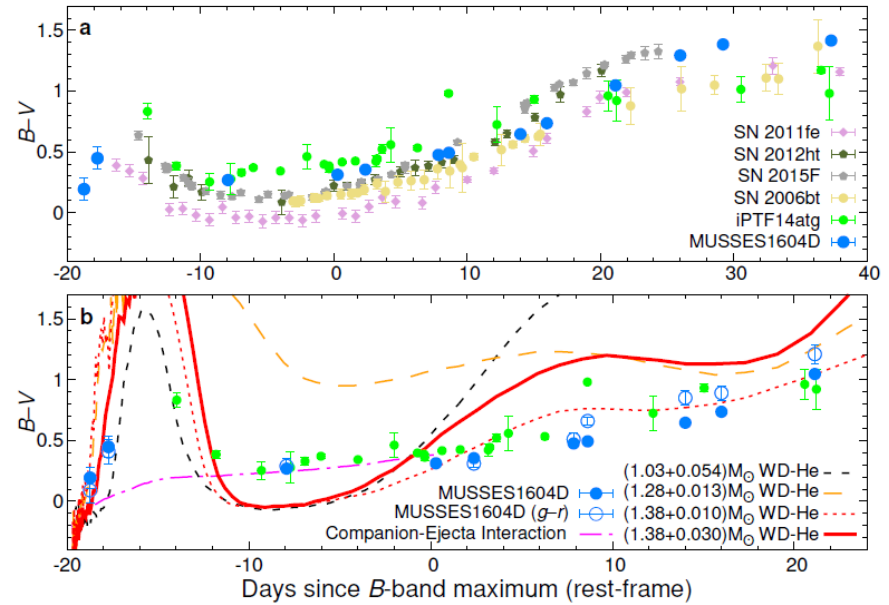
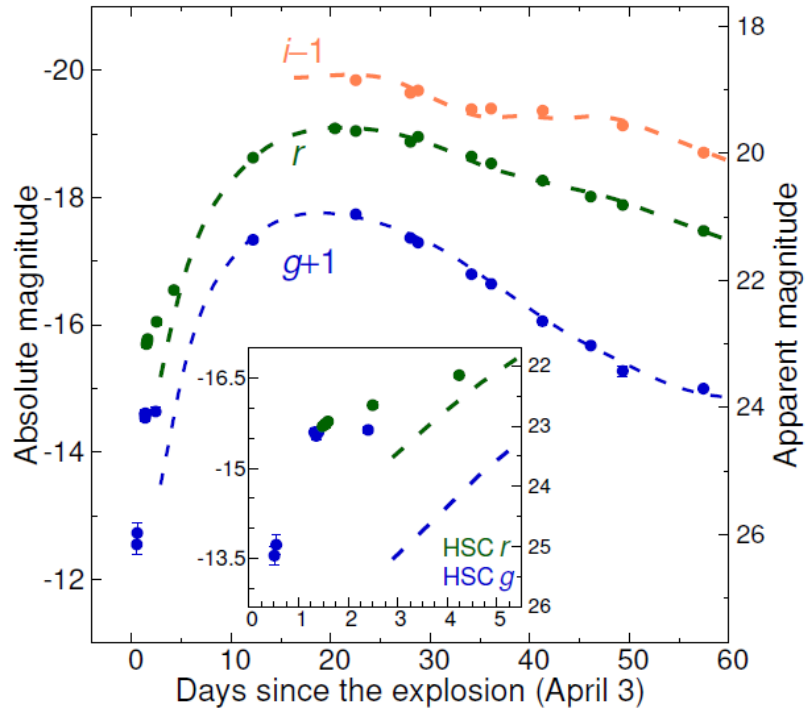
The high-cadence observation enables to select low-mass active black holes (BHs) at galaxy centers. Spectroscopic follow-up observation identify an active $2.7 \times 10^6 M_{\text{sun}}$ BH at $z = 0.164$.

Fig. 5. Detectable BH mass as a function of redshift.

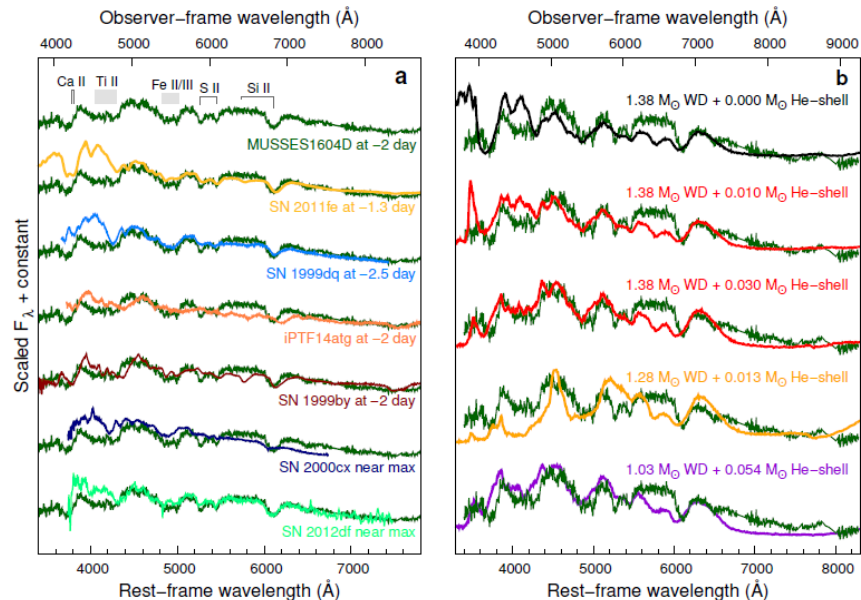
An Eddington ratio of 0.1 is assumed. A variability amplitude of 10% is also assumed.



Pick-up results -A hybrid SN Ia-



A red optical flash at ~ 0.5 days after explosion can be explained with a SN explosion triggered by a detonation of a thin helium shell.



Future plan

- HSC-SSP: transient survey in SXDS
 - Aug 2019 – Jan 2020
 - 3 fields in deep (4 month), 1 field in UD (6 month)
 - high-cadence in g-band (18 fields nights)
- Openuse
 - 1hr cadence survey
 - May 1-4, 2019: 4 x 0.5 nights (+GMOS follow-up 2 nights)
 - 30 sec exp. in g (<25.1mag), 60 fields (106deg²)
or 1min exp. in g (<25.5mag), 45 fields (80deg²)
- More future
 - HSC-intensive or HSC-SSP-2 for a transient survey?