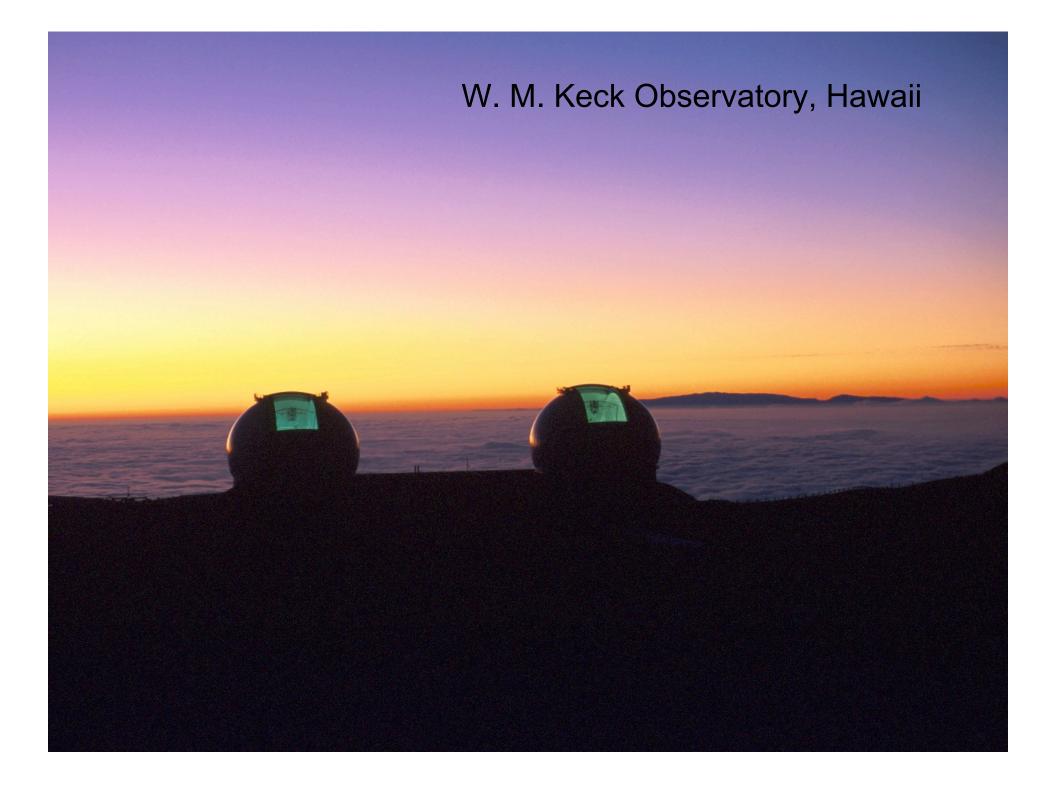
OPTICAL & RADIO TRANSIENTS

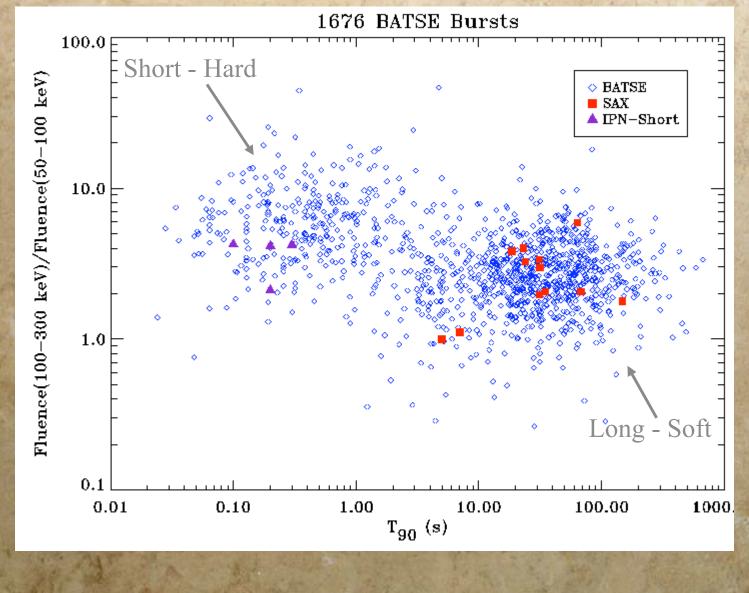
S. R. KULKARNI CALIFORNIA INSTITUTE OF TECHNOLOGY DIRECTOR, CALTECH OPTICAL OBSERVATORIES CHAIRMAN, SPACE INTERFEROMETRY MISSION



Taking the Measure of the Universe... Planets, Parallaxes & More

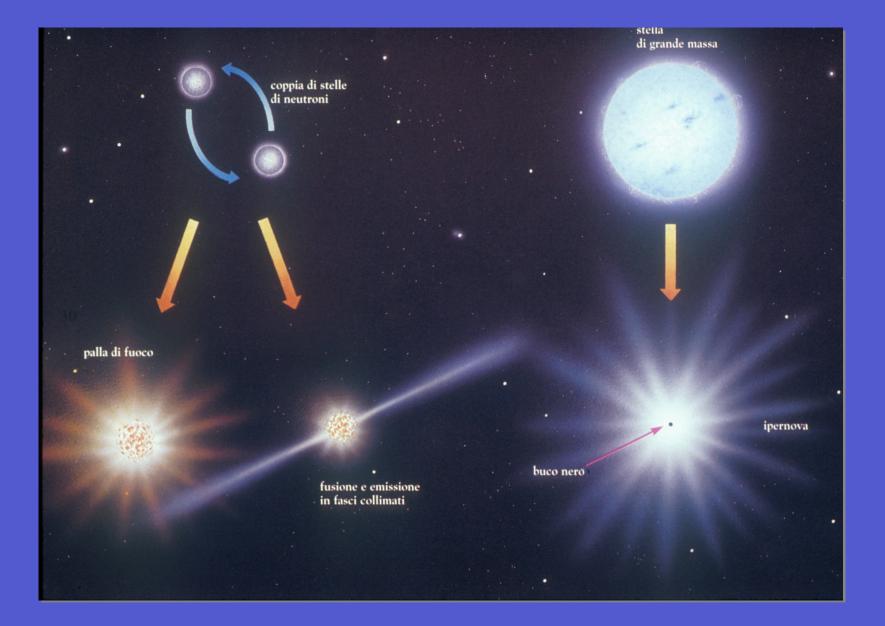
Space Interferometry Mission, PlanetQuest

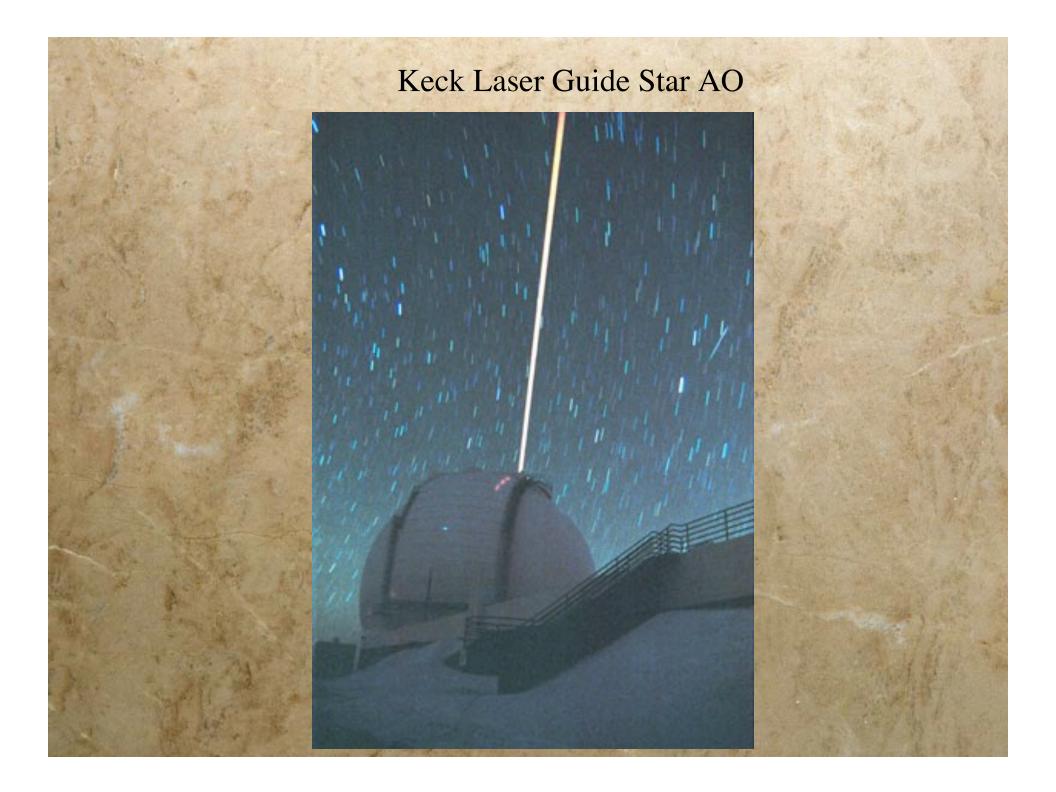
TWO CLASSES OF GRBS



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POPULAR MODELS





GRB 050724 Host Galaxy Keck/LGSAO/Narrow Camera K'-Band

Red elliptical z=0.258 L=1.6 L_{*} SFR< $0.03 M_{\odot} yr^{-1}$

Kulkarni & Cameron

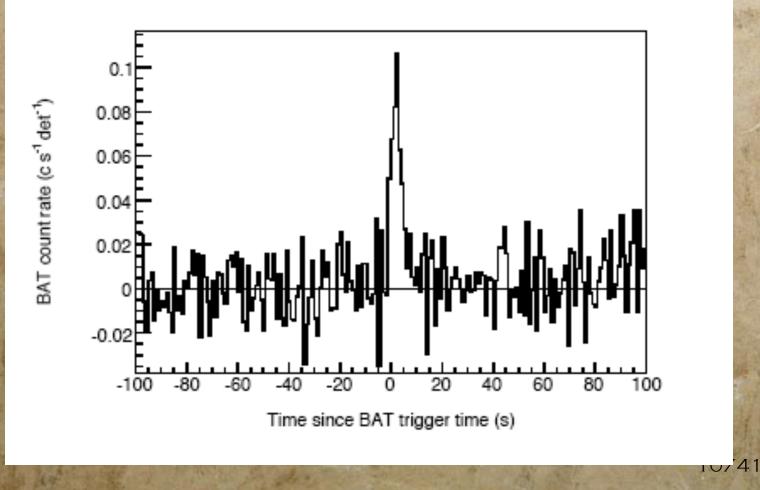
+

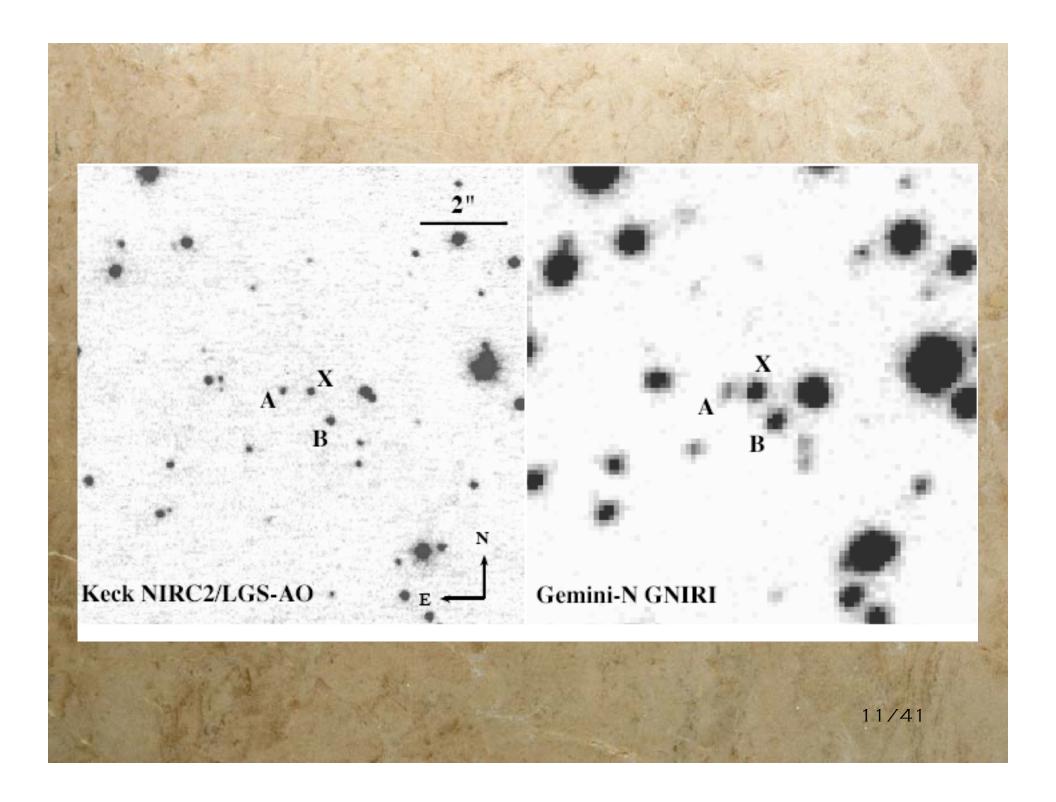
HST IMAGING & SEARCH FOR SUPERNOVA





GRB 070610 (THE MEEK SHALL INHERIT THE SKIES)





LESSONS LEARNT

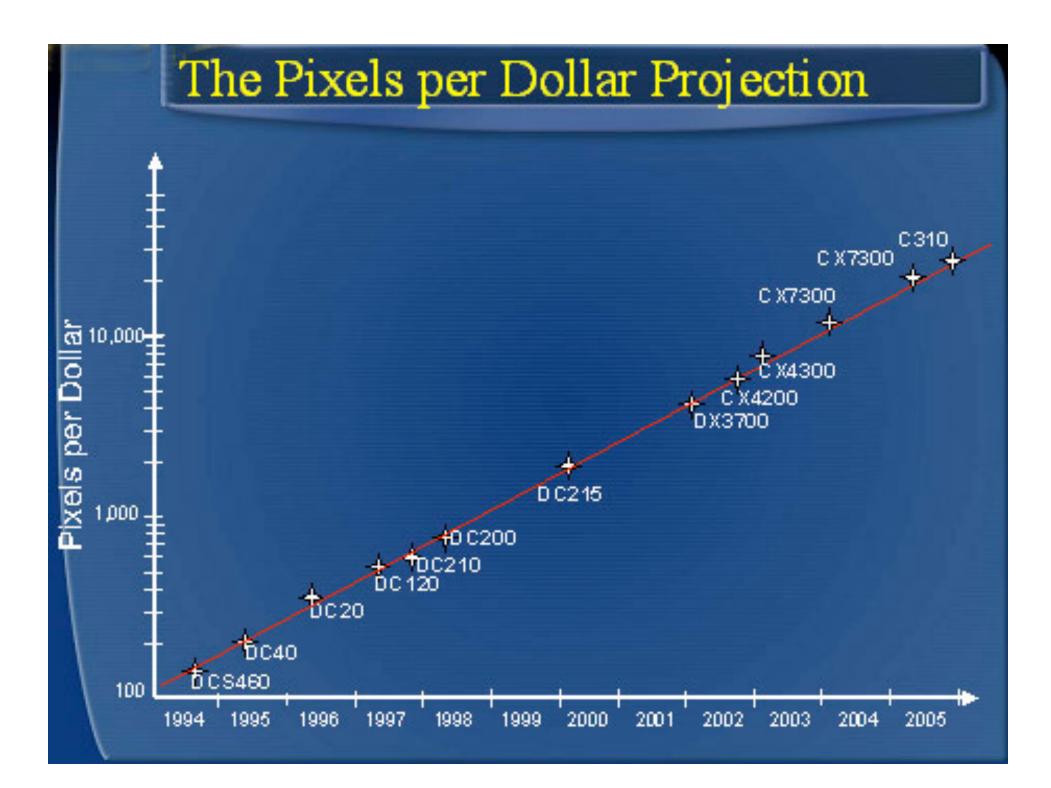
 AMBITIOUS AND RAPID MULTI-WAVELENGTH EFFORT IS THE KEY TO RAPID PROGRESS

THERE IS TREMENDOUS OPEN PHASE

"WHAT IS NOT FORBIDDEN EXISTS"

THE NEW WILD WEST: TRANSIENTS

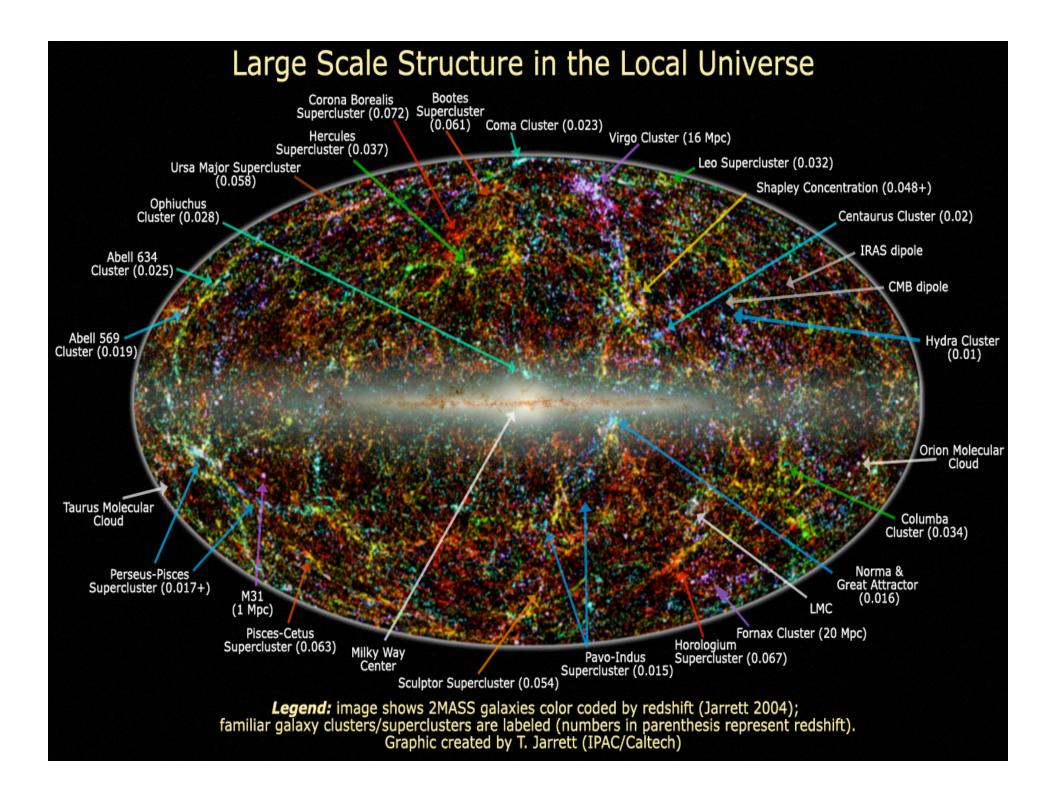
I. OPTICAL SEARCHES



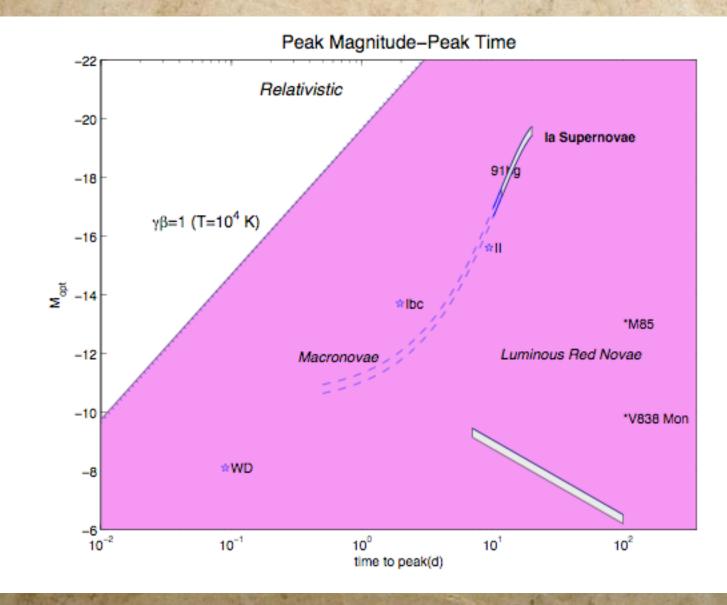
TRANSIENTS IN THE LOCAL UNIVERSE

AN EMERGING FIELD

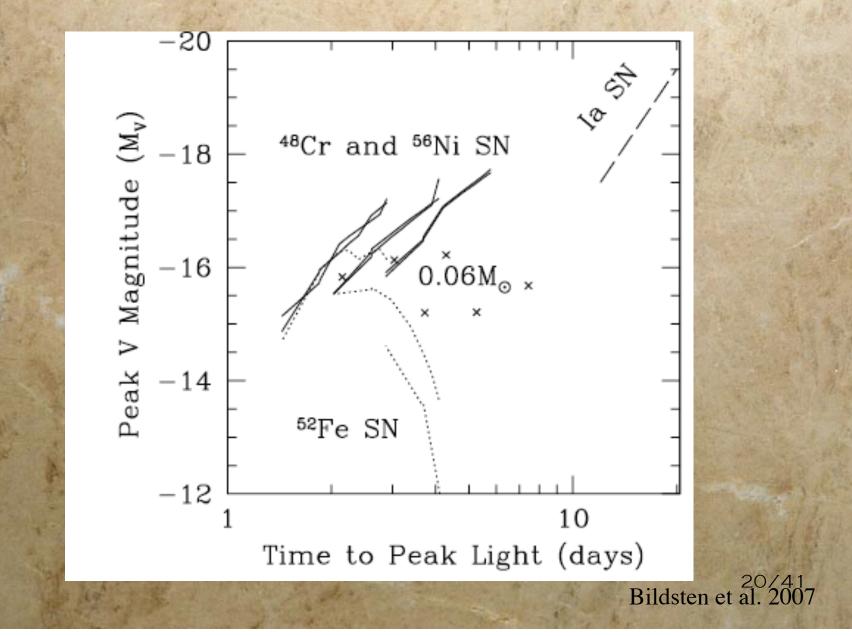
ULTRA-HIGH ENERGY COSMIC RAYS
GZK CUTOFF (PROTON-PION)
TEV SOURCES
PAIR PRODUCTION ON CMB PHOTONS
GW SOURCES (100 HZ BAND)
ENAHANCED LIGO (2009)
ADVANCED LIGO (2013)
TRANSIENTS IN THE NOVA-SUPERNOVA GAP



Focus: Objects in the Nova-Supernova Gap



Helium Flashes (0.11a Supernovae)



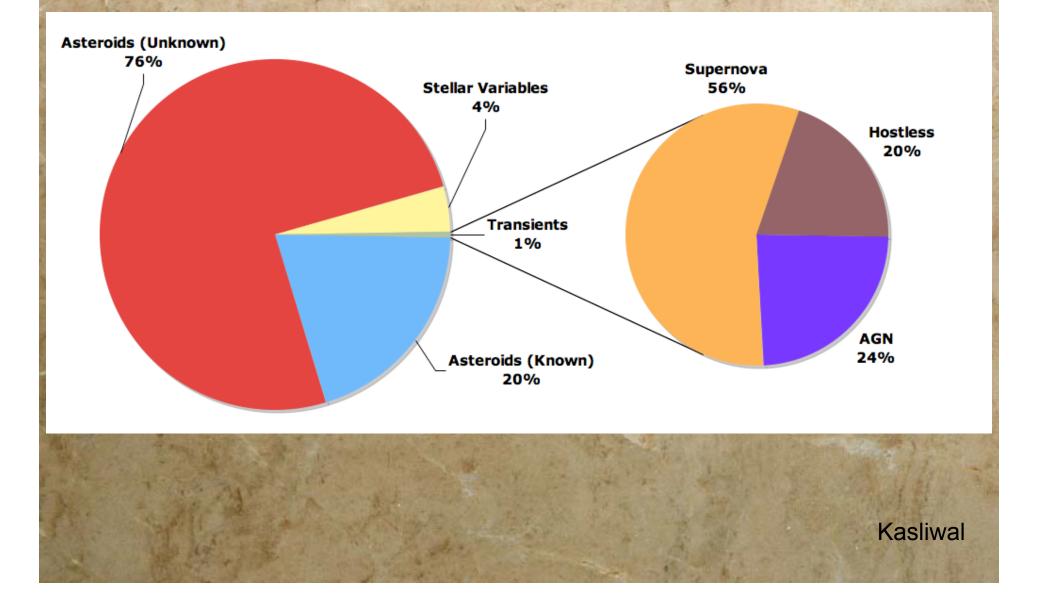
FINDING FAST TRANSIENTS AIN'T EASY

THE DEEP LENS SURVEY TRANSIENT SEARCH. I. SHORT TIMESCALE AND ASTROMETRIC VARIABILITY

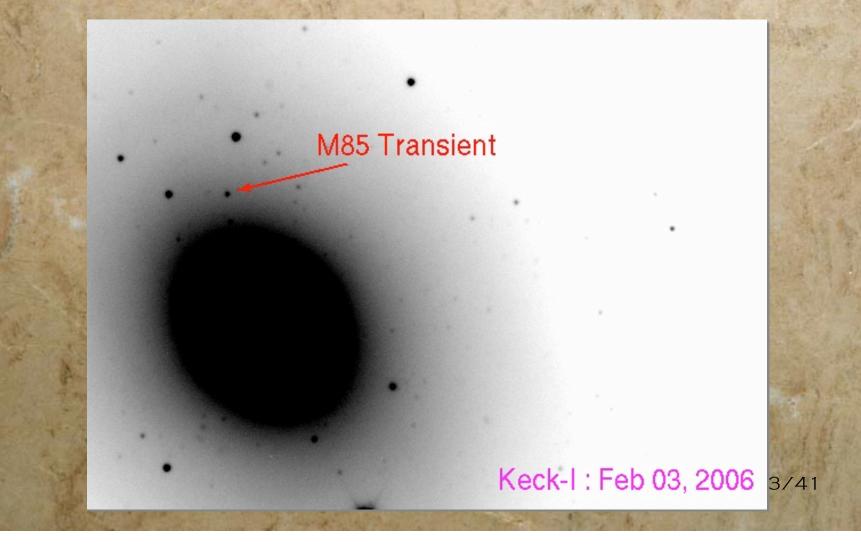
A. C. BECKER,^{1,2,3} D. M. WITTMAN,^{1,4} P. C. BOESHAAR,^{4,5} A. CLOCCHIATTI,⁶ I. P. DELL'ANTONIO,⁷ D. A. FRAIL,⁸ J. HALPERN,⁹ V. E. MARGONINER,^{1,4} D. NORMAN,¹⁰ J. A. TYSON,^{1,4} AND R. A. SCHOMMER¹¹ Received 2004 March 4; accepted 2004 April 19

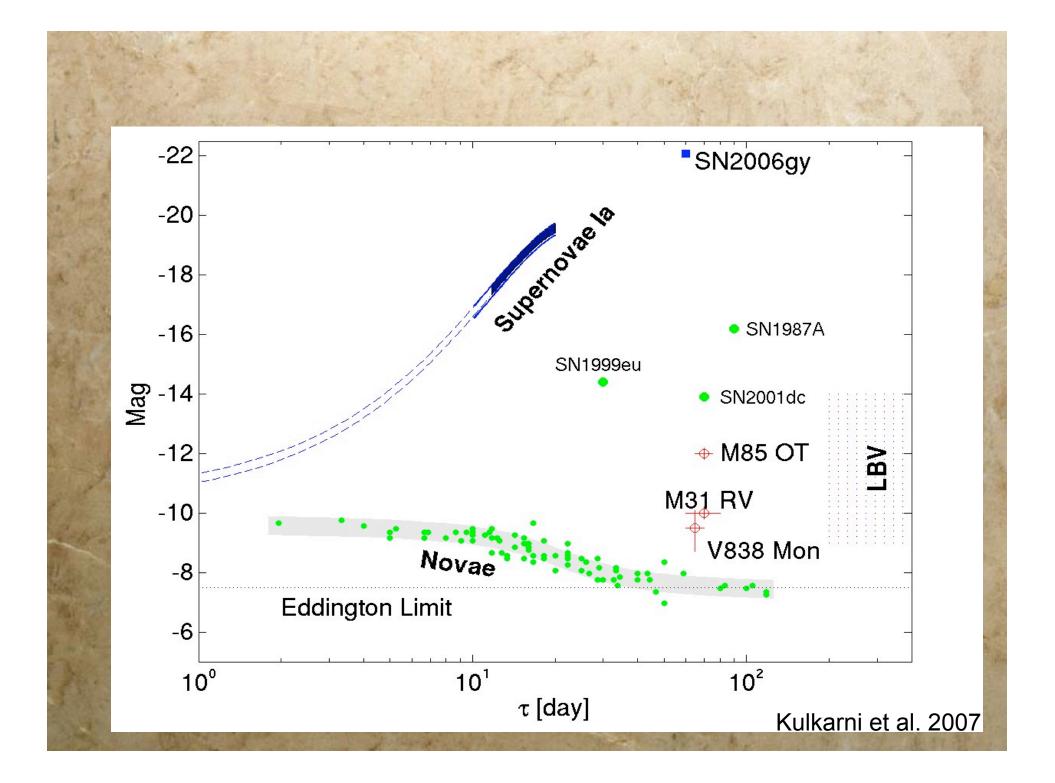


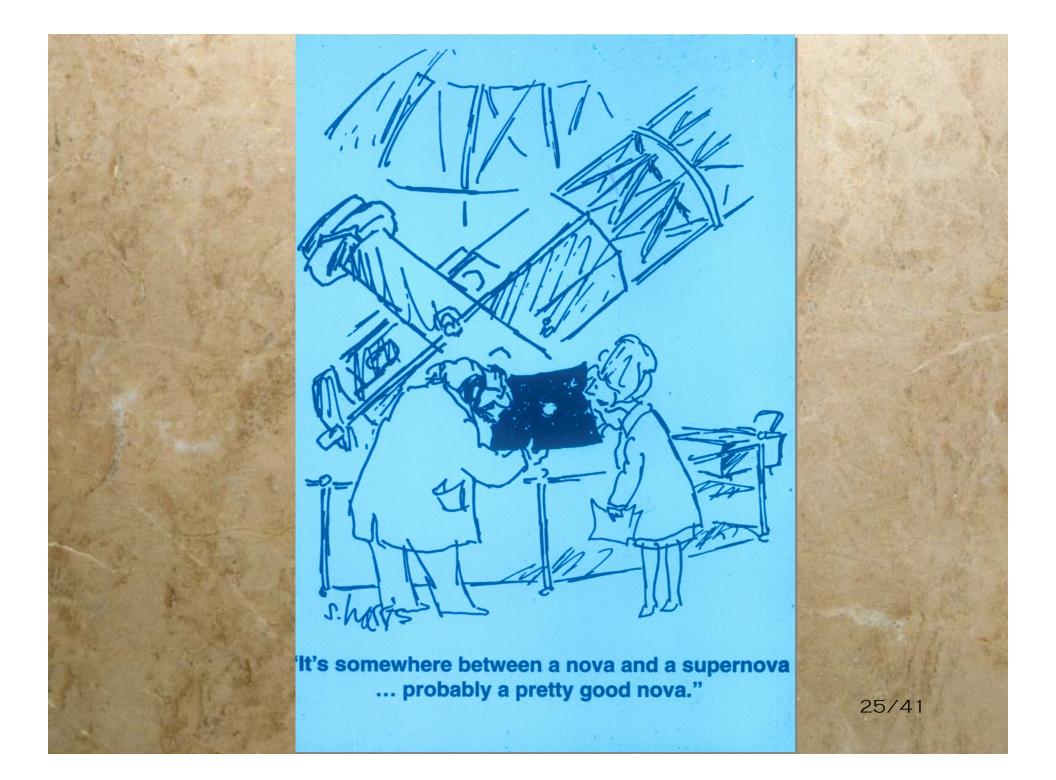
No dearth of Slow Transients either!

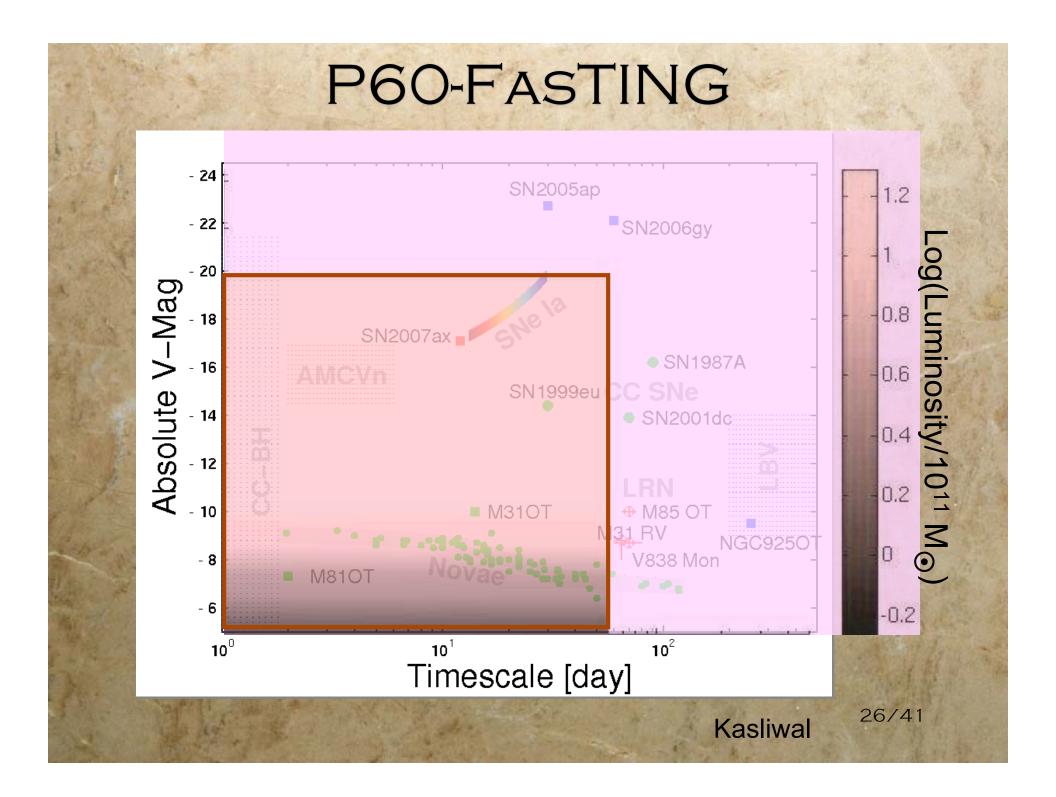


M85 OT2006-1: A REFRESHING START!

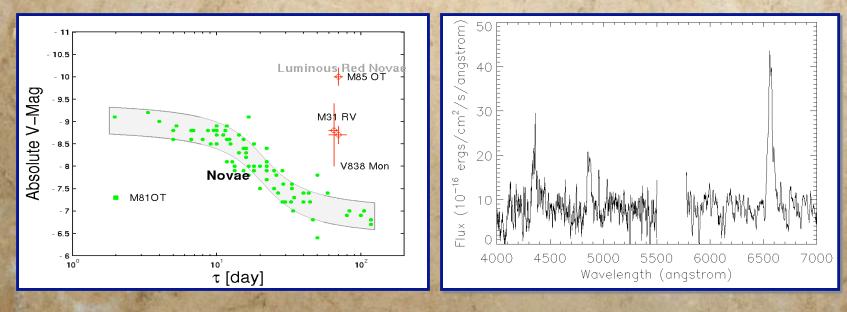








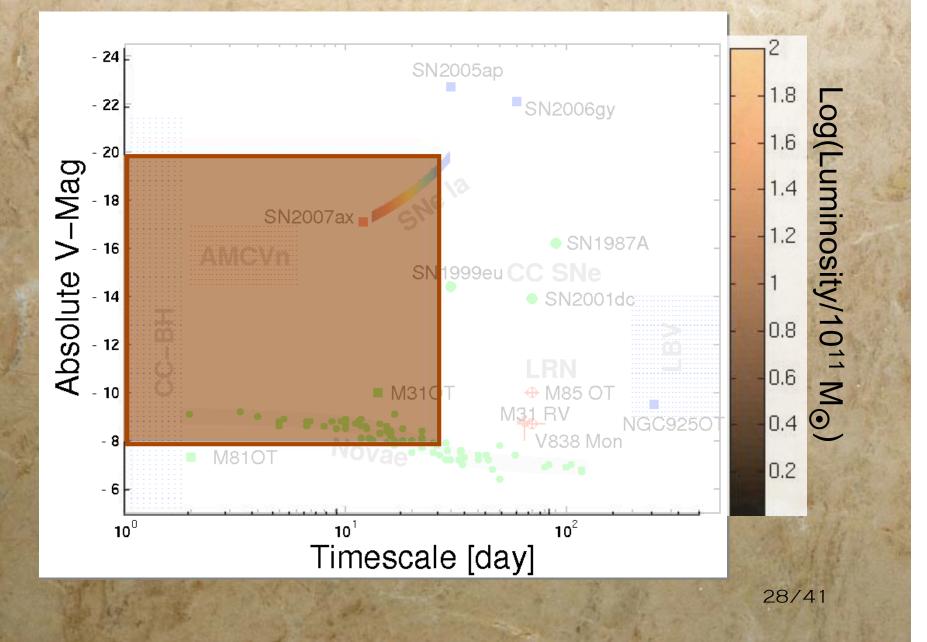
M81OT: A PECULIAR NOVA



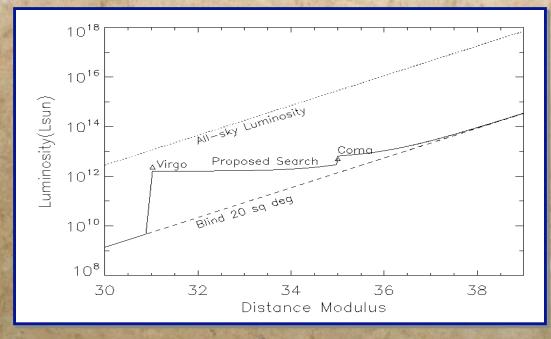
- A CLEAR DEVIANT FROM THE NOVA RELATION
- DISCOVERY OF NOVA IN M83 (STARBURST)
- DISCOVERY OF NEW LBVs
 - ARE THESE PROGENITORS OF PAIR PRODUCTION SNE?

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CFHT-COVET



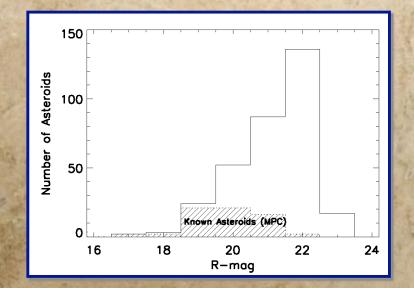
COMA & VIRGO EXPLORATION FOR TRANSIENTS



- CFHT: 3.6M + MEGACAM: 1 SQ DEG FOV @ 0.18"/PIX
- R-BAND, 60S X 2, DAILY CADENCE, DEPTH R~22
- WHY VIRGO? 25% OF UNIVERSE'S LIGHT @16 MPC IS IN VIRGO!
- PILOT PROJECT (2008): 7 SQ DEG; LARGE PROPOSAL (3 YRS): 20 SQ DEG
- PILOT PROJECT DISCOVERIES:
 - 7 TRANSIENTS (6 LIKELY SUPERNOVAE) WITH DATA ON 7 NIGHTS ONLY.
- REALTIME REDUCTION PIPELINE : <30min</p>

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LESSONS LEARNED



- ASTEROIDS : MUST BE FILTERED OUT (KINETIC SIGNATURE)
- GALACTIC M-DWARFS ARE PESTS (BLUE, RAPID, HALPHA)
- MOST "HOSTLESS" TRANSIENTS ARE DWARF NOVAE (COLOR)
 - 1% ARE UNIDENTIFIED AND NEW GALACTIC EXPLOSIVE TRANSIENTS
 - 1% ARE ENTIRELY UNKNOWN
- EXTRAGALACTIC SUPERNOVAE DOMINATE AT FAINT MAGNITUDES
- AGN NOT NEGLIGIBLE

KEY : ROBUST REAL-TIME PIPELINE AND WELL-DEFINED PROMPT MULTIBAND FOLLOW-UP STRATEGY NEED A CATALOG OF THE LOCAL UNIVERSE

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THE PALOMAR TRANSIENT FACTORY



N. LAW (PROJECT SCIENTIST) A. RAU (BOOK SCIENTIST) E. OFEK (CADENCE SCIENTIST) R. QUIMBY (PIPELINE SCIENTIST) IPAC (ARCHIVE CENTER) P. NUGENT (FAST PIPELINE) J. BLOOM (CLASSIFICATION & ALERT SCIENTIST) D. FOX (P60 AND FOLLOWUP SCIENTIST) S. KULKARNI (PRINCIPAL INVESTIGATOR)

PALOMAR TRANSIENT FACTORY

A full end-to-end facility dedicated for transients

TRIPLE CADENCE SURVEYS

- 1-DAY CADENCE
 - AM CVN
 - PECULIAR EVENTS IN NEARBY GALAXIES
- 3-DAY CADENCE
 - AGN BURPING & BLEATING
- 9-DAY CADENCE
 - SUPERNOVAE

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SCIENCE GOALS

- SUPERNOVAE: CORE COLLAPSE
 - ANONYMOUS (LOWER METALICITY) HOST GALAXIES
 - SYSTEMATIC STUDY OF IB AND IC SUPERNOVAE
 - RARE TYPES OF SUPERNOVAE (IIN, UNDERLUMINOUS)
- SUPERNOVAE: IA
 - COSMOLOGY (LUMINOSITY FUNCTION AT Z=0)
- CATCH SUPERNOVAE AT VERY EARLY TIMES

SCIENCE GOALS

TRANSIENTS IN NEARBY GALAXIES

- NOVAE
- LUMINOUS RED NOVAE
- LBV OUTBURSTS
- OTHER EVENTS (AIC, COALESCENCE, CE, 0.1 IA)

NUCLEAR BLACK HOLES:

- BLAZING
- BURPING
- TIDAL EVENTS

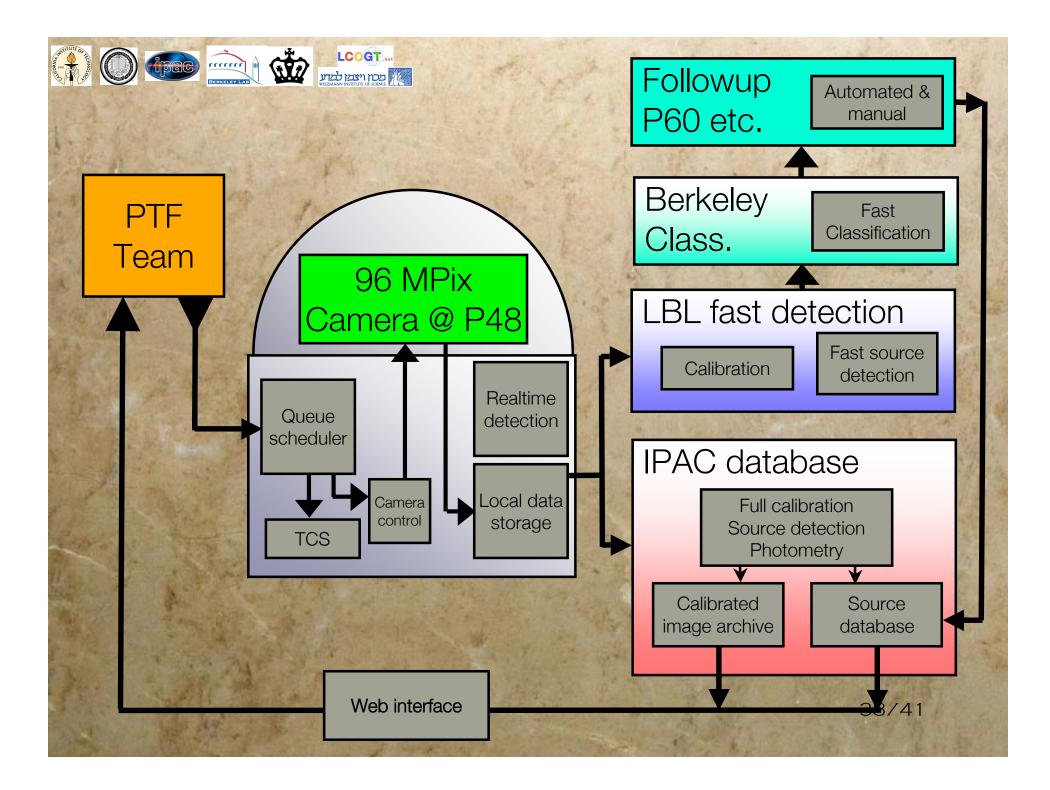
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SCIENCE GOALS

- GALACTIC TRANSIENTS:
 - AM CVN
 - HIBERNATING CVS (AND DWARF NOVAE)
 - FLARE STARS
 - PRE-MAIN SEQUENCE STARS
 - PLANETARY TRANSITS

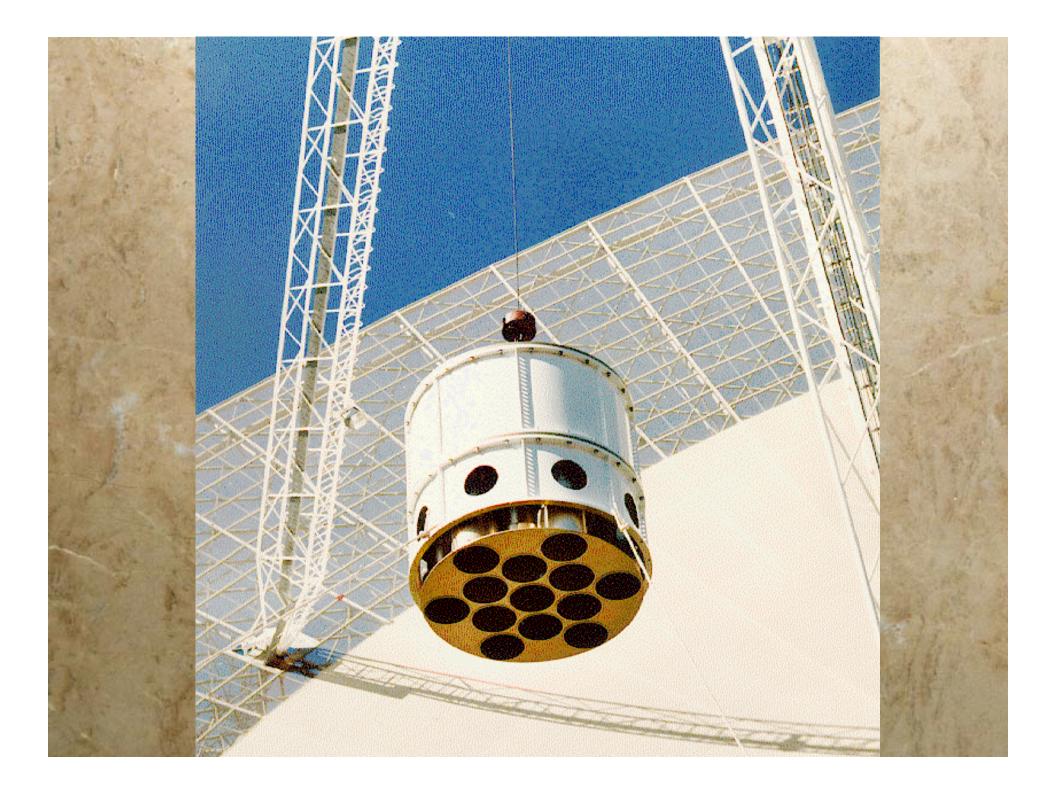
CFH12K DETECTOR

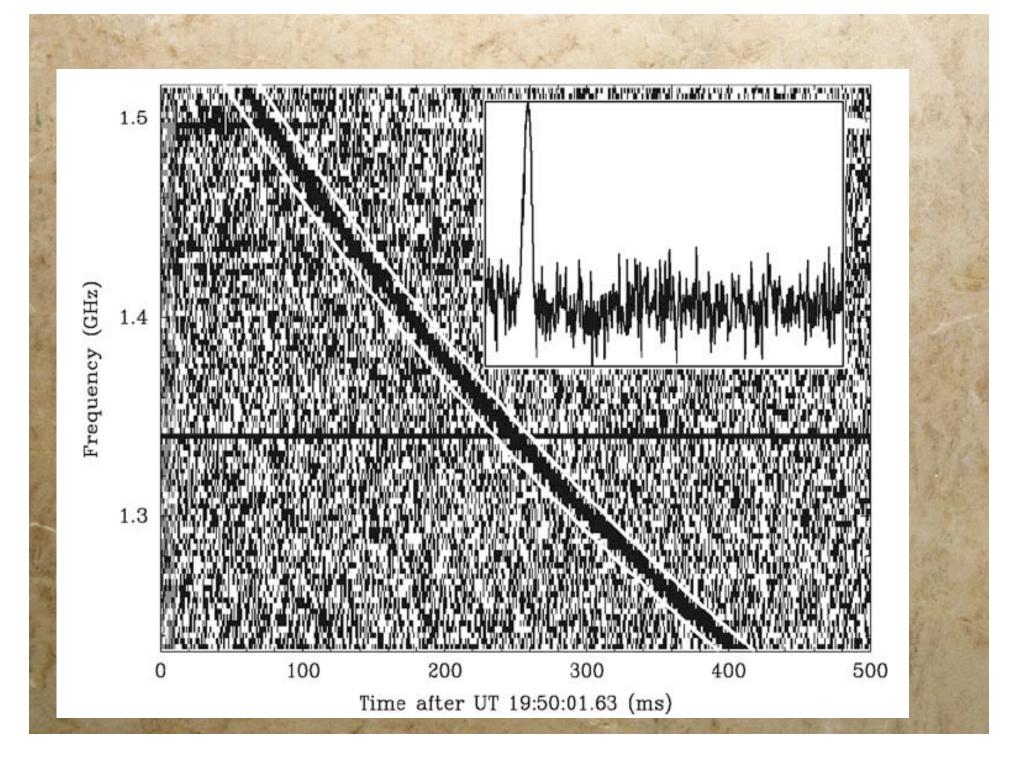
At Caltech now



II. HIGHLY MULTIPLEXED IMAGING AT RADIO WAVELENGTHS





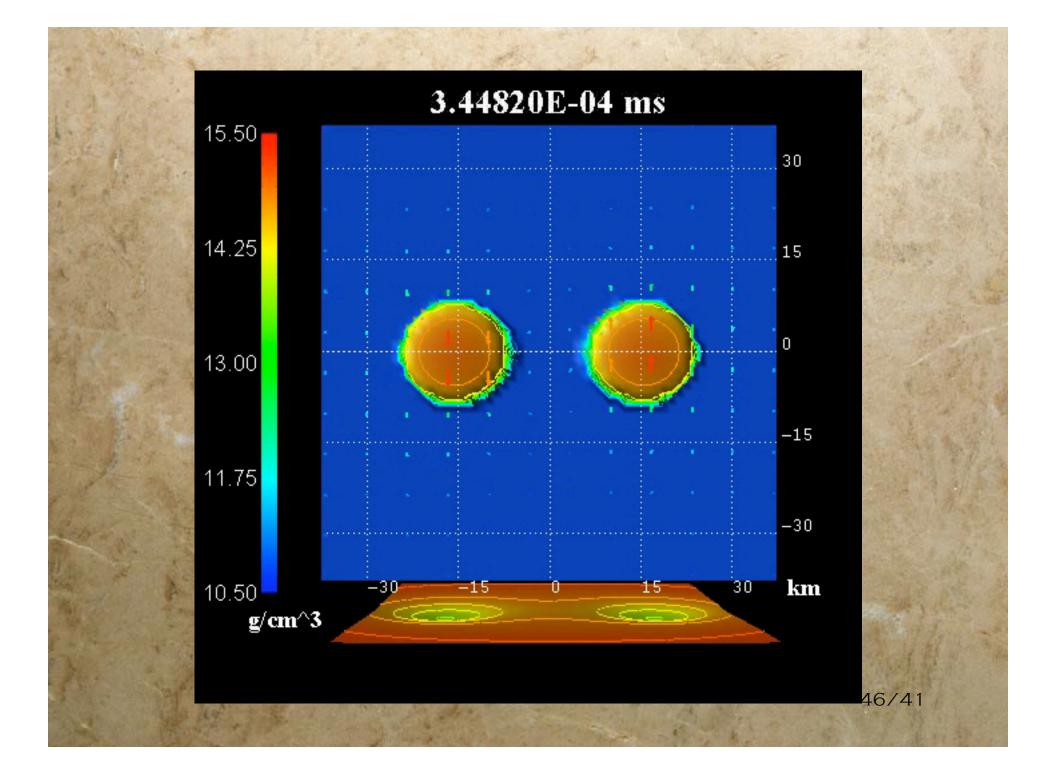


Sparkers cannot come from cataclysmic events

Class	Type	Φ	Ref
		$\mathrm{Gpc}^{-3}\mathrm{yr}^{-1}$	
LSB (low)	BC	100 - 1800	[1,2]
LSB (high)	Obs	1	[1]
	BC	100 - 550	[1]
SHB	Obs	> 10	[3]
	BC	$< 10^{5}$	[3]
SGR	Obs	$< 1.5 \times 10^4$	[4]
In-spiral	Th	$3 imes 10^3$	[5]
Core Collapse	Obs	3×10^5	[6,7]
Sparkers	Obs	732.7×10^{5}	[8]



III. NEW MESSENGERS



MACRONOVA MODEL

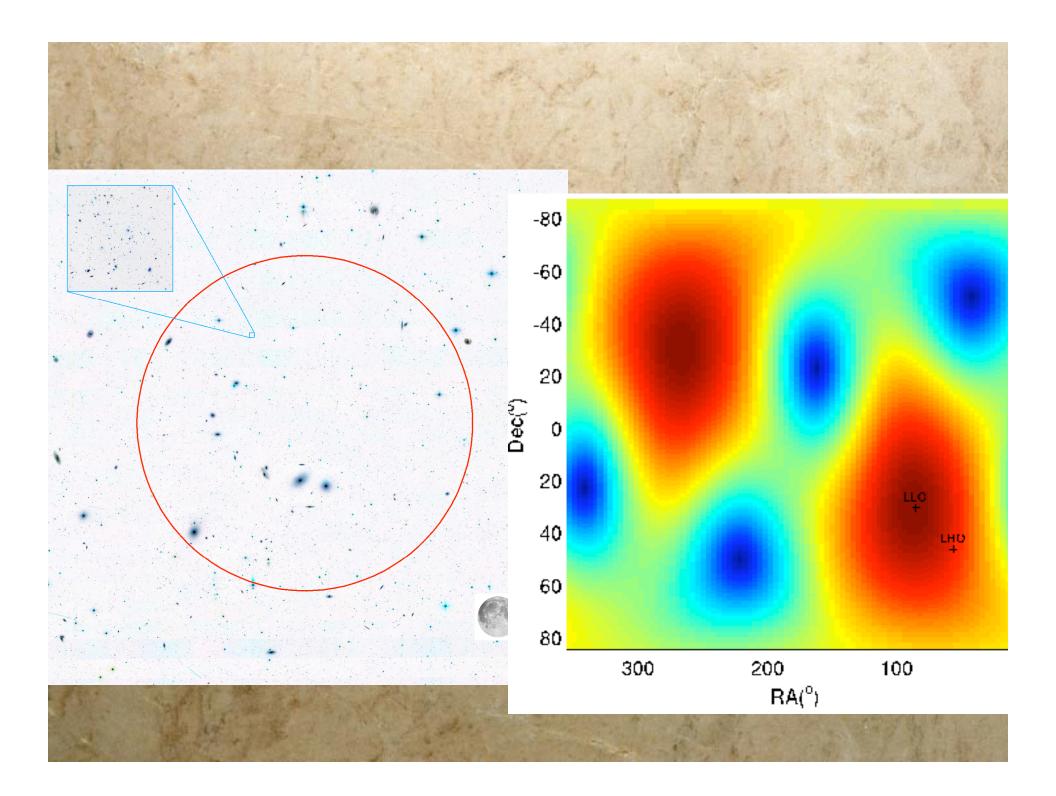
- PARAMETERS: $M_{EJECTA} \& V = \beta C$
- COMPOSITION
 - FREE NEUTRONS
 - RADIOACTIVE NICKEL
 - NEUTRON RICH MATERIAL (NON-RADIOACTIVE)
- INJECTION OF ENERGY ESSENTIAL FOR MACRONOVA TO SHINE AND BE DETECTABLE

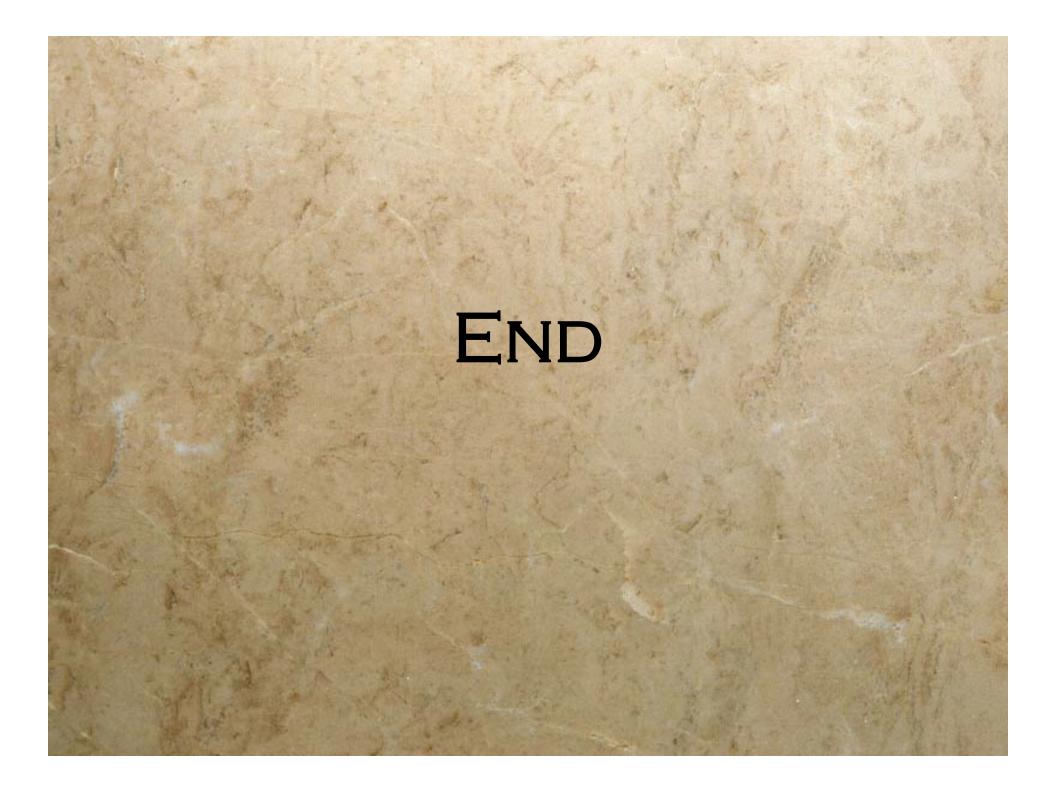
Li & Paczynski (1998) Kulkarni (2005)

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Gravitational Wave Observatories







What do you know? RUMSFELD

"Reports that say that something that has not happened is always interesting to me...

The message is that there are known knowns, there are things we know that we know.

There are known unknowns, that is to say there are things that we now know we don't know.

But there are also unknown unknowns, there are things we do not know we don't know and each year we discover a few more of the unknown unknowns.''

Mr. Donald Rumsfeld, Department of Defense new briefing

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IN SUMMARY: THERE REALLY ARE MANY "UNKNOWN UNKNOWNS" IN THE SKY