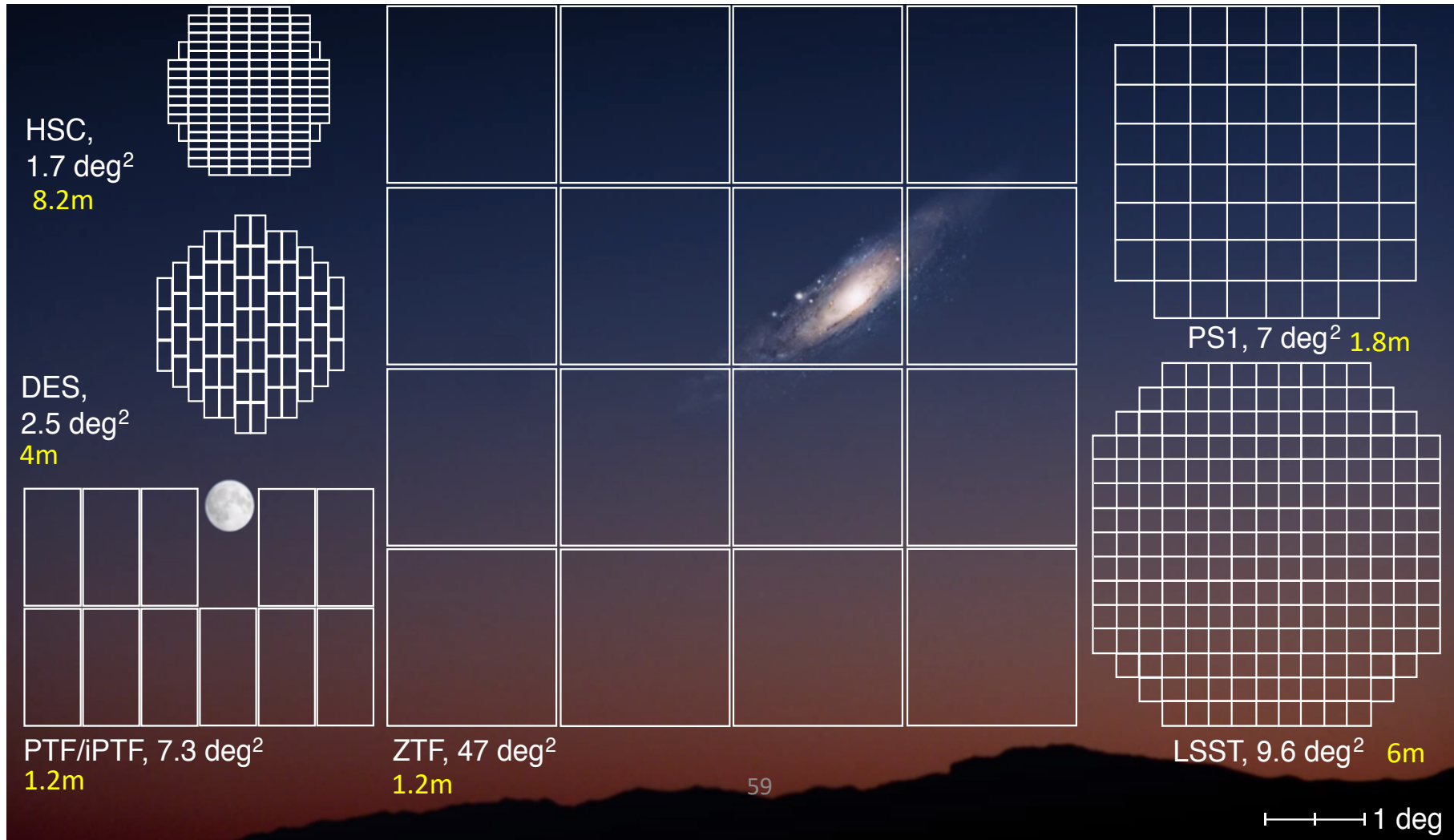


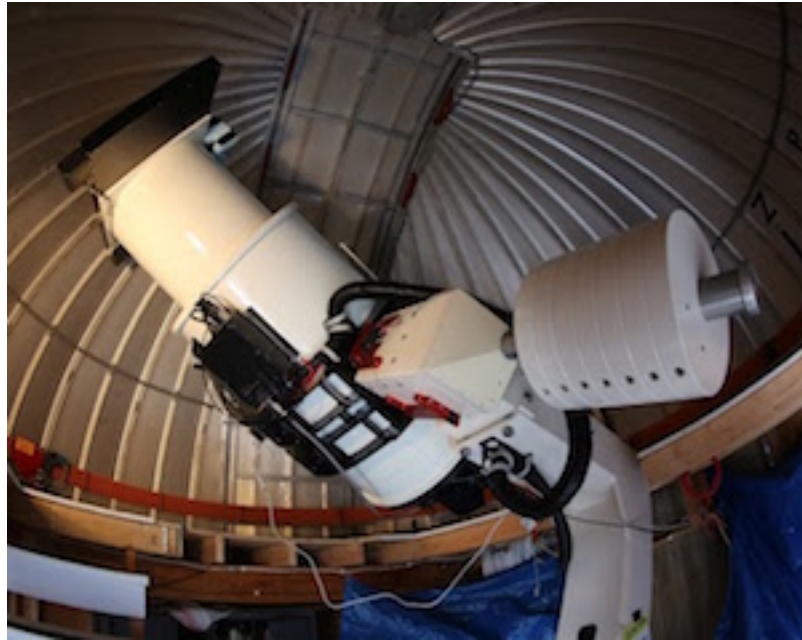
The Landscape of Time Domain Surveys

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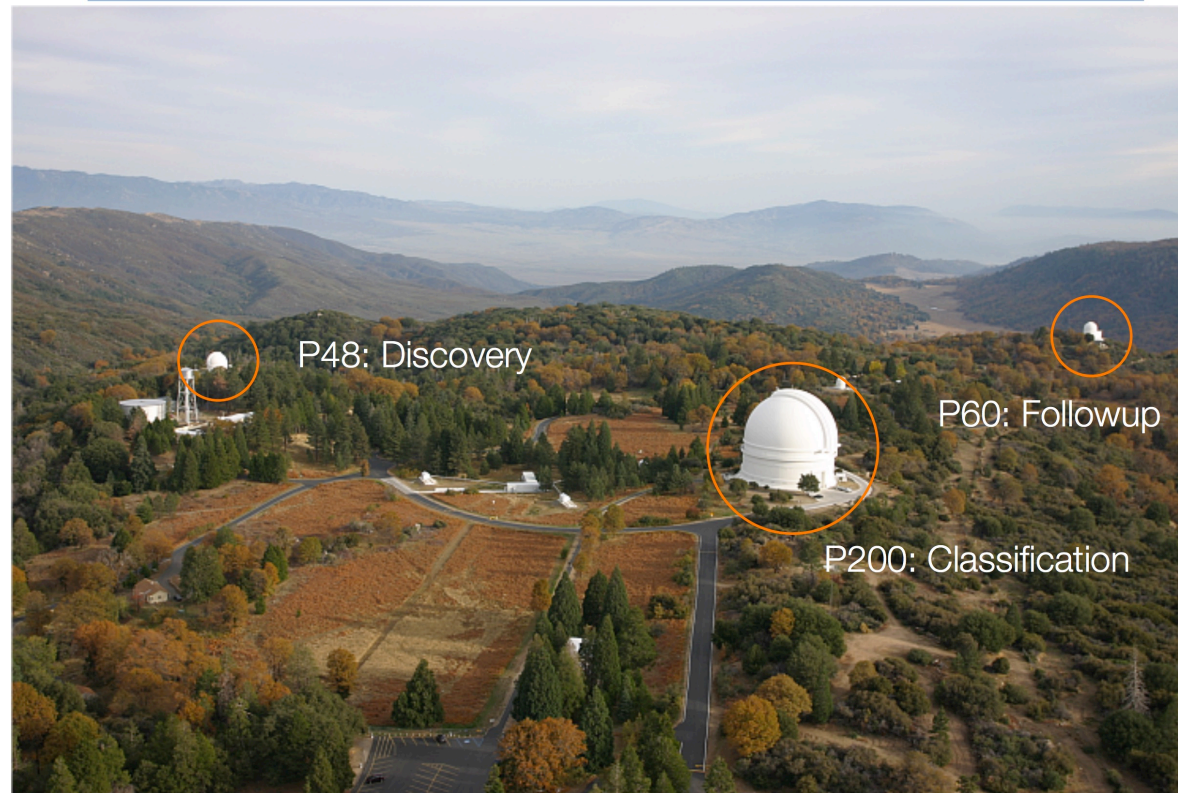
A collage of facilities





Zwicky Transient *Facility*

PTF uses the Palomar network of telescopes.



Moderate aperture survey matched to followup resources.

Dedicated large area surveys

- ASAS-SN (telephoto lens)
 - North & South, several longitudes
 - Covering all night sky with limiting mag of ~ 18 mag
 - Transients are announced over ATELS
- ATLAS (50-cm Schmidt telescopes)
 - Tuned for NEO, limiting mag about 19 (or 19.5, depending on broad color)
 - Northern sky surveyed over 2 nights (with intra-night cadence for NEOs)
 - Transients are announced via ATELS
- ZTF (1.2-m Schmidt telescope)
 - Northern Sky, multiple surveys, 20.5 mag (5 sigma; R and g band)
 - Built-in capability for classification (SEDM)
 - Prototype for LSST: industrial approach (alerts, archives etc)
 - MSIP Northern Sky Survey (3 nights, g&R)
 - Galactic Plane Patrol (nightly survey in g&R band)
 - Moderate cadence partnership survey (2000 sq deg; 3+3 times a night)
 - 2-hour deep drilling fields

Other Dedicated Facilities

- PS-1 (1.8-m)
 - Dedicated entirely for NEO studies
 - Some TDA (“foundation survey”)
- Kiso Schmidt (1-m)
 - Revolutionary CMOS camera (*)
 - 20 sq degrees behind a 1-m telescope
 - Has just been commissioned
- SkyMapper (1.35-m)
 - In Australia
 - Apparently no longer pursuing TDA
- BlackGem (65-cm telescopes)
 - In South Africa & Chile
 - 20 mag limiting mag
 - About to be commissioned

Dedicated focused projects

- Small field (Microlensing)
 - OGLE
 - MOA (*)
- Nearby SN
 - DLT40 and others (small telescopes)
 - KAIT
- SN activity has picked up a lot in China (*)
 - We will hear from XF Wang
- There are many focused projects (transits, stellar activity, GRB response)
 - Not discussed in my talk but there are opportunities here to use tens of 1-m class telescopes for transient follow up (cf. papers on AT2018cow)

Time Domain Surveys on Large telescopes (but time shared with other instruments)

- HSC on Subaru 8-m
 - 1.7 sq degrees, excellent imaging, as deep as LSST (or deeper)
 - Main focus has been large scale structure studies
 - In TDA main contribution has been fast transients (shock breakout of large stars)
- DECam on Blanco 4-m
 - 2.5 sq degrees, excellent imaging
 - Built for large scale structure & SN Ia cosmology
 - A large number of community led programs (SLSN, shock breakout, GW astronomy)

The classification bottleneck

- The local rate of SNe (core collapse & Ia) is $10^5 \text{ Gpc}^{-3} \text{ yr}^{-1}$
- The bright surveys are capable of finding 10^4 SNe per year
- In order to classify a transient you need about 100 channels of information
 - This can be spectral (SEDM approach) or light curves (Ia cosmology approach)
- In order to determine the redshift of the transient you need 1000 channels of information
 - Low resolution spectroscopy ($R=1000$) is robust. Broad band colors is adequate for some purpose (but works for higher redshift)

Enter “Massively Multiplexed Spectrographs”

- A symbiotic relation between time domain astronomy surveys and massively multiplexed spectrographs
- Use the next few years to train on TDA+DESI, TDA+SDSS V and establish large SN samples
 - Use this training set to establish filters for LSST+PFS
- Then there is the entire area of stellar & solar-system astronomy
- With Gaia and other data sets and highly cadenced ZTF data new avenues for Galactic black hole searches, compact binaries and in general binaries is a rich field.