Discussion: TDA+MMS for GW astronomy

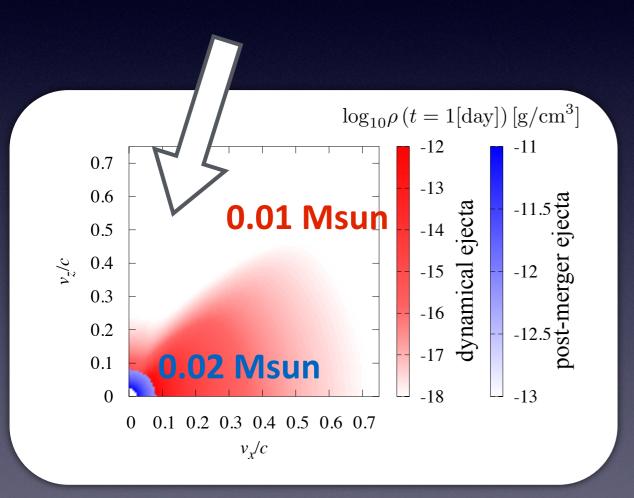
Masaomi Tanaka (Tohoku University)

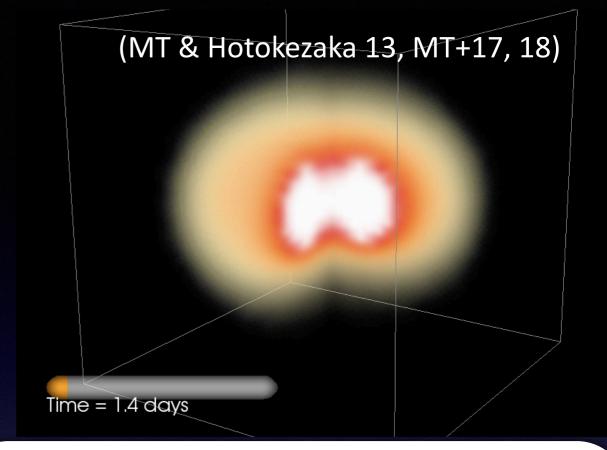
Key questions

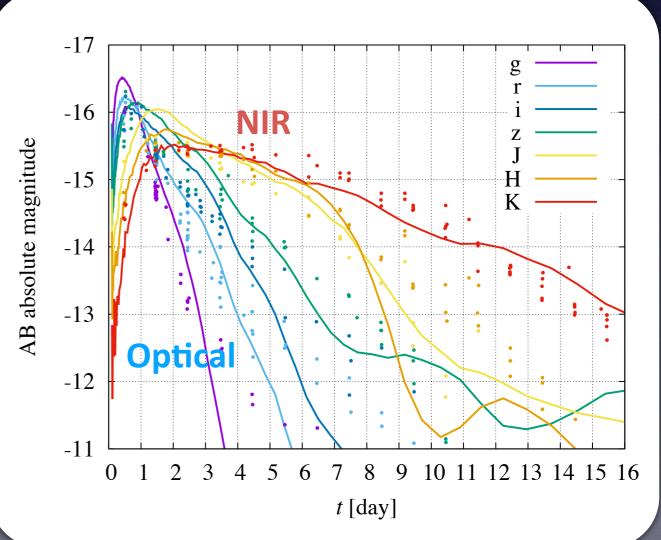
- Physics of NS merger
 - Mass ejection mechanism
 - Central remnant <= EOS of high-density matter
- Nucleosynthesis by NS mergers
 - Variety by NS mass, mass ratio?
- What about BH-NS merger?

...

Radiative transfer simulations for NS merger

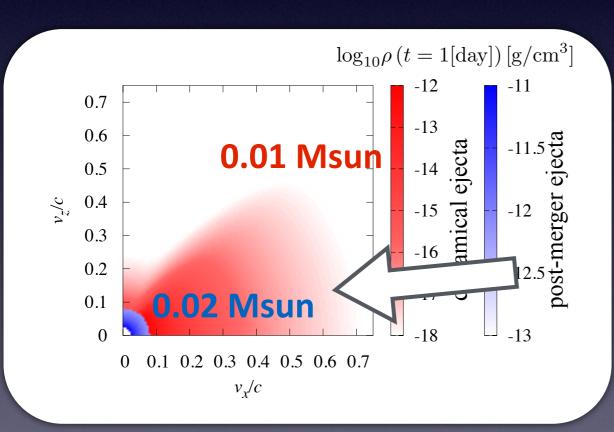


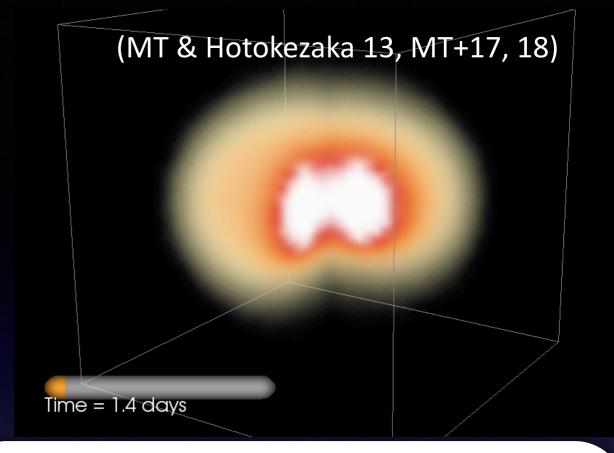


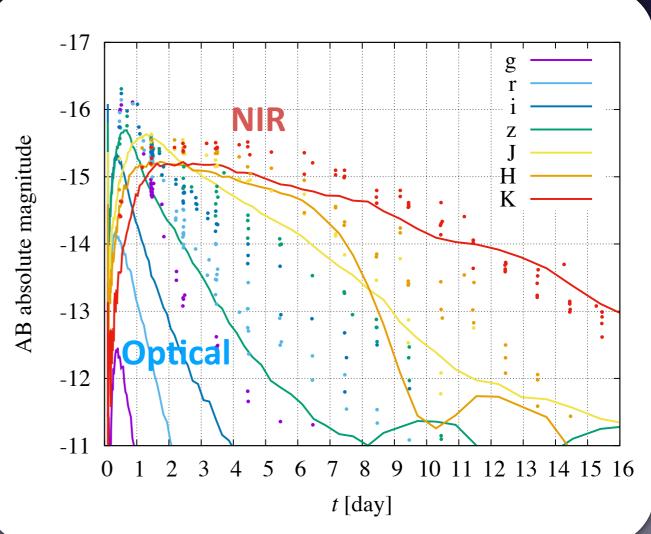


Kawaguchi, Shibata, MT 2018

Radiative transfer simulations for NS merger

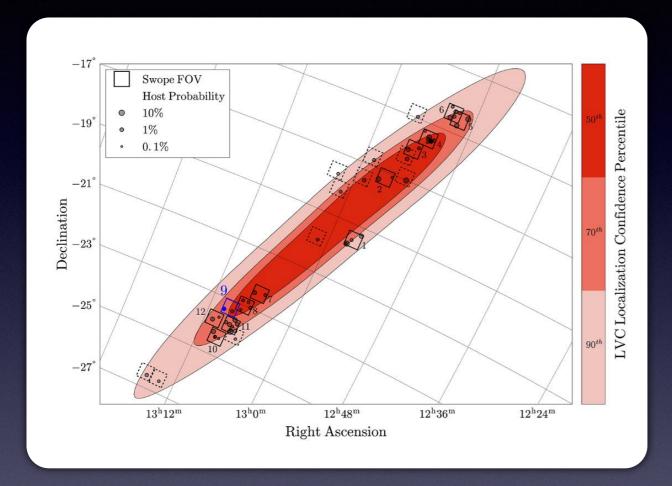






Kawaguchi, Shibata, MT 2018

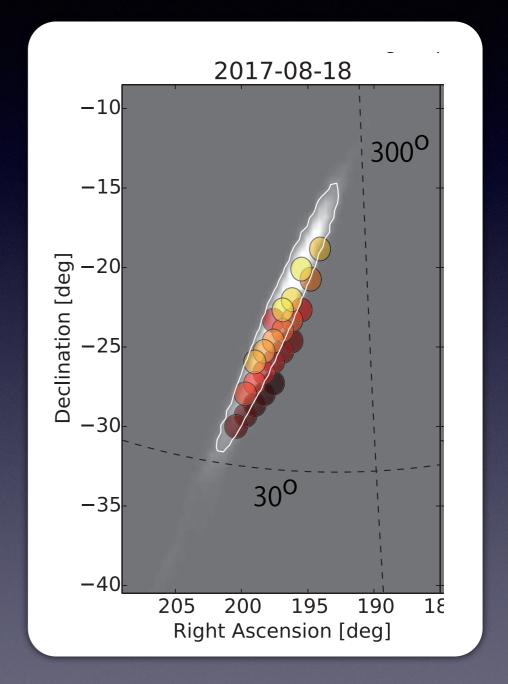
Targeted search



Coulter+17

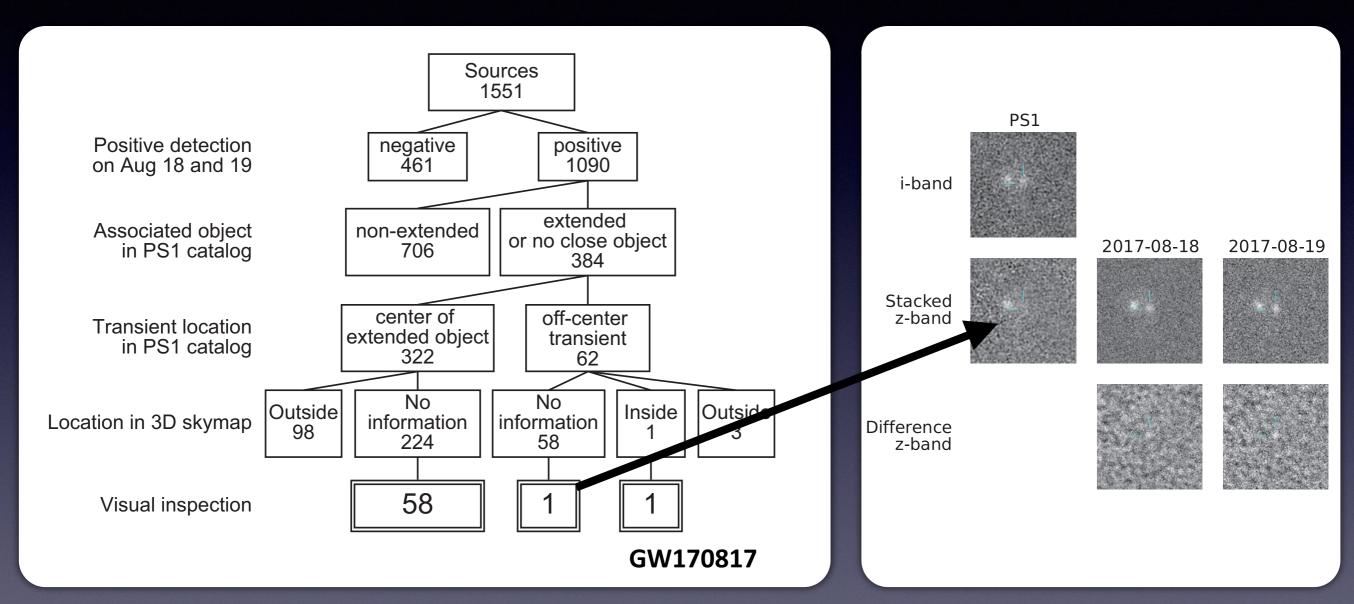
Not complete beyond >~50 Mpc

Wide field search



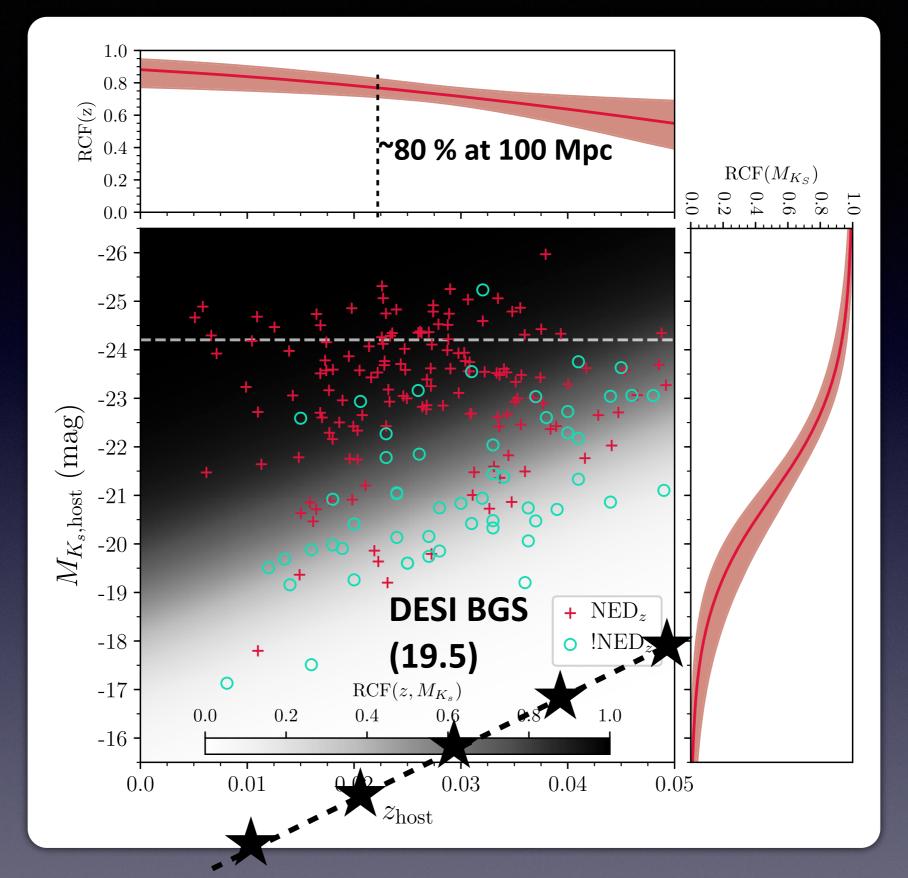
Tominaga, MT, Morokuma+18

Wide field survey Subaru/HSC 20.6 mag for 23.6 deg²



Tominaga, MT, Morokuma+18

MMS: Spectroscopy for ALL the candidates



Kulkarni+18

MMS: Complete galaxy catalog up to 200 Mpc

Role of MMS for GW astronomy

- Spectroscopy of all the transient candidates in the GW localization area
 - DESI? PFS?
 - LAMOST? (fixed along the meridian)
- Complete galaxy catalog up to 200 Mpc
 - DESI BGS
 - Targeted survey wins wide-field survey?
- Anything else?