

# Are you MKIDing me?!

*The Story of MUSIC's Detectors*



Seth Siegel

# Outline

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- Introduction to Microwave Kinetic Inductance Detectors (MKIDs)
- Recap of the 2010 DemoCam Observing Run
- The problem of dark response
- The problem of low optical efficiency
- Results from the latest 6x6 (4 Color) Array
- Future work

# MUSIC

- **576 spatial pixels**, simultaneously sensitive to **4 spectral bands**

$\nu$ (GHz)	150	230	290	350
$\Delta\nu$ (GHz)	34	45	34	21

- 14 arcminute FOV
- **Primary science goals:**
  - Pointed observations of galaxy clusters through the tSZ Effect
  - Wide blank-field surveys of dusty, star-forming galaxies
- First camera at any wavelength to use MKIDs as detectors
- Scheduled to be commissioned at the CSO in the winter of 2011/2012



*Photo courtesy of Matt Hollister*

# **Introduction to MKIDs**

# Superconductivity

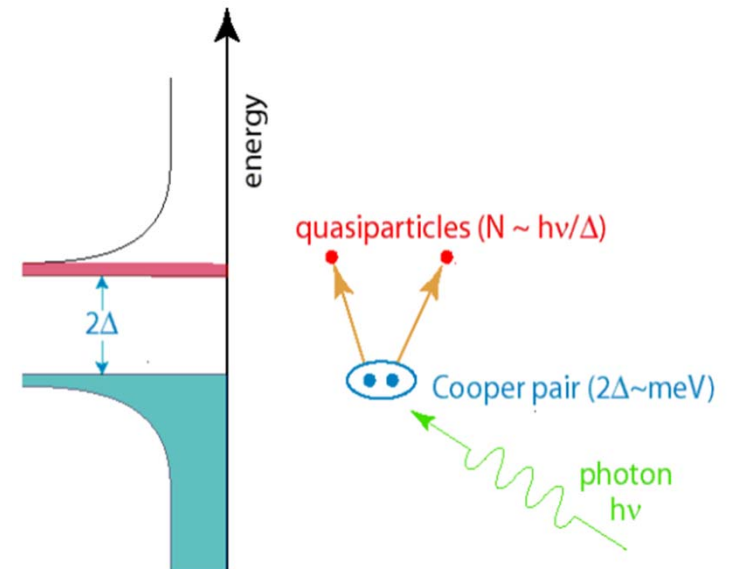
- A Cooper pair consists of two electrons bound together by the electron-phonon interaction. The energy of the Cooper pair is below the Fermi energy.
- Binding energy is weak.

**Aluminum:**  $\Delta \approx 0.18 \text{ meV}$   $\frac{2\Delta}{h} \approx 80 \text{ GHz}$

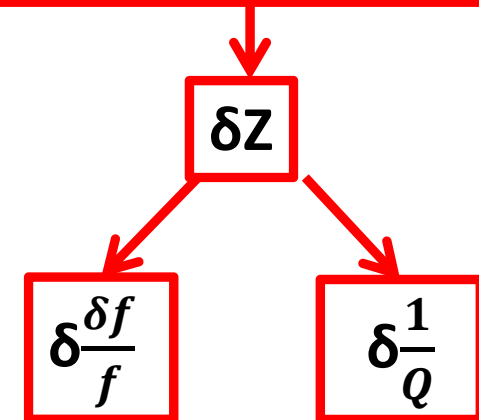
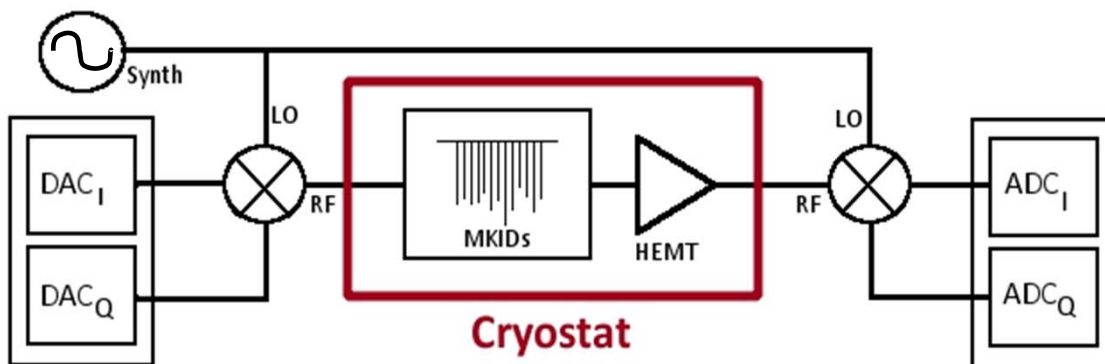
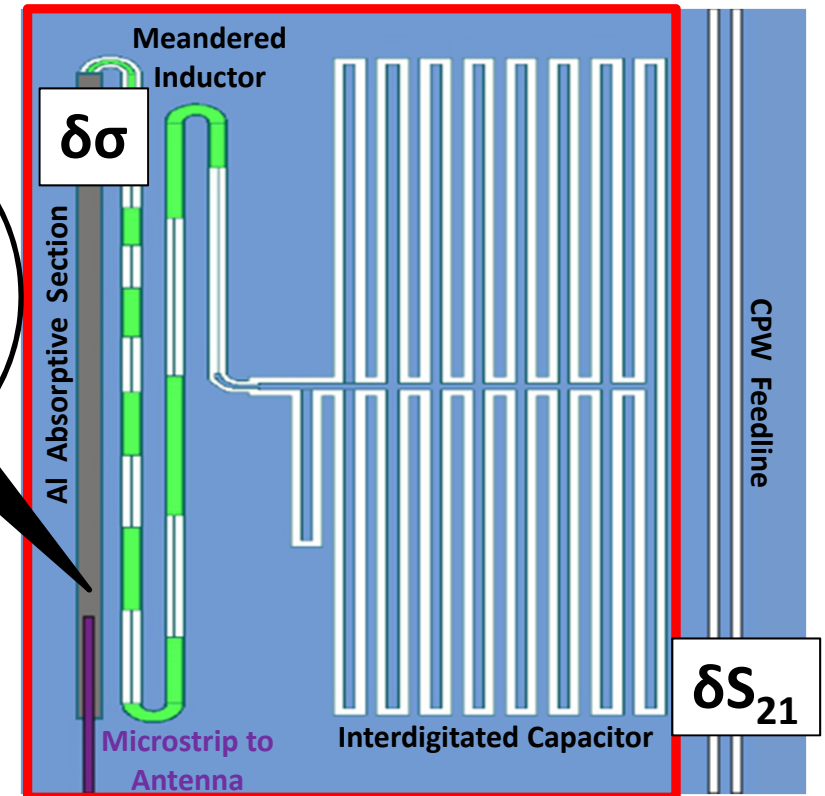
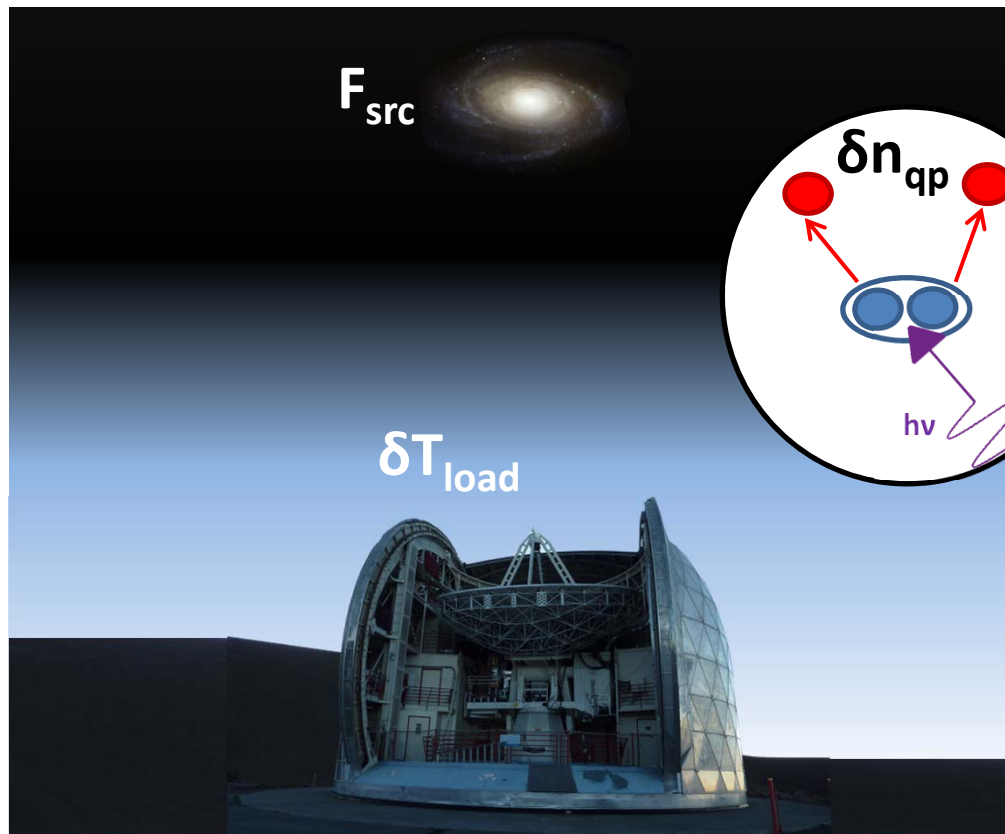
**Niobium:**  $1.45 \text{ meV}$   $700 \text{ GHz}$

- Still, prevents inelastic scattering resulting in zero DC resistance.
- But Cooper pairs have inertia. Results in a “phase-lag” between current and an AC electric field that has a form equivalent to an inductance:

**Kinetic Inductance**



# Principle of Detection



# Basics

- MKIDs are **superconducting RLC circuits**, capacitively coupled to a feedline.
- Near the resonant frequency the complex transmission  $S_{21}$  of a microwave probe signal sweeps out a circle

$$S_{21}(f) = 1 - \frac{Q/Q_c}{1 + 2iQ \frac{f-f_0}{f_0}}$$

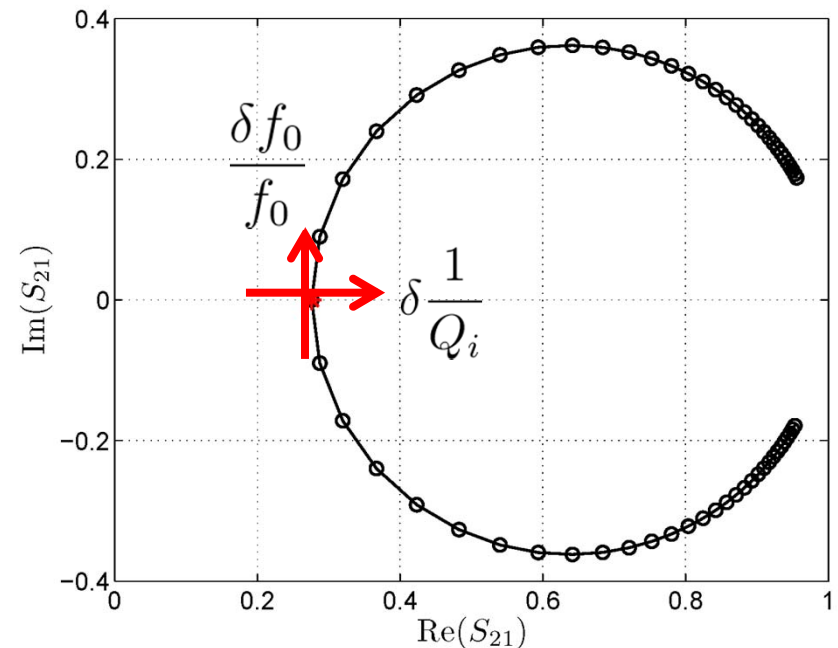
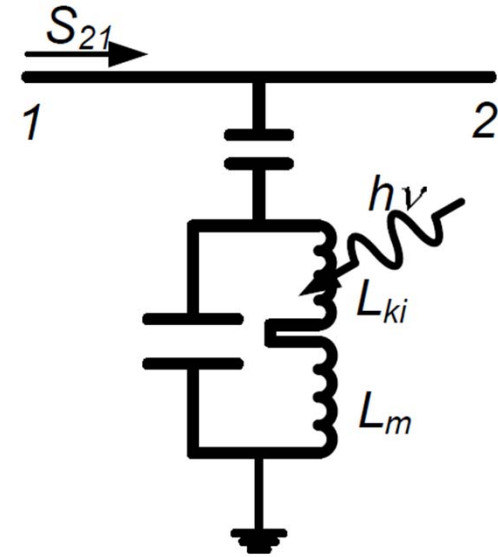
where

$$\frac{1}{Q} = \frac{1}{Q_i} + \frac{1}{Q_c}$$

- Consider small perturbations in frequency and loss, this results in perturbations in the transmission **on resonance** given by

$$\delta S_{21} = \frac{Q^2}{Q_c} \left( \delta \frac{1}{Q_i} - 2i \frac{\delta f_0}{f_0} \right)$$

Amplitude
Phase



# Basics

- The frequency  $f$  and loss  $1/Q$  are determined by the impedance of the circuit

$$Z = R + i\omega L + \frac{1}{i\omega C}$$

Quasi-particles      Cooper Pairs      Geometry  
+ Geometry

- Define the **kinetic inductance fraction**

$$\alpha = \frac{L_k}{L} = \frac{L_k}{L_k + L_m}$$

- Frequency and loss then given by

$$f_0 \propto \frac{1}{\sqrt{LC}}$$

$$\delta f_0 = f_0 \left( \frac{\sqrt{L}}{\sqrt{L + \delta L}} - 1 \right) = -\frac{f_0}{2} \frac{\delta L}{L}$$

$$\boxed{\frac{\delta f_0}{f_0} = -\frac{\alpha}{2} \frac{\delta L_k}{L_k}}$$

$$\frac{1}{Q_i} = \frac{R}{\omega L}$$

$$\boxed{\delta \left( \frac{1}{Q_i} \right) = \alpha \frac{\delta R}{\omega L_k}}$$



# Basics

- Changes in resistance and inductance of the resonator due solely to changes in the complex conductivity of the superconductor

$$\sigma = \sigma_1 - i\sigma_2$$

- For a superconducting thin film

$$\frac{\delta Z_s}{Z_s} = \frac{\delta \sigma}{\sigma}$$

- Lets measure relative to the values at zero temperature, so that

$$\sigma_1(0) = 0 \qquad R(0) = 0$$

and

$$\frac{\delta R}{\omega L_k(0)} = \frac{\delta \sigma_1}{\sigma_2(0)} \qquad \frac{\delta L_k}{\omega L_k(0)} = -\frac{\delta \sigma_2}{\sigma_2(0)}$$

- Now just need a theory for the complex conductivity of a thin superconducting film under an AC electromagnetic field.

# Mattis - Bardeen Theory

- In 1958, D.C. Mattis and J. Bardeen use BCS theory to derive an expression for the complex conductivity

$$\sigma = \sigma_1 - i\sigma_2$$

of a superconducting thin film:

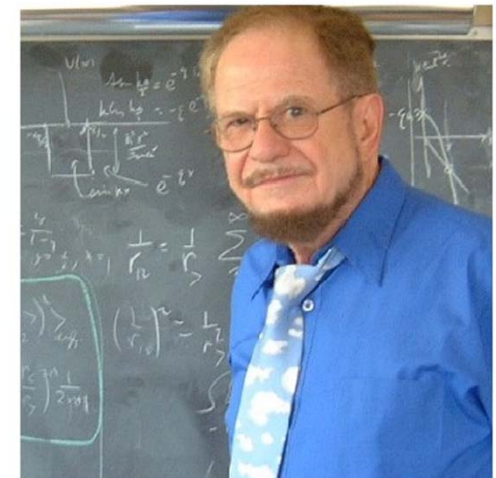
$$\begin{aligned}\frac{\sigma_1}{\sigma_n} &= \frac{2}{\hbar\omega} \int_{\Delta}^{\infty} \frac{[f(E) - f(E + \hbar\omega)](E^2 + \Delta^2 + \hbar\omega E)}{\sqrt{E^2 - \Delta^2} \sqrt{(E + \hbar\omega)^2 - \Delta^2}} dE \\ &\quad + \frac{1}{\hbar\omega} \int_{\Delta - \hbar\omega}^{-\Delta} \frac{[1 - 2f(E + \hbar\omega)](E^2 + \Delta^2 + \hbar\omega E)}{\sqrt{E^2 - \Delta^2} \sqrt{(E + \hbar\omega)^2 - \Delta^2}} dE \\ \frac{\sigma_2}{\sigma_n} &= \frac{1}{\hbar\omega} \int_{\max\{\Delta - \hbar\omega, -\Delta\}}^{\Delta} \frac{[1 - 2f(E + \hbar\omega)](E^2 + \Delta^2 + \hbar\omega E)}{\sqrt{\Delta^2 - E^2} \sqrt{(E + \hbar\omega)^2 - \Delta^2}} dE.\end{aligned}$$

- Here  $f(E)$  is the Fermi distribution

$$f(E; \mu^*, T) = \frac{1}{1 + e^{\frac{E - \mu^*}{kT}}}$$



*John Bardeen*



*Daniel Mattis*

# Mattis – Bardeen Theory

- In the limits applicable to our detectors  $\hbar\omega \ll \Delta$  and  $kT \ll \Delta$  the first order approximation is valid

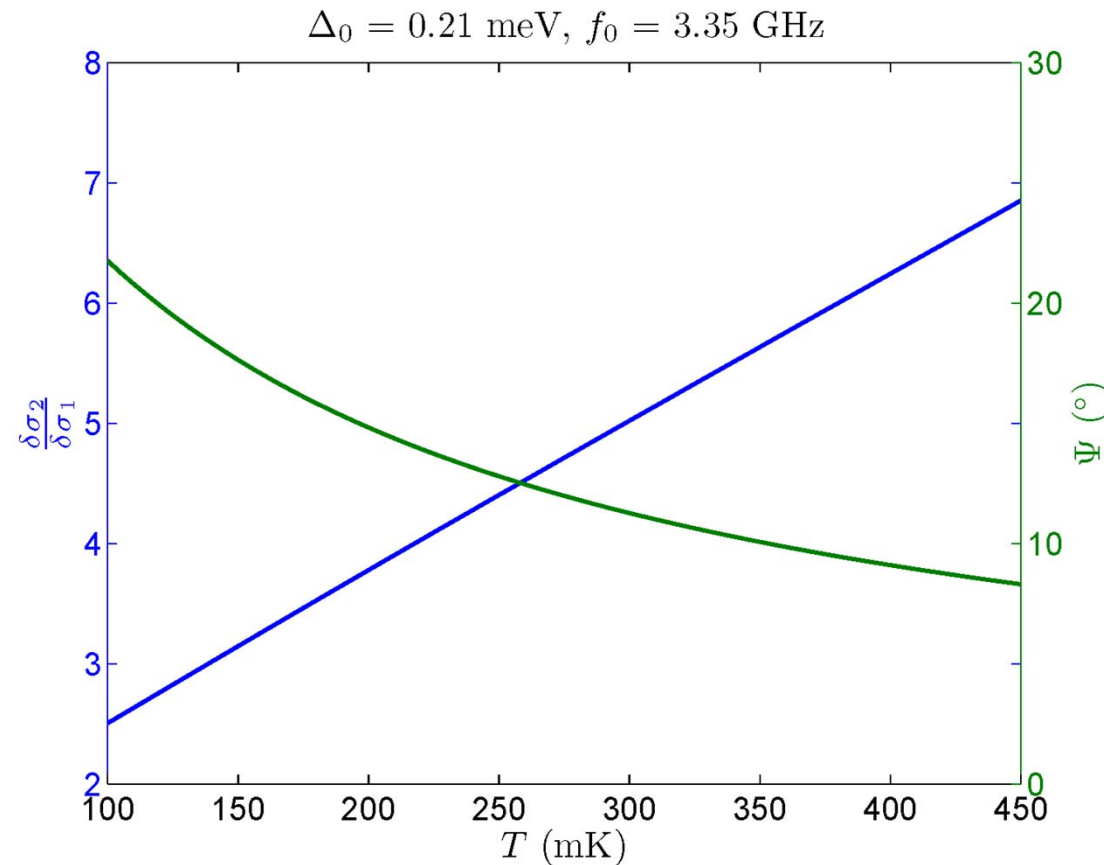
$$\frac{\delta\sigma_1}{\sigma_2(0)} = \frac{1}{\pi N_0} \sqrt{\frac{2}{\pi kT \Delta_0}} \sinh\left(\frac{\hbar\omega}{2kT}\right) K_0\left(\frac{\hbar\omega}{2kT}\right) n_{qp} = \kappa_1(\omega, T, \Delta_0) n_{qp}$$

$$\frac{\delta\sigma_2}{\sigma_2(0)} = -\frac{1}{2\Delta_0 N_0} \left[ 1 + \sqrt{\frac{2\Delta_0}{\pi kT}} e^{-\frac{\hbar\omega}{2kT}} I_0\left(\frac{\hbar\omega}{2kT}\right) \right] n_{qp} = \kappa_2(\omega, T, \Delta_0) n_{qp}$$

- Hence, complex conductivity is proportional to quasi-particle density

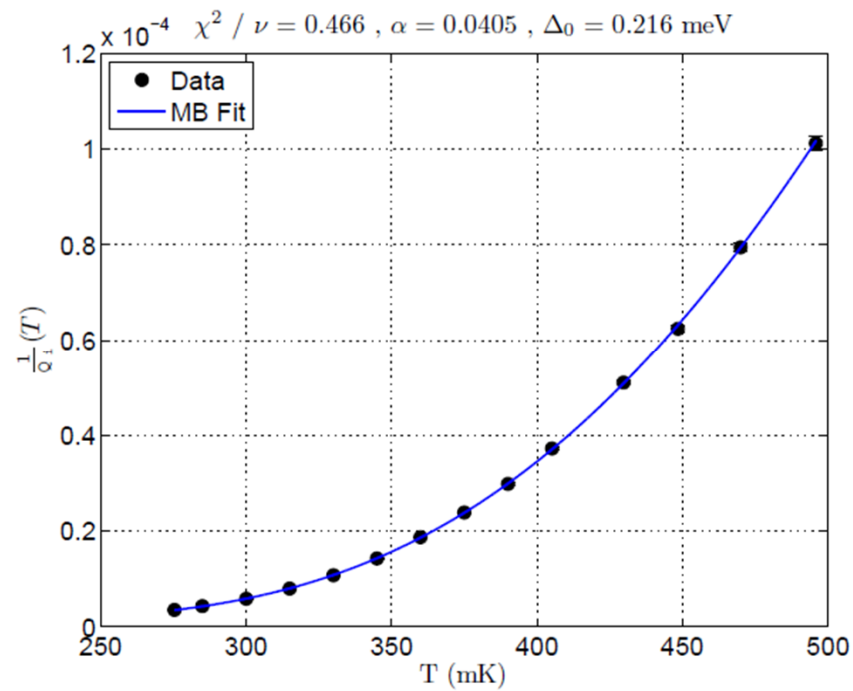
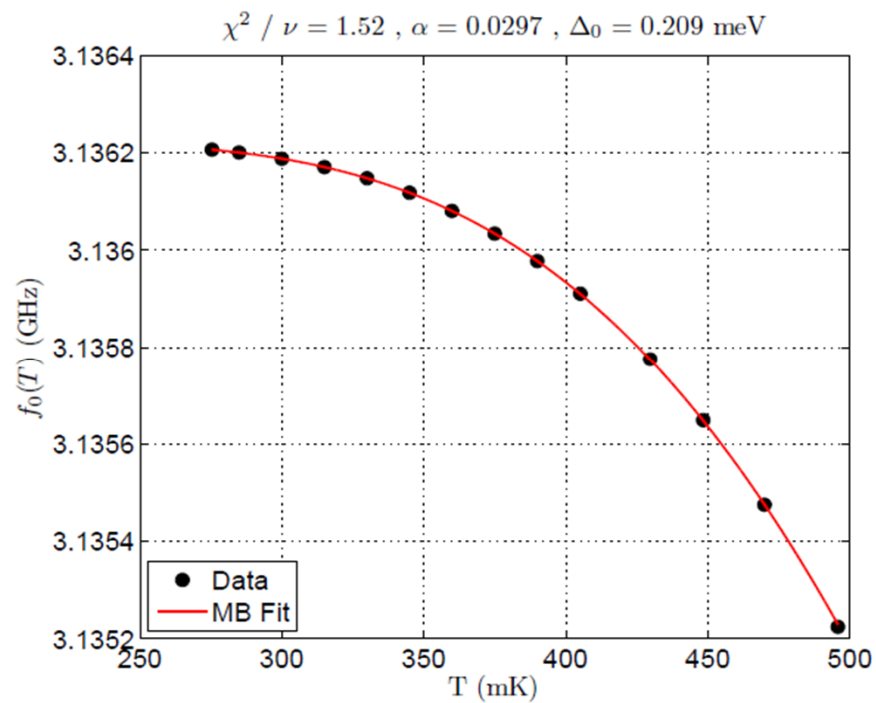
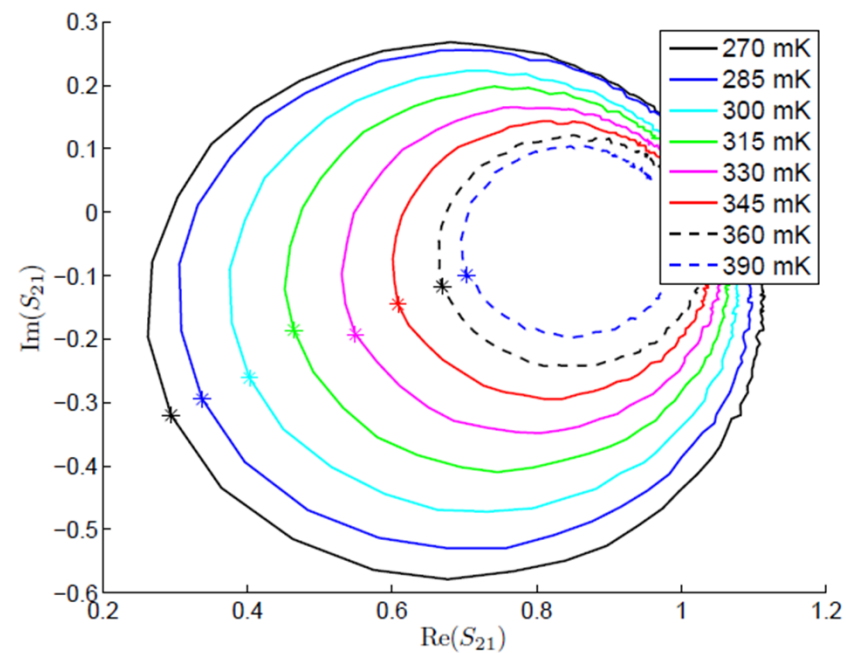
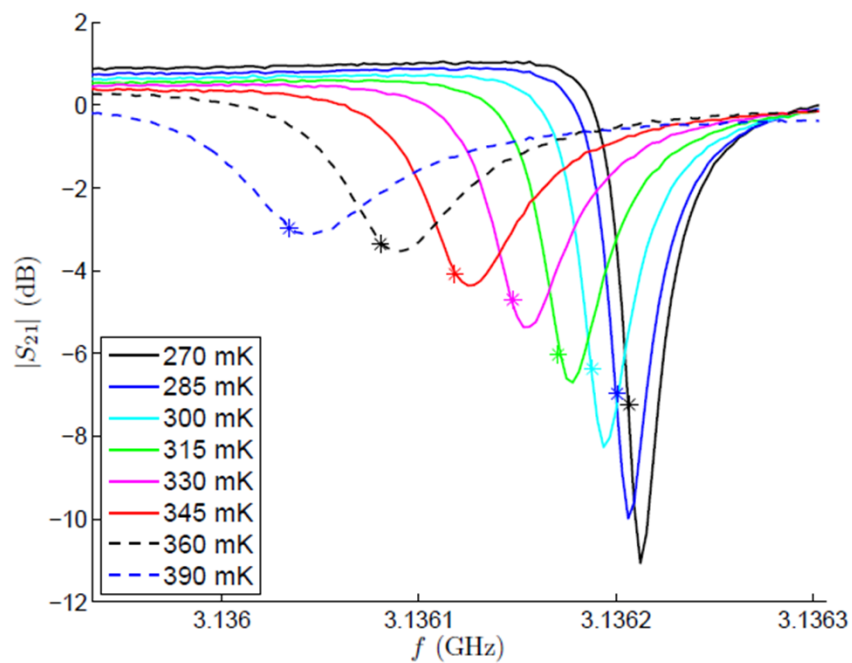
$$n_{qp} = 2N_0 \sqrt{2\pi kT} e^{\frac{\Delta - \mu^*}{kT}}$$

# Mattis – Bardeen Theory



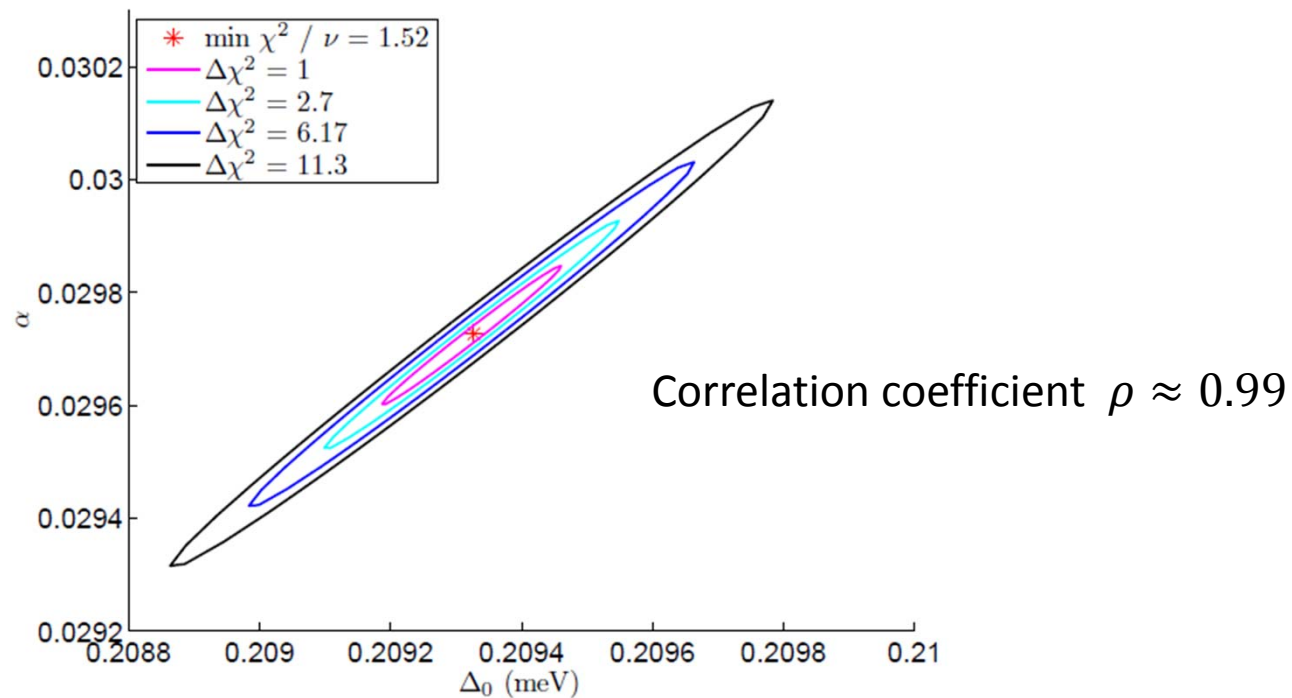
- Frequency response inherently larger than dissipation response

$$\frac{\delta\sigma_1}{\delta\sigma_2} = \frac{\kappa_1(\omega, T, \Delta_0)}{\kappa_2(\omega, T, \Delta_0)} = \tan(\Psi)$$



# Mattis - Bardeen Theory

- Obtain very good estimates of  $\alpha$  and  $\Delta_0$  from MB fits to dark temperature sweeps

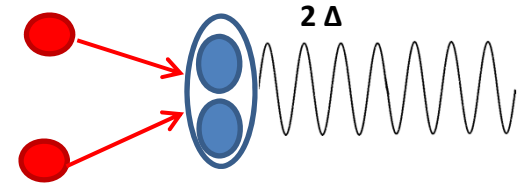
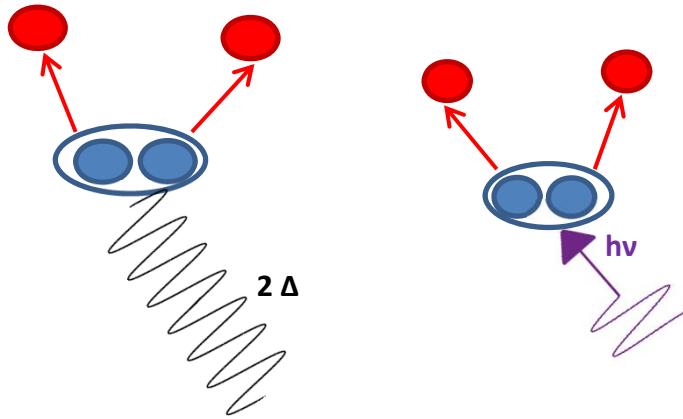


- However, results from fits to frequency and loss inconsistent.

# The G-R Equation

In a steady state

$$\Gamma_{th}^G + \Gamma_{opt}^G = \Gamma^R$$



## Optical Generation

$$\Gamma_{opt}^G = \frac{\eta P}{\Delta} = \frac{\eta k(T_{load} + T_{exc})B}{\Delta}$$

## Recombination

$$\Gamma^R = \frac{V n_{qp}}{\tau}$$

where

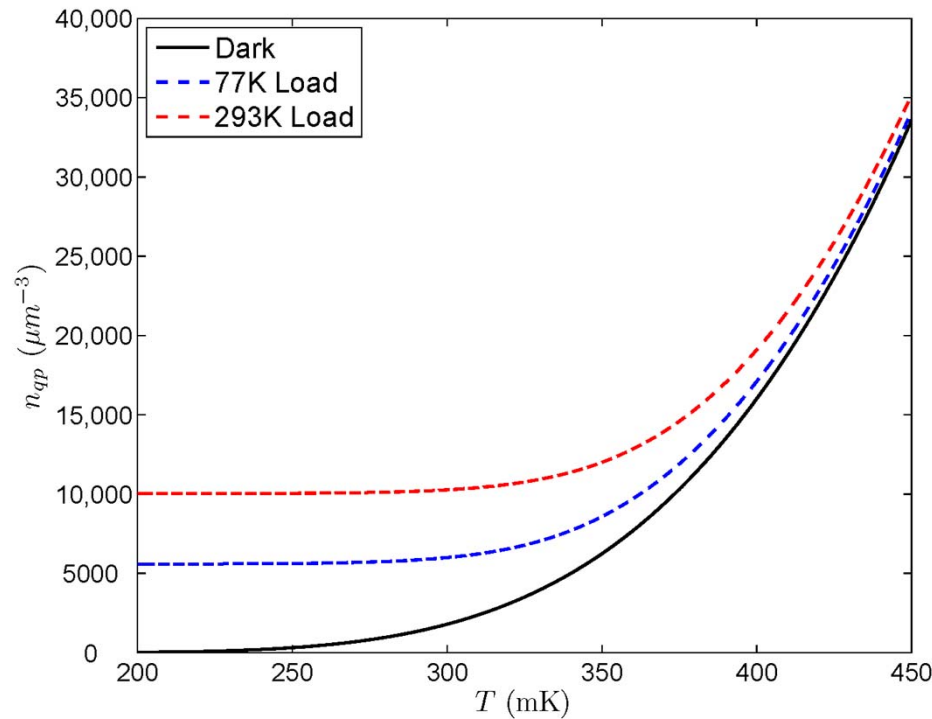
$$\frac{1}{\tau} = R n_{qp} + \frac{1}{\tau_0}$$

$$n_{qp} = n_{qp,opt} + n_{qp,th}(T, \alpha, \Delta_0)$$

$$\Rightarrow \Gamma_{th}^G + \frac{\eta k(T_{load} + T_{exc})B}{\Delta} = R V n_{qp}^2 + \frac{V n_{qp}}{\tau_0}$$

- Obtain the following quadratic equation for quasi-particle density

$$\frac{\eta k (T_{load} + T_{exc}) B}{RV \Delta} = n_{qp}^2 - n_{qp,th}^2 + \frac{1}{R\tau_0} (n_{qp} - n_{qp,th})$$



- Taking the derivative with respect to  $T_{load}$

$$\frac{dn_{qp}}{dT_{load}} = \frac{\eta k B}{RV \Delta \left( 2n_{qp} + \frac{1}{R\tau_0} \right)}$$



# Responsivity

$$\text{Responsivity} = \left[ \frac{dS_{21}}{d(\frac{\delta f}{f})} \right] \left[ \frac{d(\frac{\delta f}{f})}{d(\frac{\delta \sigma_2}{\sigma_2})} \right] \left[ \frac{d(\frac{\delta \sigma_2}{\sigma_2})}{dn_{qp}} \right] \left[ \frac{dn_{qp}}{dT_{load}} \right] \left[ \frac{dT_{load}}{dF_{src}} \right]$$

$$\frac{dS_{21}}{d(\frac{\delta f}{f})} = \frac{2Q^2}{Q_c}$$

$$\frac{d(\frac{\delta f}{f})}{d(\frac{\delta \sigma_2}{\sigma_2})} = \frac{\alpha}{2}$$

$$\frac{d(\frac{\delta \sigma_2}{\sigma_2})}{dn_{qp}} = \frac{1}{2N_0\Delta_0} \left[ 1 + \sqrt{\frac{2\Delta_0}{\pi kT}} e^{-\frac{\hbar\omega}{2kT}} I_0 \left( \frac{\hbar\omega}{2kT} \right) \right] = \kappa_2(\omega, T, \Delta_0)$$

$$\frac{dn_{qp}}{dT_{load}} = \frac{\eta k B}{RV \Delta_0 (2n_{qp} + \frac{1}{R\tau_0})}$$

$$2kT_{src}\Delta\nu = F_{src}A_{tel}\eta_{tel}\Delta\nu \quad \Rightarrow \quad \frac{dT_{load}}{dF_{src}} = \frac{\eta_{tel}A_{tel}}{2k}$$

## What makes for a responsive MKID?

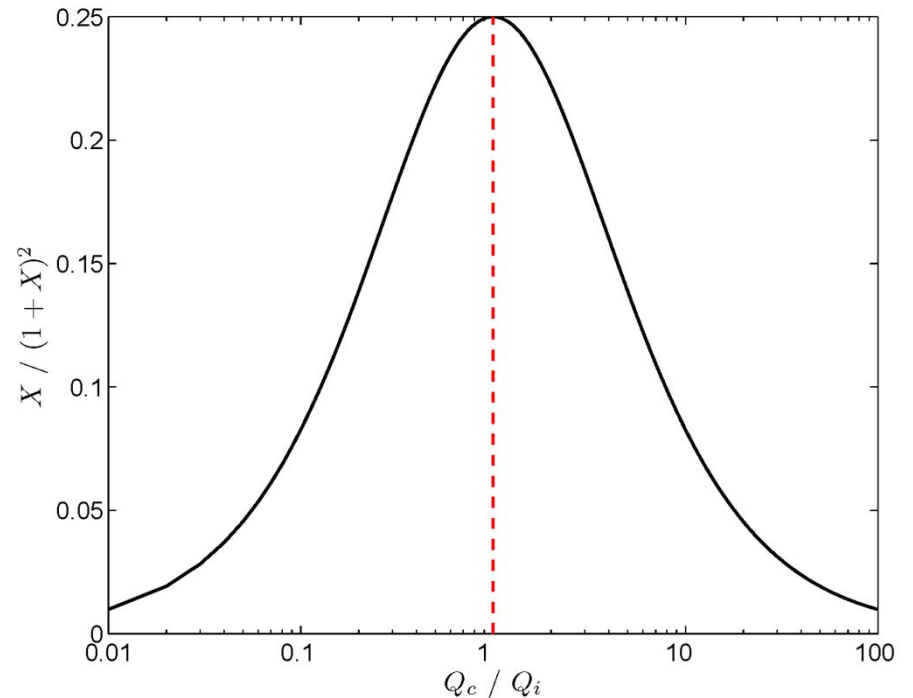
$$\text{Responsivity} = \frac{\kappa_2(\omega, T, \Delta_0)}{\Delta_0} \left( \frac{\eta_{tel} A_{tel}}{2} \right) \left( \frac{\alpha}{V} \frac{\eta B}{(2Rn_{qp} + \frac{1}{\tau_0})} \right) \left( \frac{Q^2}{Q_c} \right)$$

Under typical loading conditions and bath temperatures  $Rn_{qp} \gg \frac{1}{\tau_0}$

Neglecting the factors intrinsic to the material we are using (i.e.  $R, N_0, \Delta_0$ ) we obtain the following simplified expression:

$$\text{Responsivity} \propto \frac{\eta B \alpha^2 Q_i^2}{V} \frac{X}{(1 + X)^2}$$

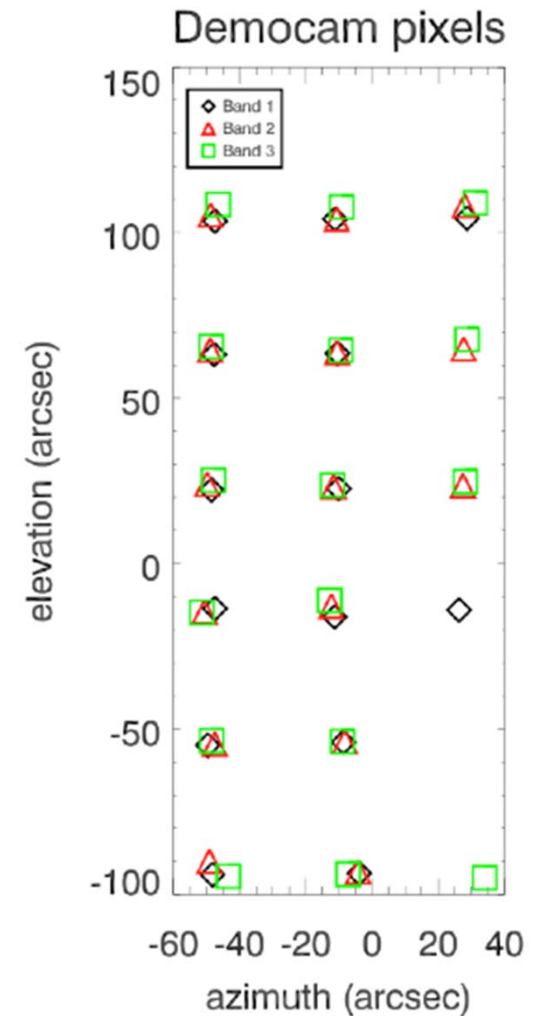
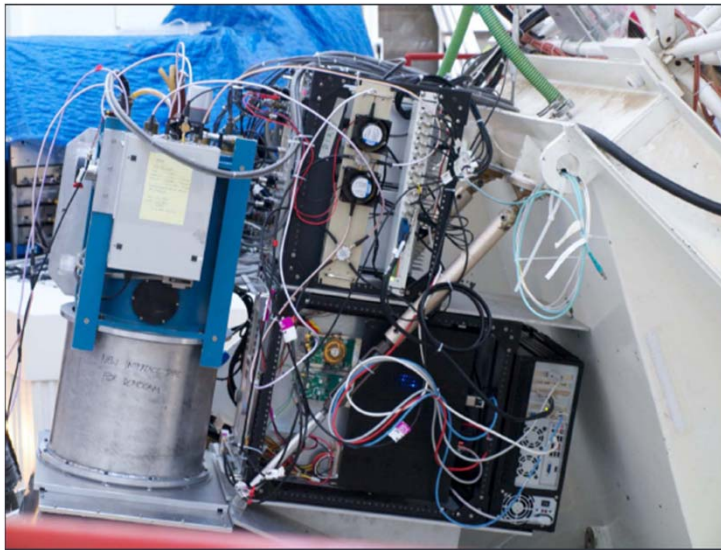
$$\text{where } X = \frac{Q_c}{Q_i}$$



**One year ago...**

# 2010 Observing Run

- Beam shapes and locations matched expectations
- Successfully read out **100+ carriers (60 resonators)**
- Demonstrated a path to the full-scale camera
- Plagued by low-frequency electronics noise
- Sensitivity 200-500 mJy  $s^{1/2}$
- Factor of 10-100 above BLIP



*Credit: Jack Sayers*

# The Problem of Dark Response

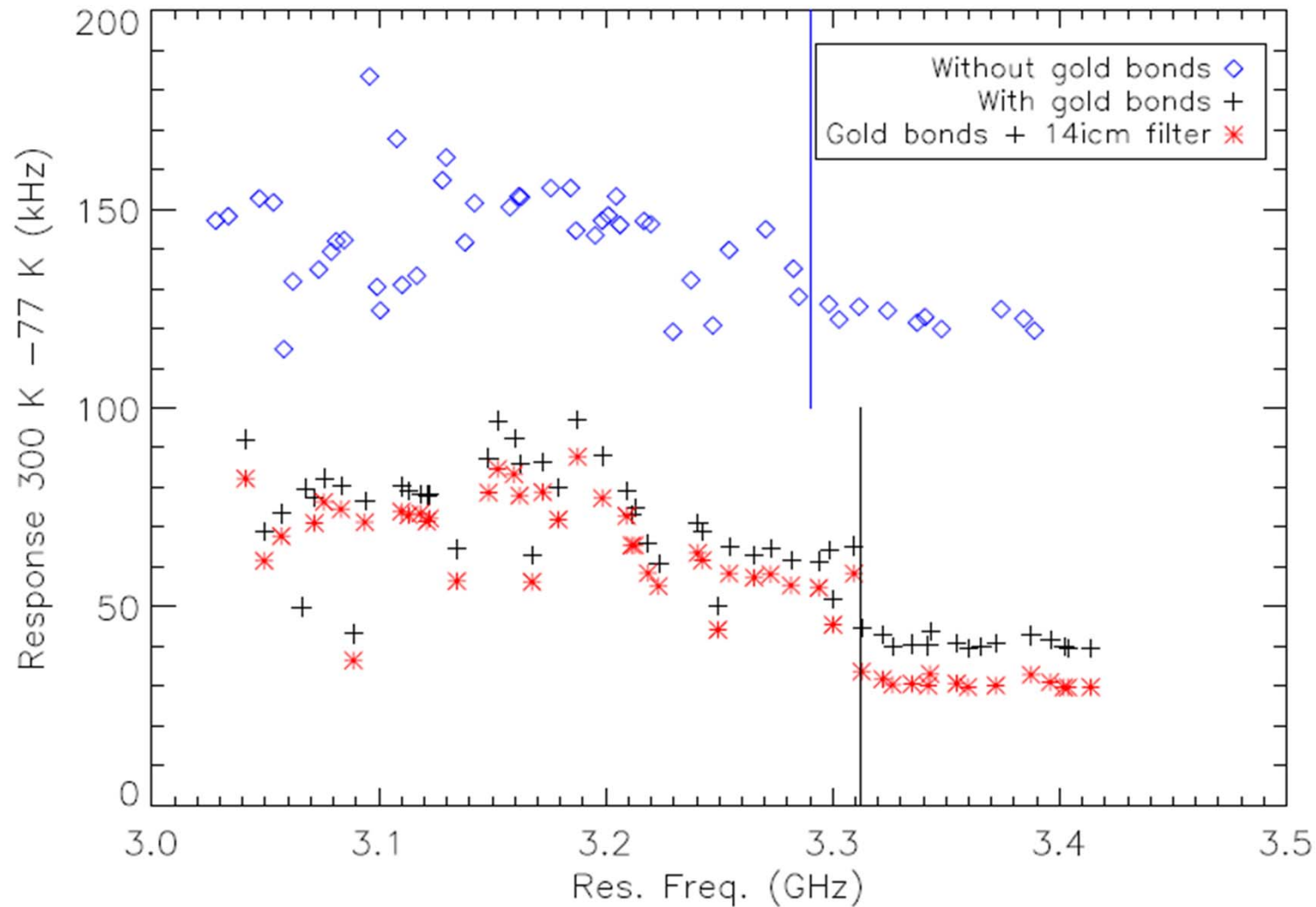
*Resonators not-coupled to an antenna show significant response to outside-the-dewar loading.*

## Possible causes?

- **Heating of the substrate**
- **Direct pick-up**
  - Direct absorption in the aluminum section
  - Coupling through the interdigitated capacitor (IDC)

# Substrate Heating


Result of adding gold wire bonds and 14 icm (420 GHz low-pass) filter

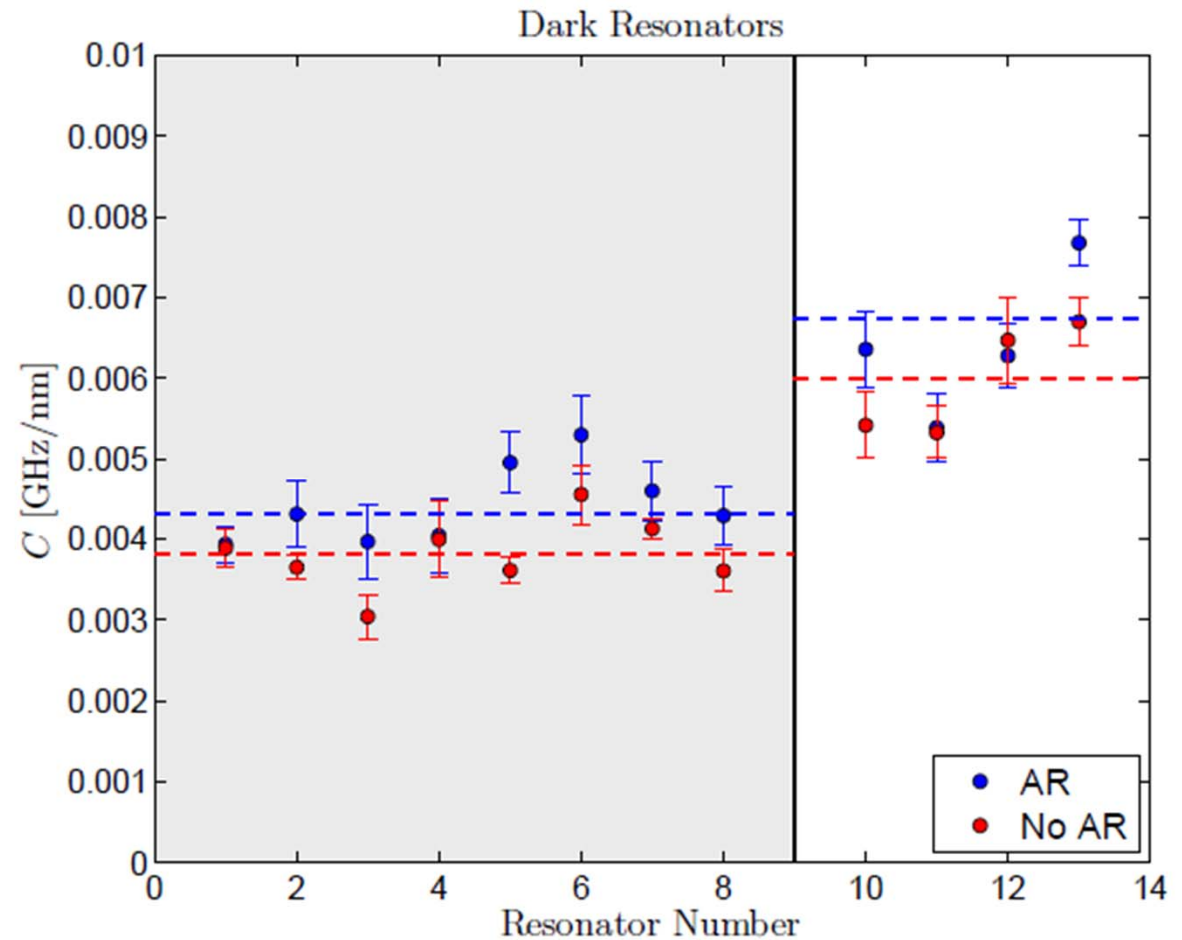


Credit: James Schlaerth

# Half - Aperture Test

Placed aluminum aperture over half of the array.  
Measured efficiency of the non-antenna coupled resonators.

$$C = \frac{\eta B}{d}$$




# ε Test Devices

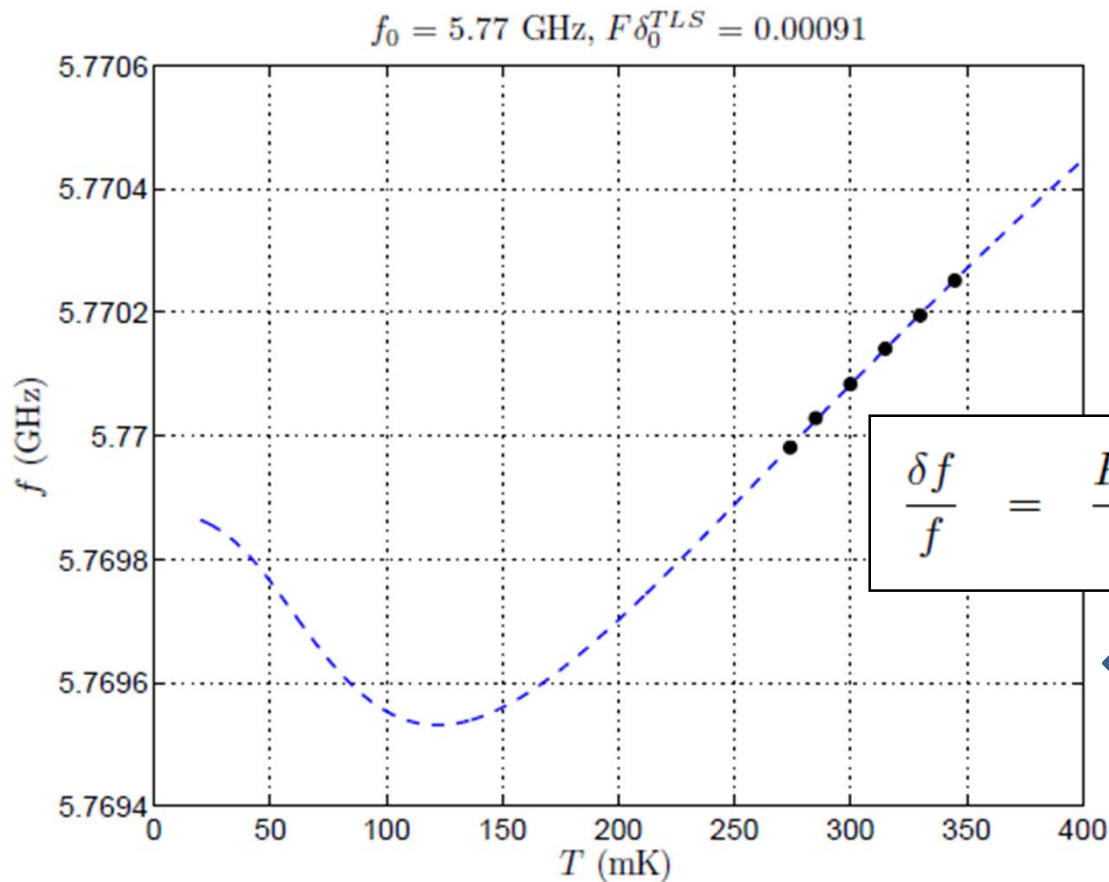
All Niobium  $\Rightarrow T_c = 9.2 \text{ K}$

$$\frac{2\Delta}{h} \approx 700 \text{ GHz}$$

- Cannot absorb mm –radiation (no direct pick-up)
- Thermal quasi-particle density negligible

Resonator frequency  
determined by TLS effects

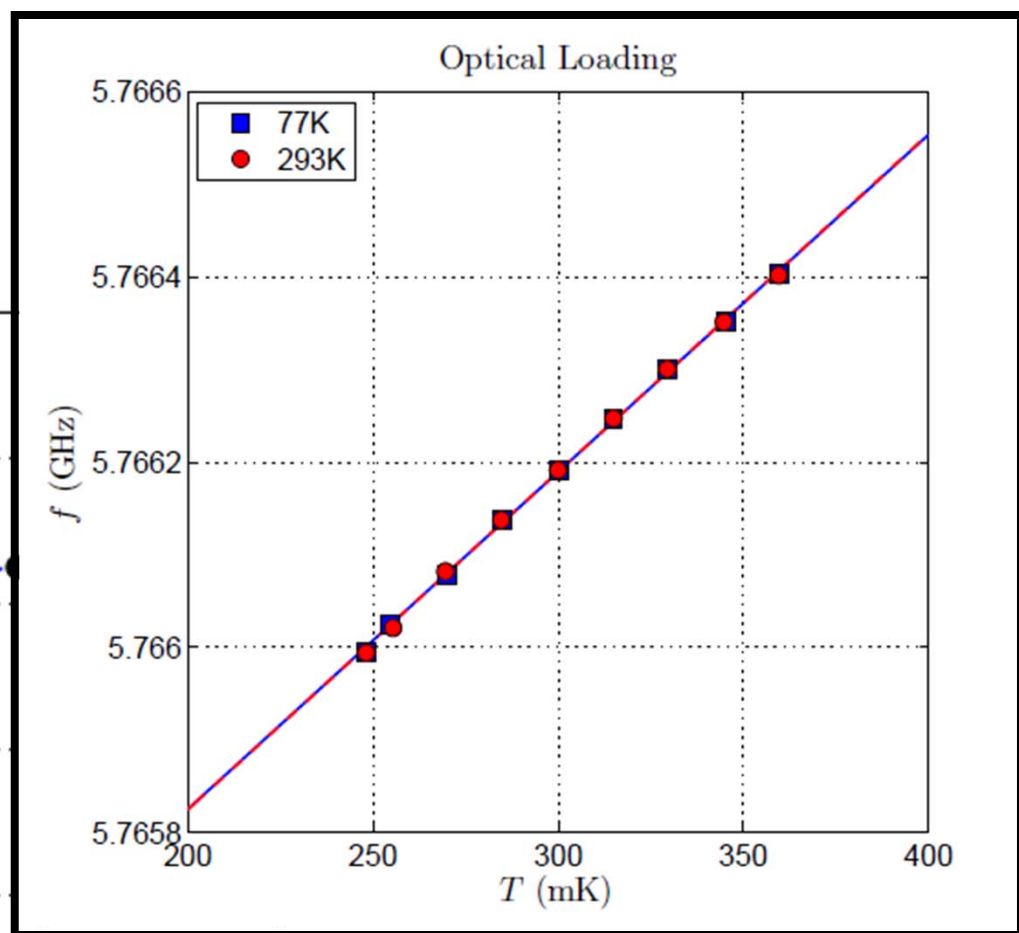
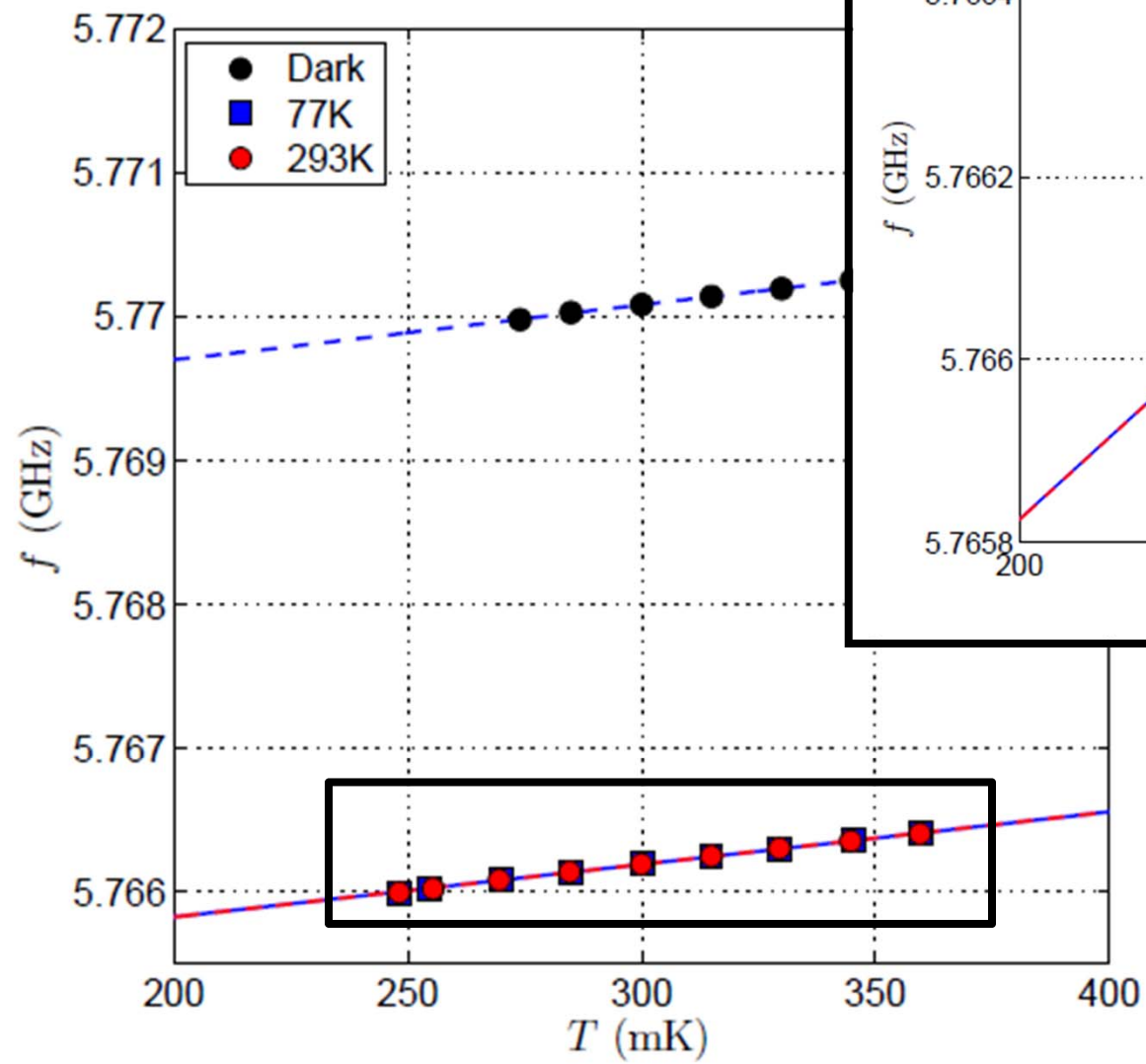
Use as a **TLS thermometer**



$$\frac{\delta f}{f} = \frac{F\delta_0^{TLS}}{\pi} \left[ \text{Re}\Psi \left( \frac{1}{2} - \frac{\hbar f_0}{ikT} \right) - \log \frac{\hbar f_0}{kT} \right]$$

$\leftarrow$  Dark calibration





# ε Test Devices

- No change in the resonant frequency between 77K and 293K loads. Places tight upper-limit on the change in substrate temperature:

$$T_{293K} - T_{77K} \leq 1.0 \text{ mK}$$

- The large response in our normal resonators between hot/cold loads simply cannot be explained by such a small change in substrate temperature.

 **It's Direct Pick-Up**

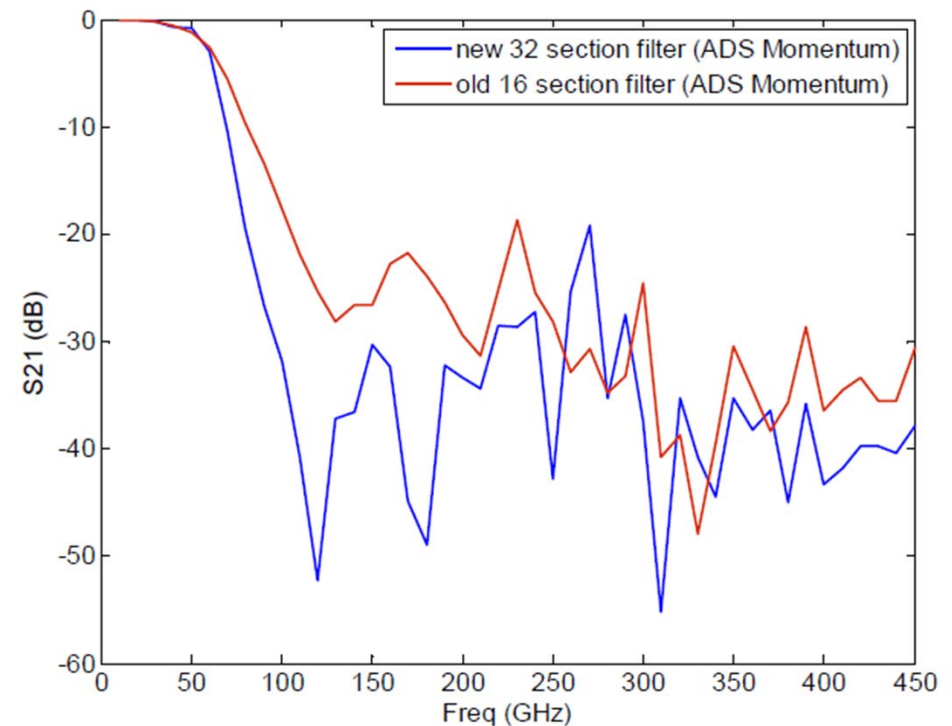
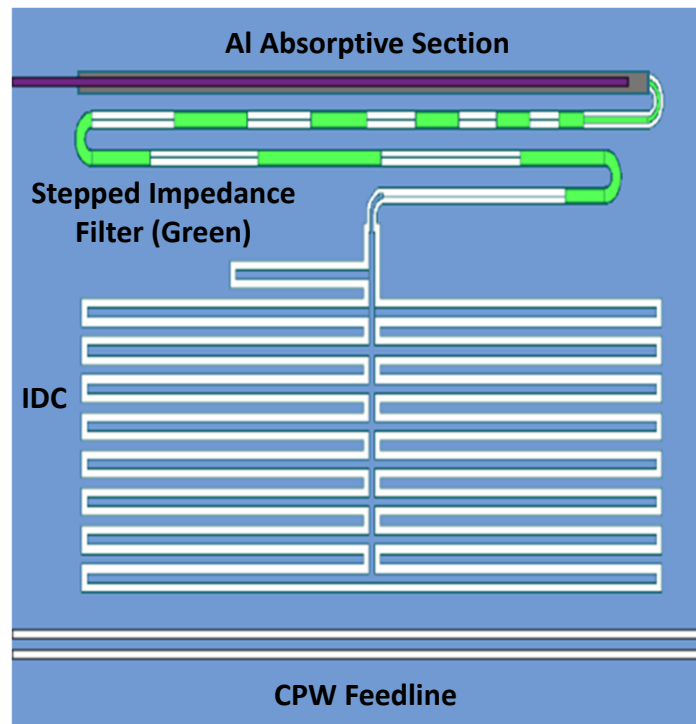
- Epsilon test devices also provide us with an estimate of the loss tangent of the dielectric

$$F\delta_0^{TLS} = 0.00085$$

- And an estimate of the dielectric constant after simulating the magnetic inductance  $L_m$ .

# Coupling through the IDC

- Originally implemented IDC to reduce TLS noise
- Observed large increase in direct pick-up. Due to the geometry of the capacitor coupling to broadband radiation.
- Stepped impedance filter introduced between IDC and AI section to block frequencies  **$80 \text{ GHz} < \nu < 700 \text{ GHz}$**



Credit: Omid Noroozian

# Direct Pick-Up Device

- Currently testing a device without antennas
- Contains three groups of five resonators:
  1. IDC with 16-element stepped impedance filter
  2. IDC with **new** 32-element stepped impedance filter
  3. CPW
- In each group the Al absorptive section is varied in length between 0 mm and 1 mm with the remainder Nb.
- If pick-up is through the aluminum, the response should have some dependence on Al length.
- If there is still pick-up through the IDC we should see drastic differences in the response between the three groups with **CPW < IDC-32 < IDC-16**

# The Problem of Low Optical Efficiency

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*Not so much a problem as an insufficient understanding of the factors that effect the overall absorption of power in our detectors.*

To this end we pursued:

1. A systematic characterization of all sources of loss in our system.
2. Extensive hot/cold measurements, an **improved model** of our detectors, and a **robust fitting procedure** to obtain better estimates of the actual system efficiency.

# Expected Efficiency

## DemoCam

Band	Filters	Lyot Stop	Antenna Efficiency	Fringing Efficiency	Reflection at MKID	Phonon Emission	Overall Efficiency
0	0.73	0.18	0.42	0.69	1.0	0.66	<b>0.025</b> +/- 0.005
1	0.76	0.36	0.47	0.68	1.0	0.58	<b>0.05</b> +/- 0.01
2	0.73	0.52	0.49	0.76	1.0	0.58	<b>0.08</b> +/- 0.02
3	0.68	0.63	0.41	0.96	1.0	0.58	<b>0.10</b> +/- 0.02

## MUSIC

Band	Filters	Lyot Stop	Antenna Efficiency	Fringing Efficiency	Reflection at MKID	Phonon Emission	Overall Efficiency
0	0.73	0.28	0.65	0.69	1.0	0.66	<b>0.06</b> +/- 0.01
1	0.74	0.51	0.58	0.68	1.0	0.58	<b>0.09</b> +/- 0.01
2	0.67	0.68	0.61	0.76	1.0	0.58	<b>0.12</b> +/- 0.01
3	0.58	0.76	0.54	0.96	1.0	0.58	<b>0.13</b> +/- 0.01

# Full Model

## Data

Hot /cold measurements of resonator frequency and quality factor made over a wide range of base temperatures (200 – 400 mK)

## Model

Mattis-Bardeen Theory + Generation-Recombination Equation

## Fit To

Frequency Shift

$$\Delta f(T) = f(T, 293\text{K}) - f(T, 77\text{K})$$

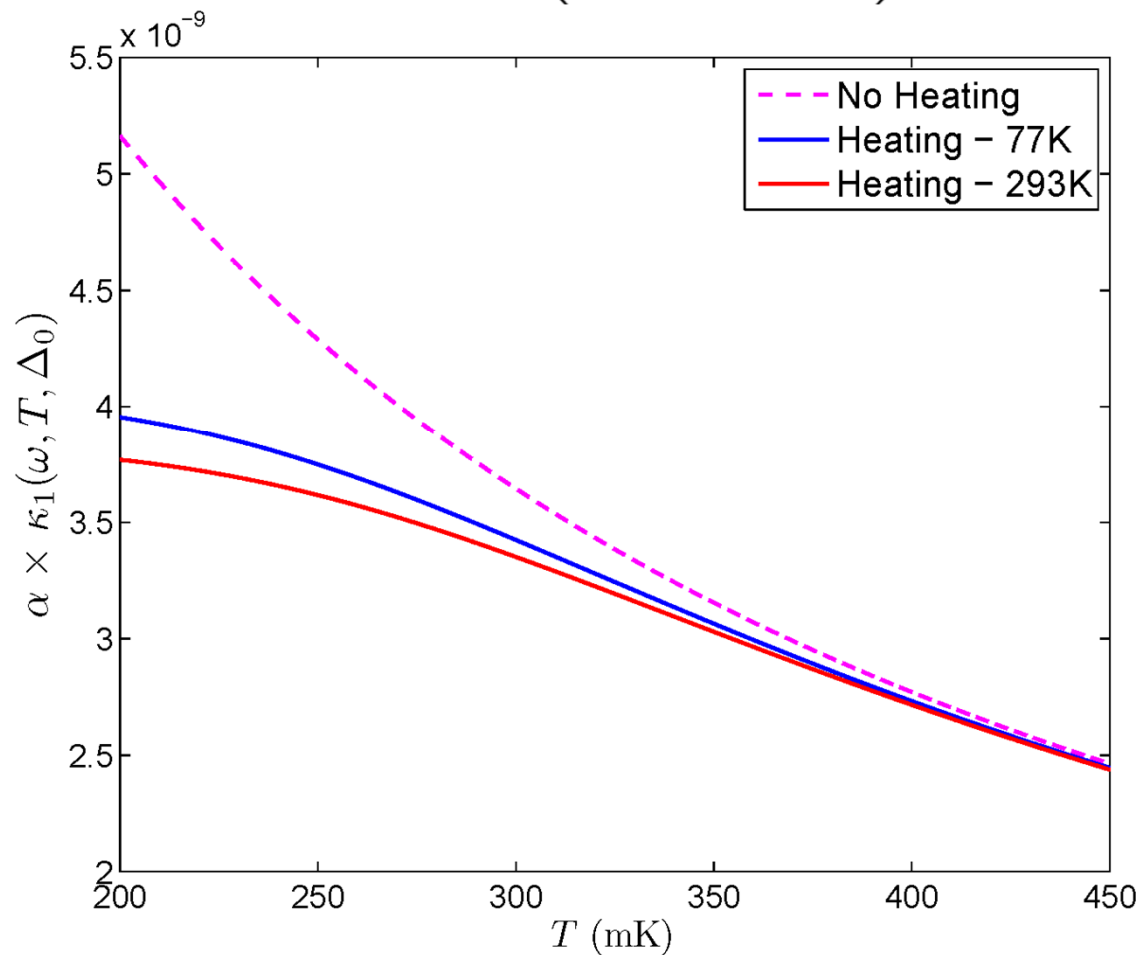
and Naïve Excess Load

$$T_{exc}^{naive} = \frac{293/Q_i^2(T, 77\text{K}) - 77/Q_i^2(T, 293\text{K})}{1/Q_i^2(T, 293\text{K}) - 1/Q_i^2(T, 77\text{K})}$$

# Full Model

- Allow for an elevated “effective temperature”  $T_{\text{eff}}$ , parameterized by a conductance law of the form

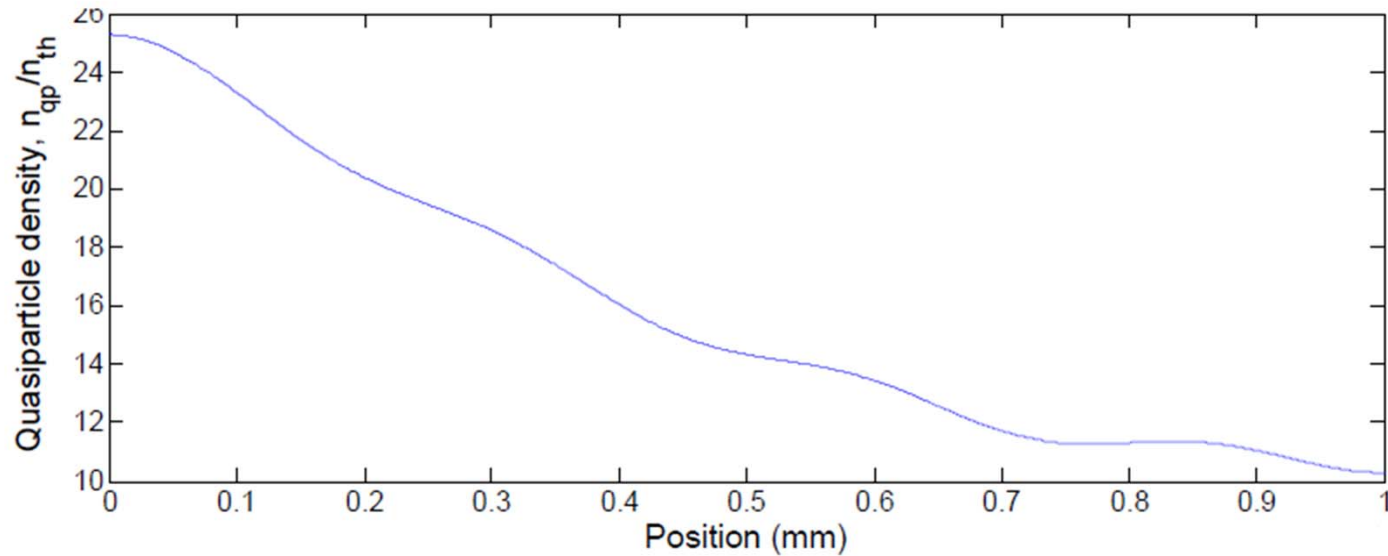
$$P = g \left( T_{\text{eff}}^{\beta} - T^{\beta} \right)$$





# Full Model

- **Do not** take into account non-uniform absorption of radiation

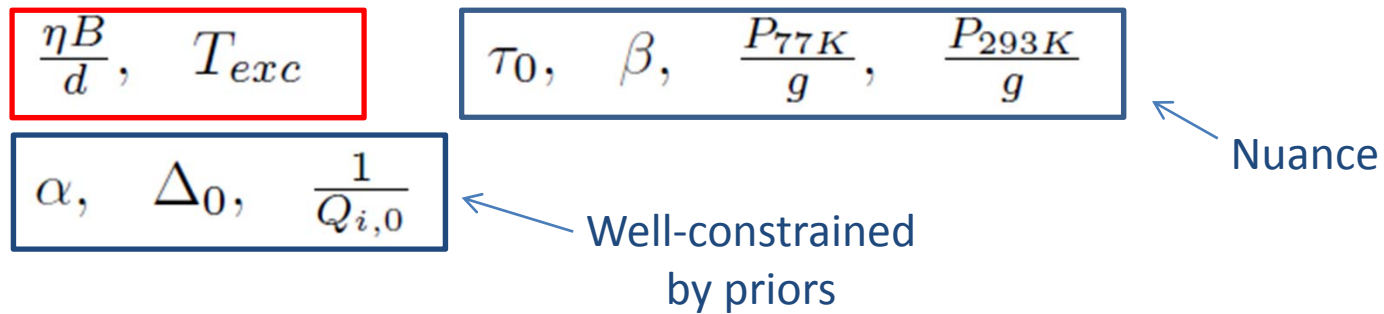


*Credit: Jonas Zmuidzinas*

- Instead include a **1.15** correction factor to the efficiency obtained from the fit

# Fitting Procedure

- Total of 9 parameters, 3 have prior information from dark Mattis-Bardeen fits



- Implement a MCMC using the **Metropolis Algorithm** and **simulated annealing** to efficiently explore the parameter space and search for global minimum of the likelihood function
  - Accounts for uncertainty in bath temperature, loading conditions, etc.
  - Correctly capture the large correlation between  $\alpha$  and  $\Delta_0$

# **Most Recent Results**

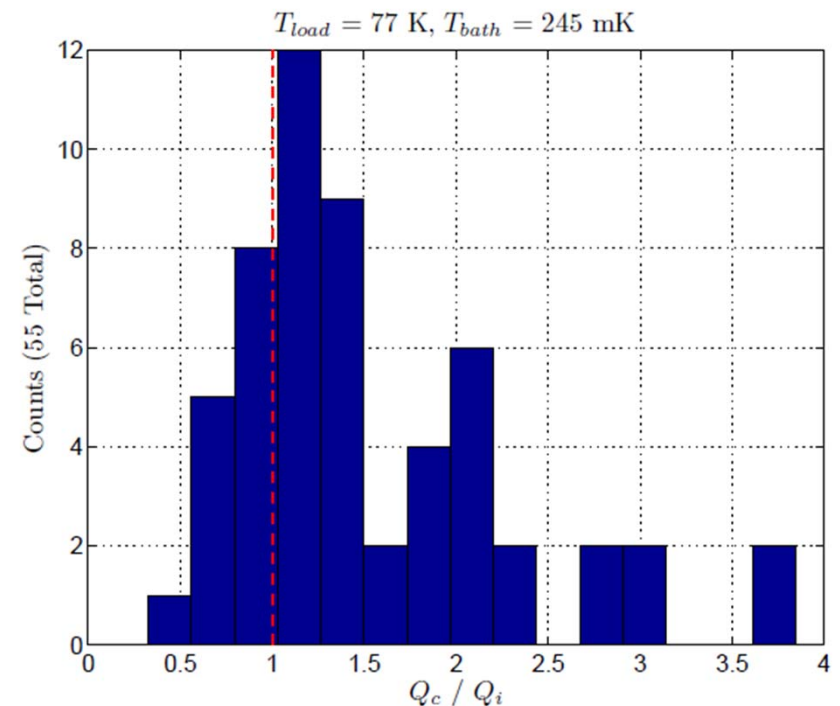
# Array Description

## Changes Made

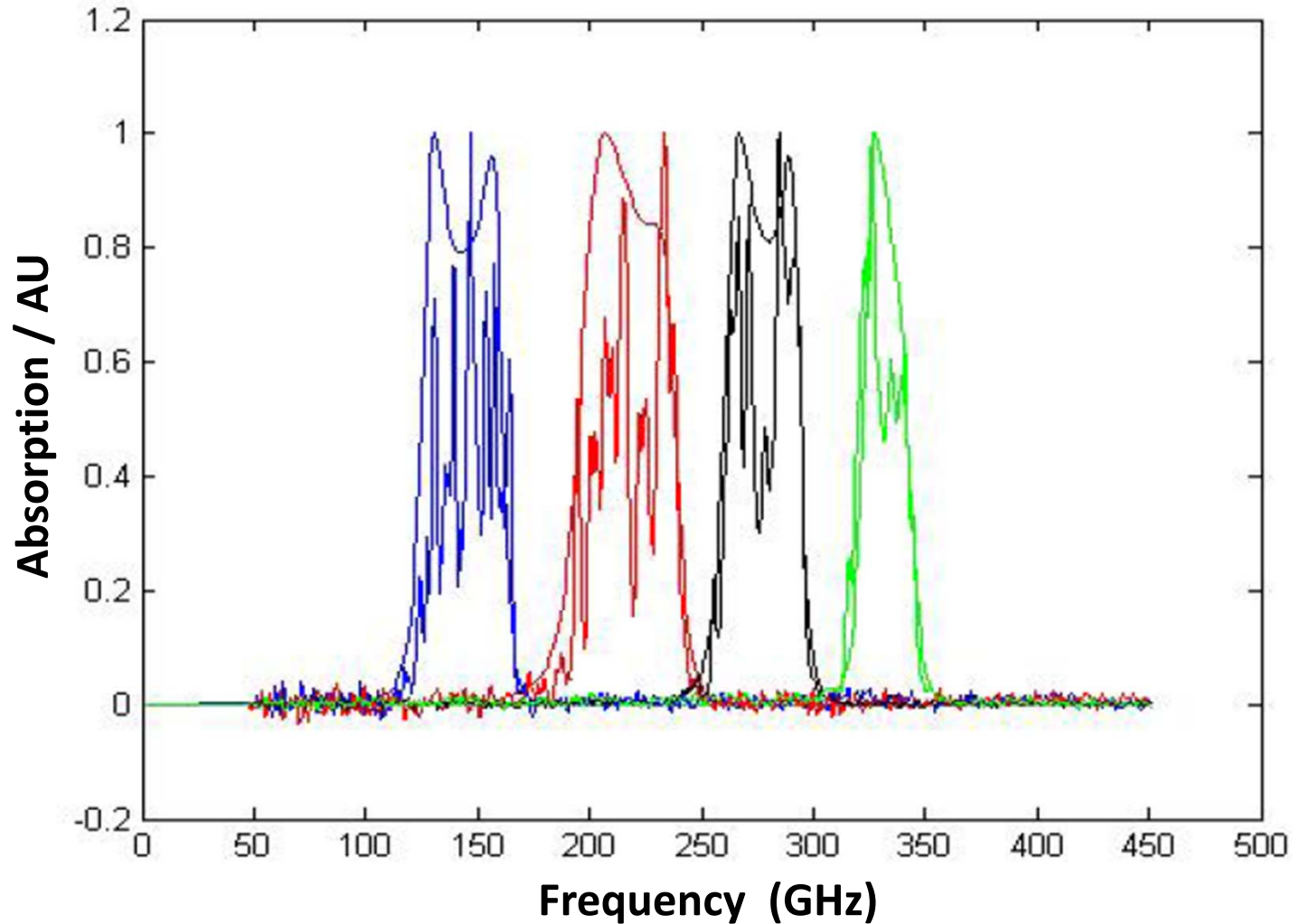
- Switched dielectric in microstrip and bandpass filters from  $\text{SiO}_2$  to  $\text{Si}_3\text{N}_4$
- Increased antenna slot and tap impedance for improved efficiency in lower frequency bands
- Included “index” and “loss” test devices to measure the phase velocity and loss in the microstrip
- Added a **fourth band at 150 GHz (Band 0)**

## Broad Characteristics

- Approximately 100 out of 144 resonators survived (**70% yield**)
- Internal and coupling Q's well matched at base temperature under 77K load (comparable to sky loading at CSO).



# Bandpasses



Epsilon of  $\text{Si}_3\text{N}_4$  : 7.0

*Credit: Ran Duan*

# Optical Efficiency

Band	Expected Efficiency	Measured Efficiency
0	0.025 +/- 0.005	0.055 +/- 0.002
1	0.05 +/- 0.01	0.082 +/- 0.003
2	0.08 +/- 0.02	0.073 +/- 0.001
3	0.10 +/- 0.02	0.075 +/- 0.004

Note: Quoted uncertainties are only statistical. **Actually sensitive to**

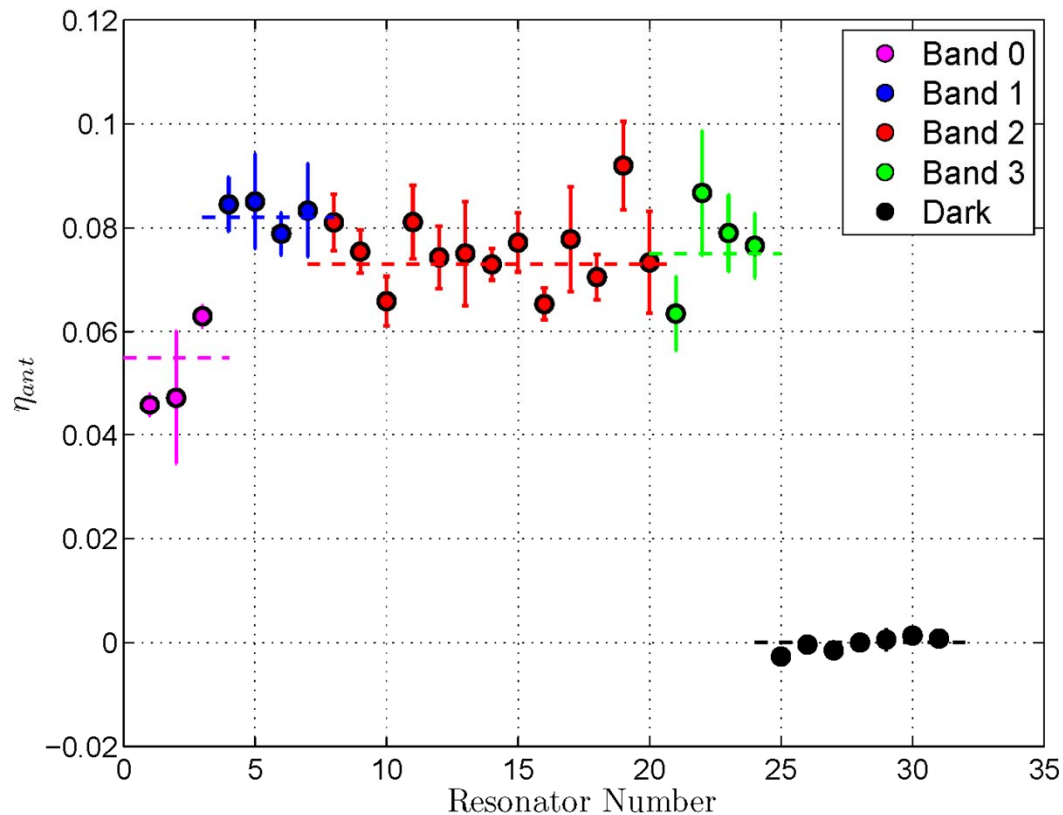
$$\frac{\eta}{RV}$$

Assume that

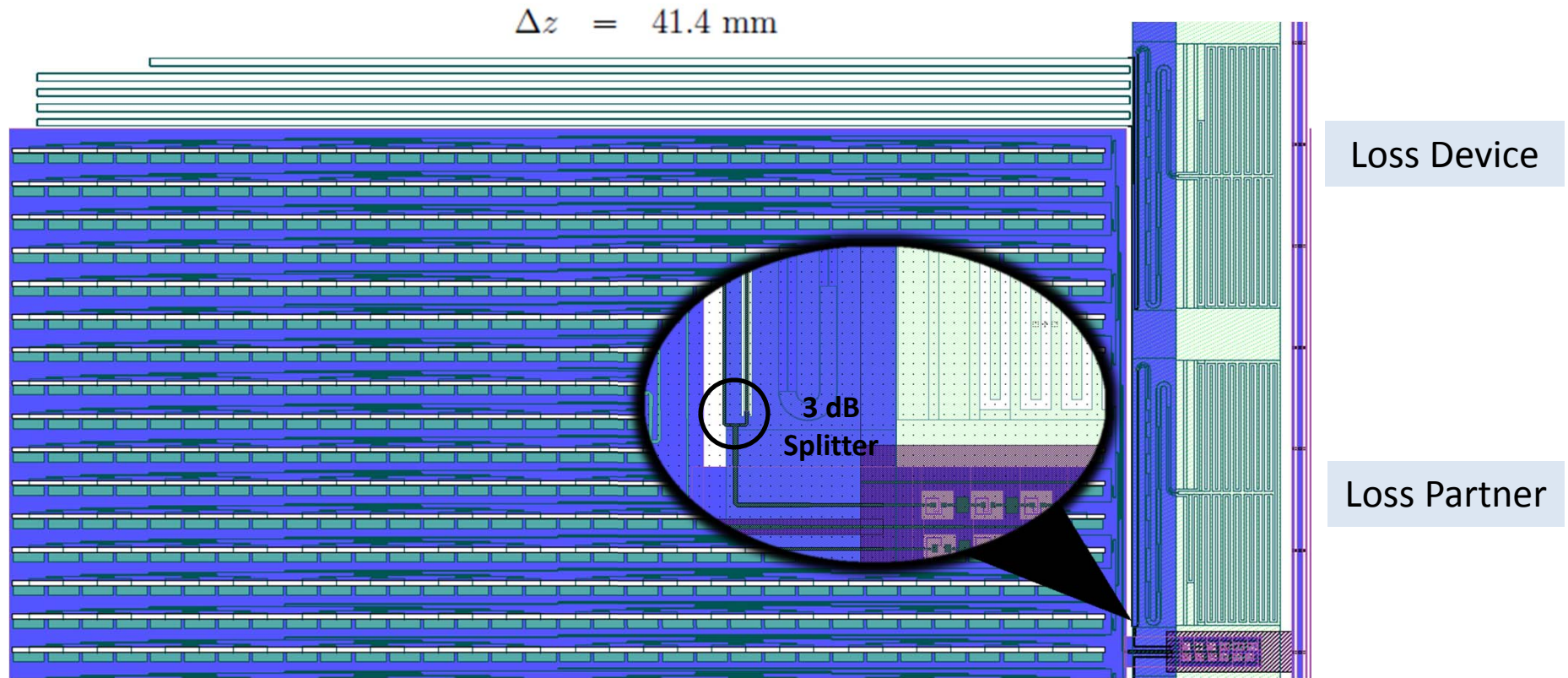
$$R = 9.4 \mu\text{m}^3 \text{ s}^{-1}$$

and

$$V = 1000 \mu\text{m} \times 6 \mu\text{m} \times (55 \text{ nm})$$



# Loss Test Devices



Latest mask contained devices designed to measure the loss tangent of  $\text{Si}_3\text{N}_4$ . These are based on devices designed by Chao-Lin Kuo for CMB polarization experiments, scaled to higher frequencies.

Loss Partner -- Loss Device efficiency ratio:  $\eta_{LP} / \eta_L = 3.9 \pm 0.4$

Loss tangent defined as the angle in the complex plane between the resistive component of an EM field and the reactive component

$$\tan \delta = \frac{\omega \epsilon'' + \sigma}{\omega \epsilon'}$$

Power dissipated in microstrip as:

$$P = P_0 e^{-\delta k z}$$



$$\delta = \frac{\ln(P_{LP}/P_L)}{k \Delta z}$$

**Loss tangent of  $\text{Si}_3\text{N}_4$  :**

$$\delta = 0.0016$$

8.4 mm of microstrip  
between antenna and device

Band	Transmission
0	0.88
1	0.83
2	0.79
3	0.75

- From epsilon test devices  $F\delta_0^{TLS} = 0.00085$
- Roughly equal to previously measured values for  $\text{SiO}_2$
- Factor of 2.25 larger than value measured in BICEP2 for  $\text{Si}_3\text{N}_4$

*Credit: Sunil Golwala*



# Excess Load

- Load from the dewar being absorbed by our dark detectors:

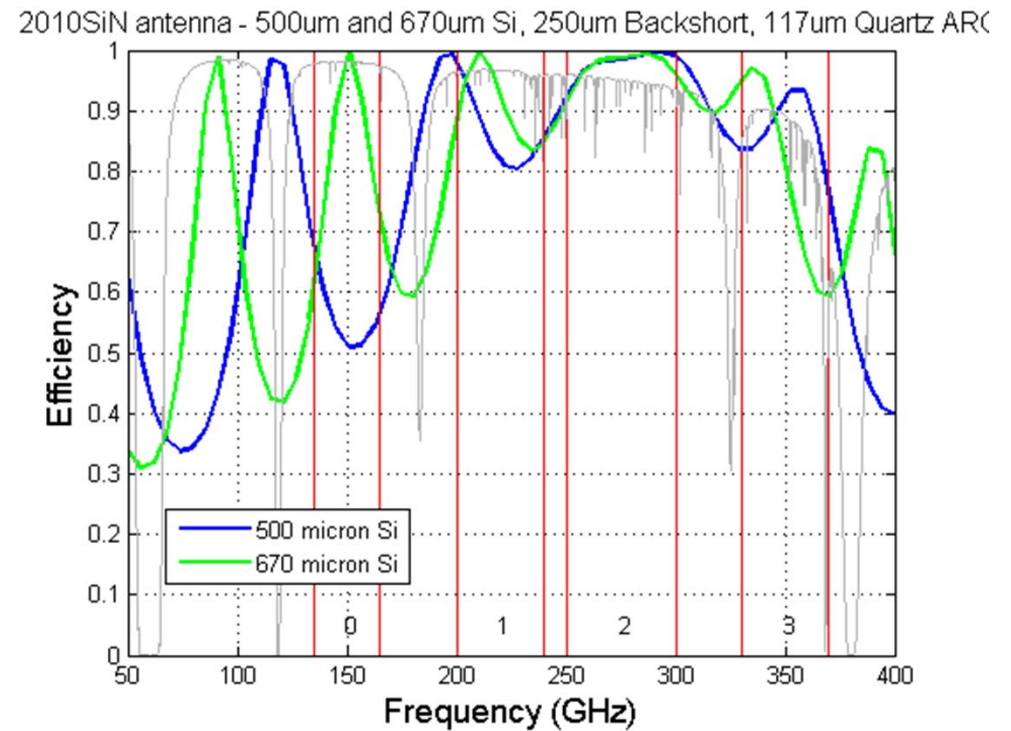
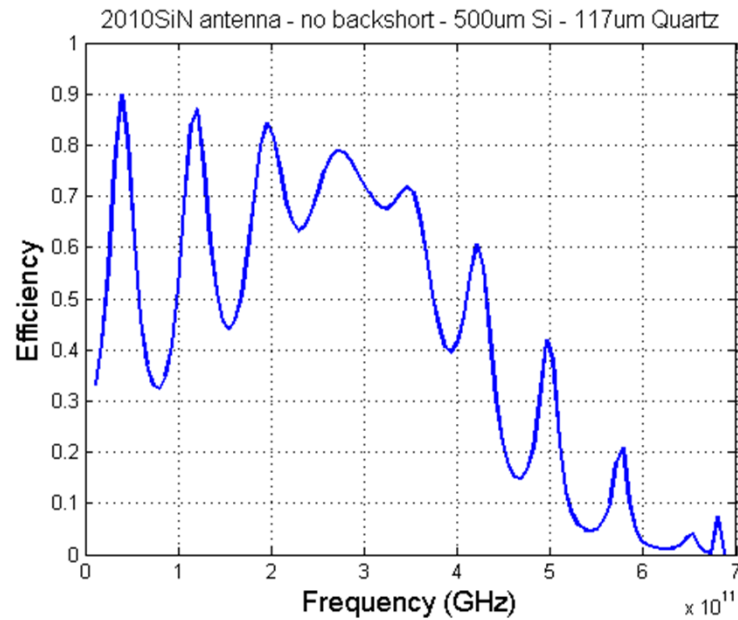
$$P_{\text{dewar}} = 0.4 \text{ pW}$$

- Load from the dewar being absorbed by our antenna coupled detectors:

$$P_{\text{dewar}} = 1.0 - 1.5 \text{ pW} \quad (\text{depending on band})$$

- Corresponds to an excess load of 30 - 40 K
  - 10 – 15 K due to direct pick-up
  - 20 – 25 K due to antenna beam spilling onto inside of dewar

# Nb Backshort



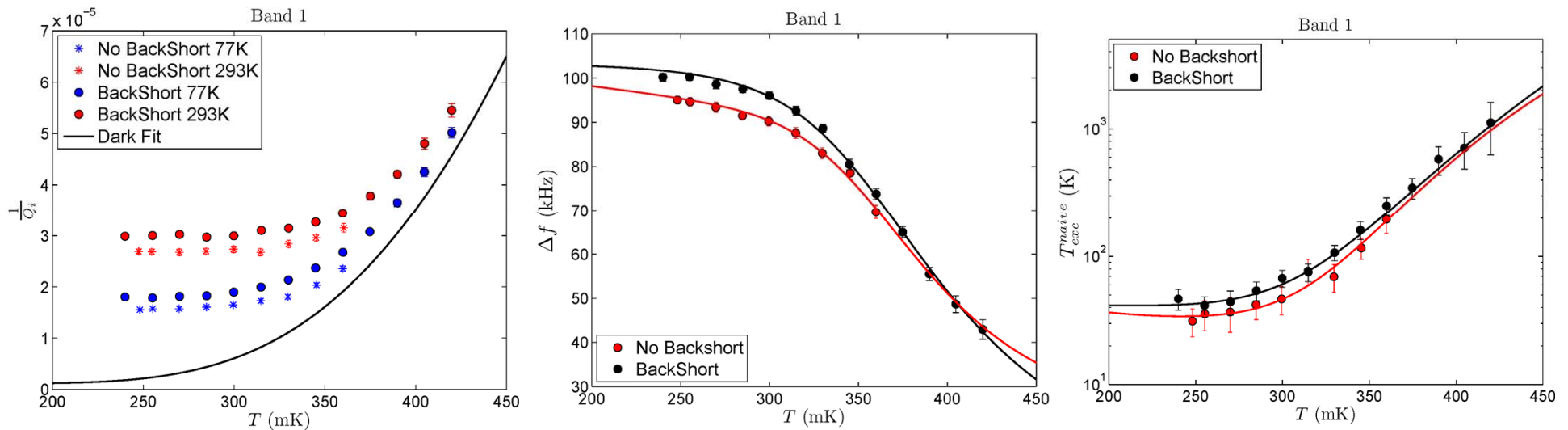
*Credit: Peter Day*



Current array has **500  $\mu$ m Si** with **117  $\mu$ m Quartz AR Coat.**

Implemented a **250  $\mu$ m Nb backshort.**

# No Backshort vs. Backshort



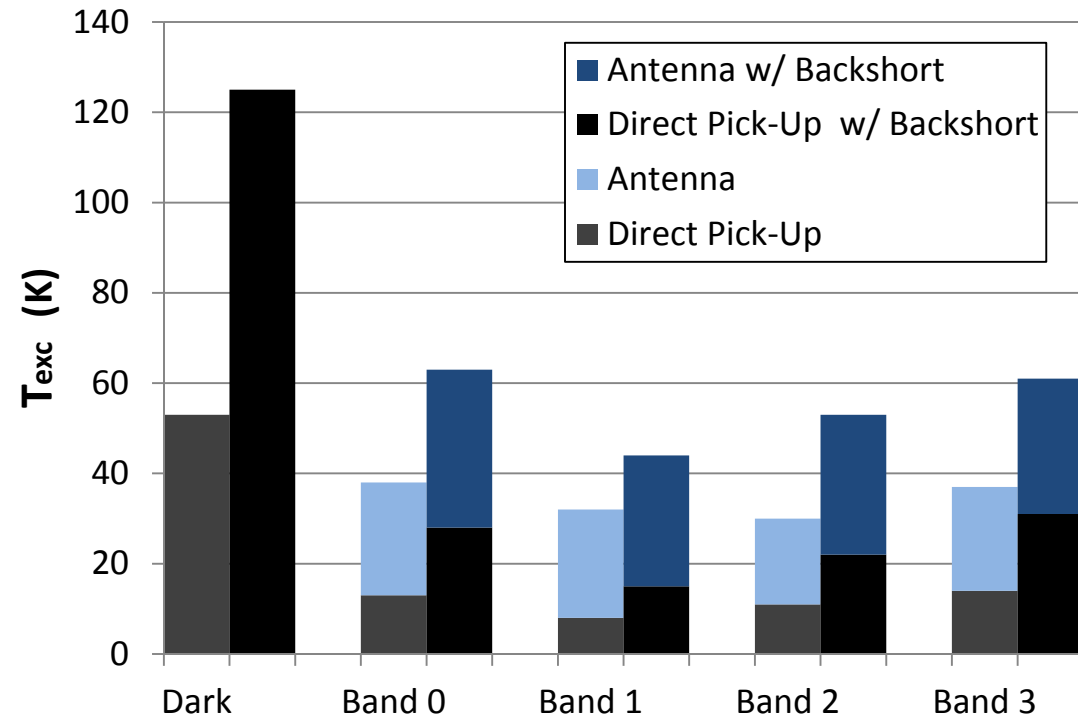
- Improvement in optical efficiency

Band	Measured Efficiency	Measured Improvement	Expected Improvement
0	0.060	9%	-8%
1	0.099	21%	32%
2	0.085	17%	26%
3	0.093	24%	29%

- Efficiency of direct pick-up remained the same

# No Backshort vs. Backshort

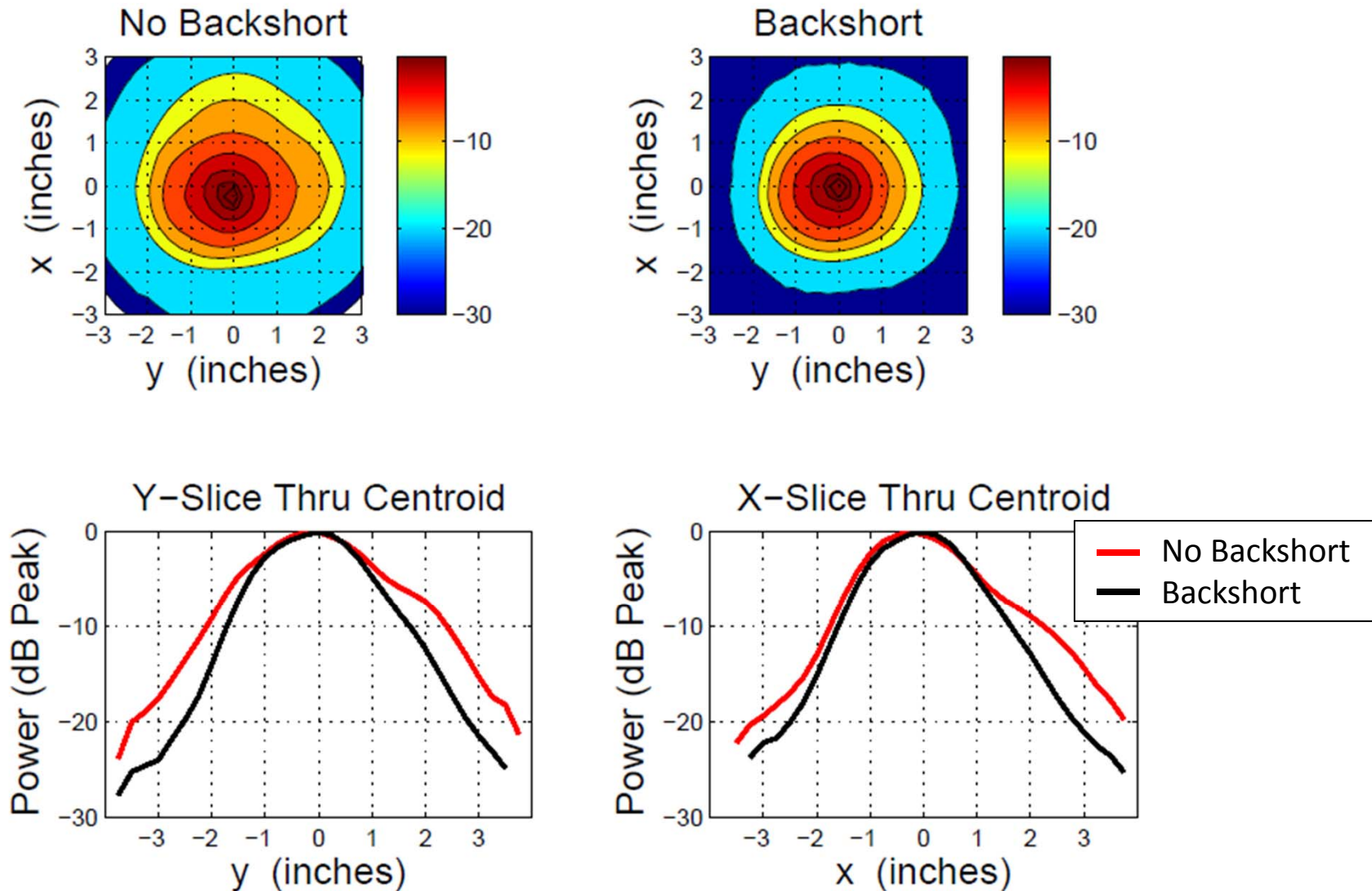
- Note that with the addition of the back short, we also **blackened the inside of the magnetic shield**
- We saw a large increase in excess load
  - $P_{\text{dewar}}$  doubled for both dark resonator and antenna coupled resonators



- Could be due to two competing effects
  1. Increased efficiency for direct absorption due to the backshort (increases  $\eta_{\text{dir}}$ ).
  2. Decreased coupling to outside world due to blackening of the mag shield. Large angle response now terminates inside dewar (decreases  $\eta_{\text{dir}}$ ).

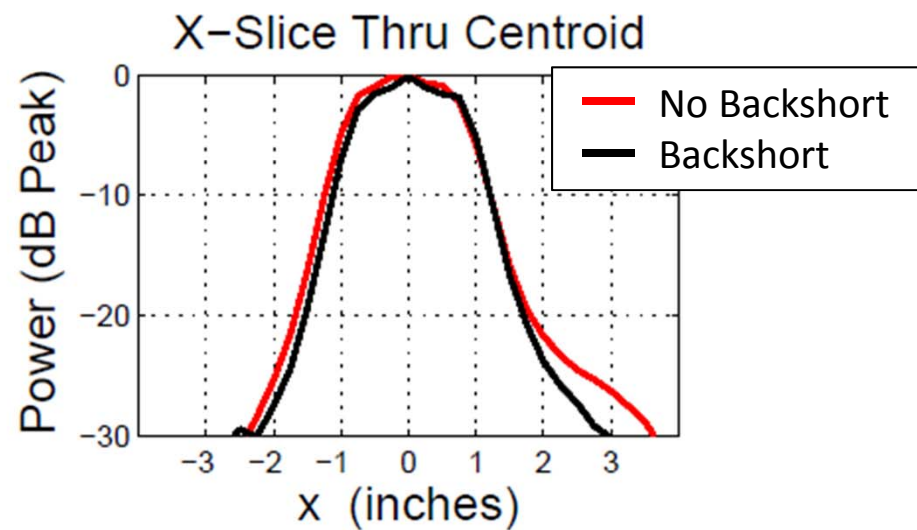
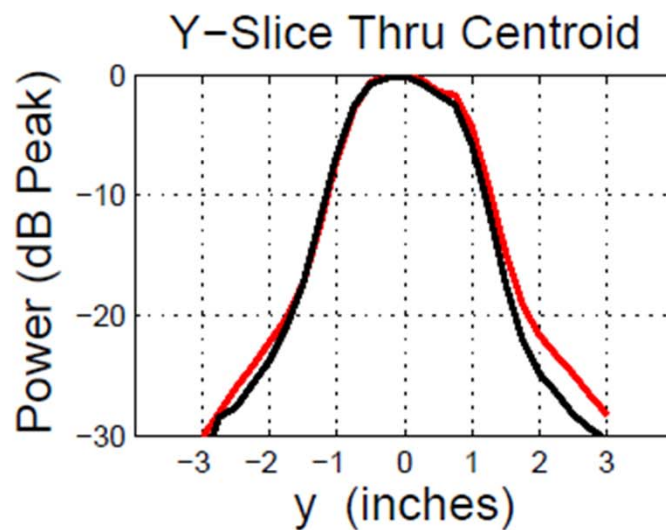
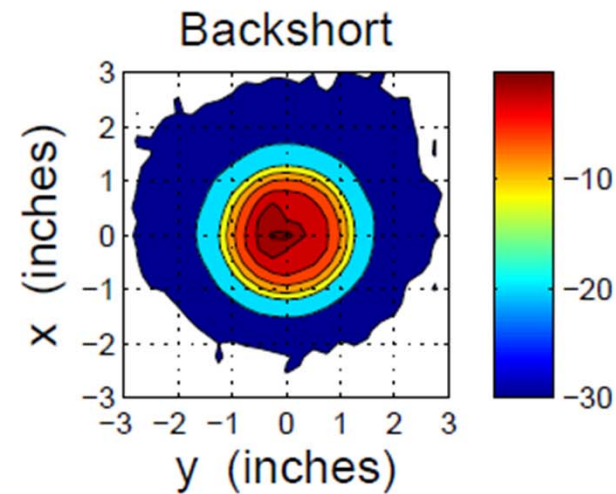
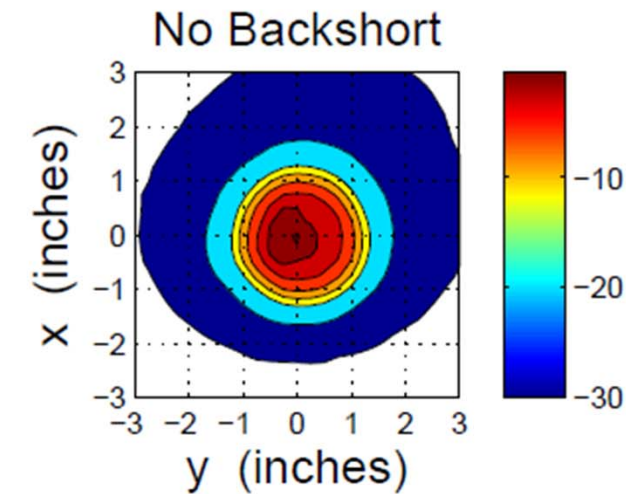
# Beammaps

## Stacked Map for Non-Antenna Coupled Resonators



# Beammaps

## Stacked Map for Band 3 Resonators

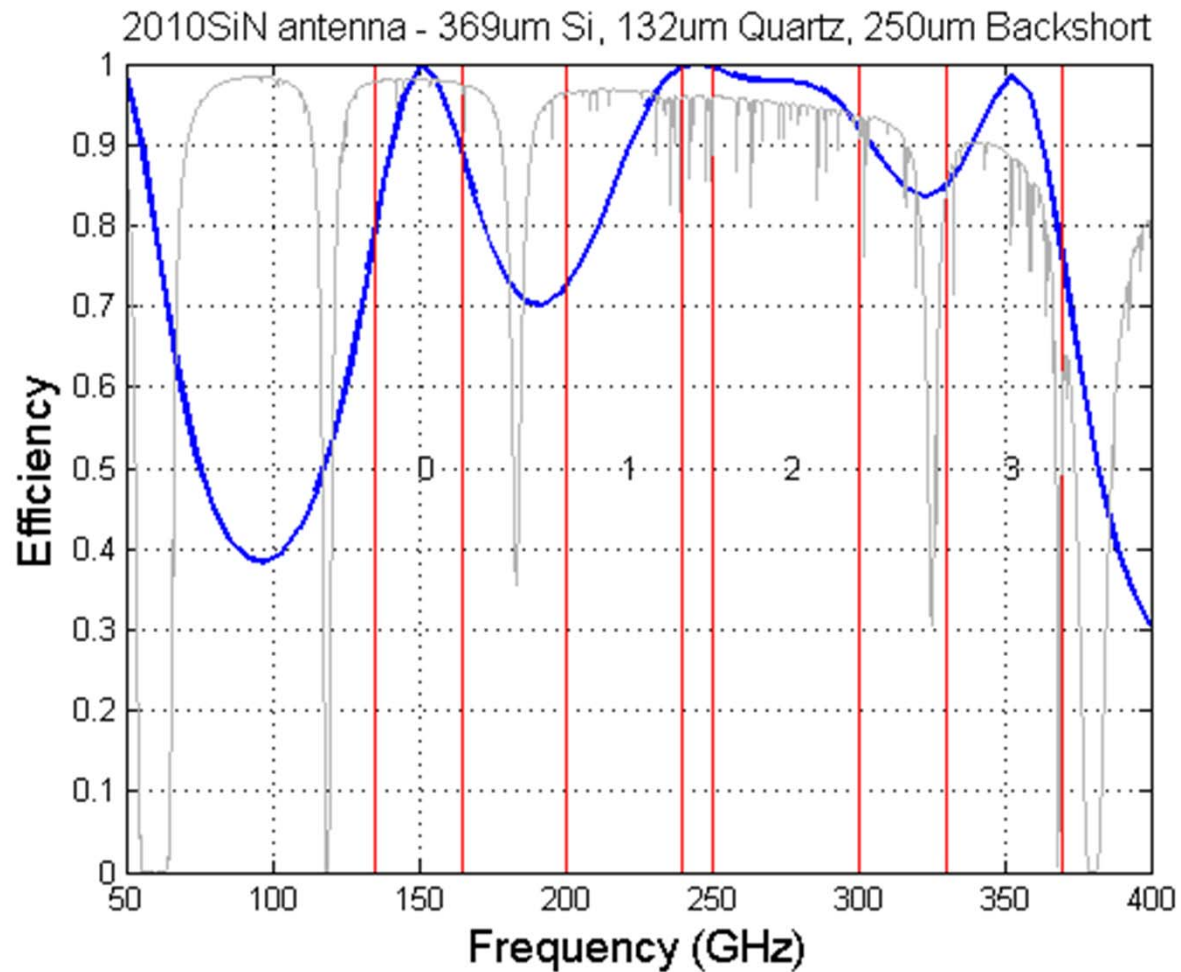




# Nb Backshort

Room for improvement, particularly in Band 0.

**Transmission for optimized Si and AR thickness:**



*Credit: Peter Day*

# NEP

*Defined as the signal power required to achieve a signal-to-noise ratio of 1.*

$$\text{NEP}(\tilde{f}) = \sqrt{S_X(\tilde{f})} \left| \frac{dX}{dP_{\text{opt}}} \right|^{-1}$$

## Dominant Sources of Noise:

- Photon:  $\text{NEP}_{\gamma}^2 = 2P_{\text{opt}} h\nu_0 + \frac{2P_{\text{opt}}^2}{\Delta\nu}$

- Two Level System (TLS) noise:  $S_{\delta f/f}^{\text{TLS}}(\tilde{f}) \propto T^{-2} P_{\text{read}}^{-1/2} \tilde{f}^{-1/2}$

- HEMT noise:  $S_{\delta S_{21}}^{\text{HEMT}}(\tilde{f}) = \frac{kT_n}{2P_{\text{read}}}$  where  $T_n = 5 \text{ K}$

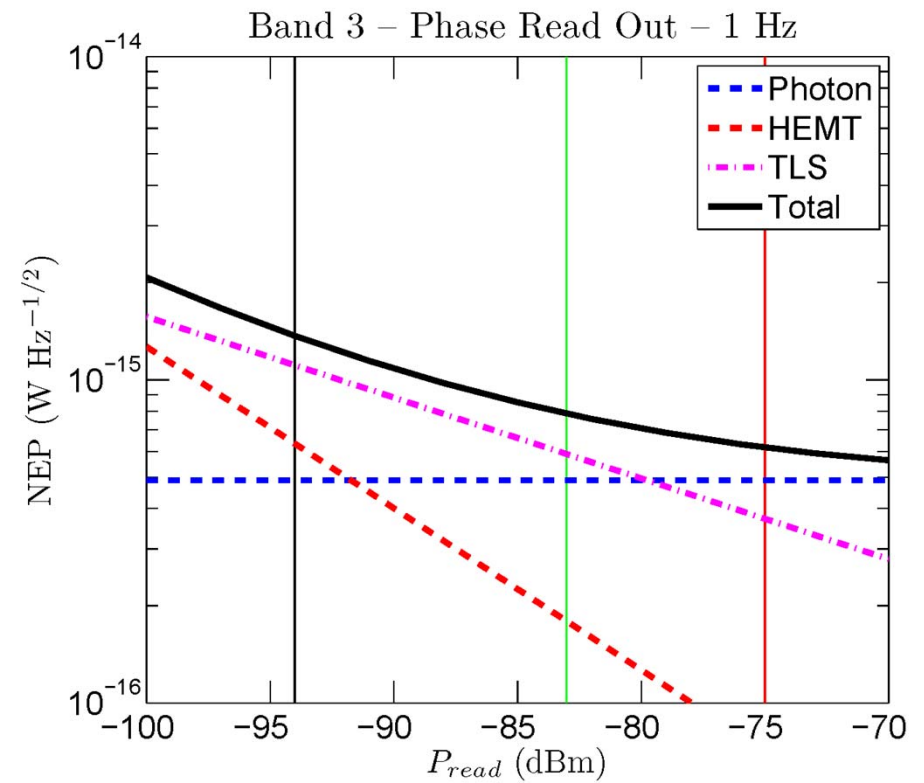
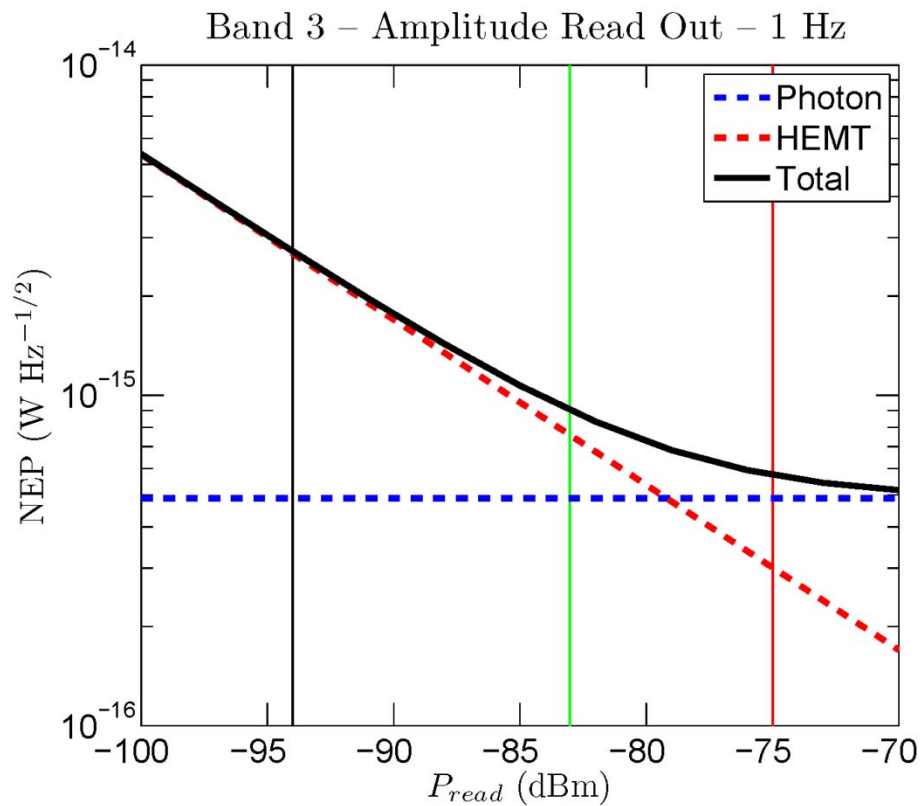


- Loading at CSO with MUSIC will be different from in lab, 77K load

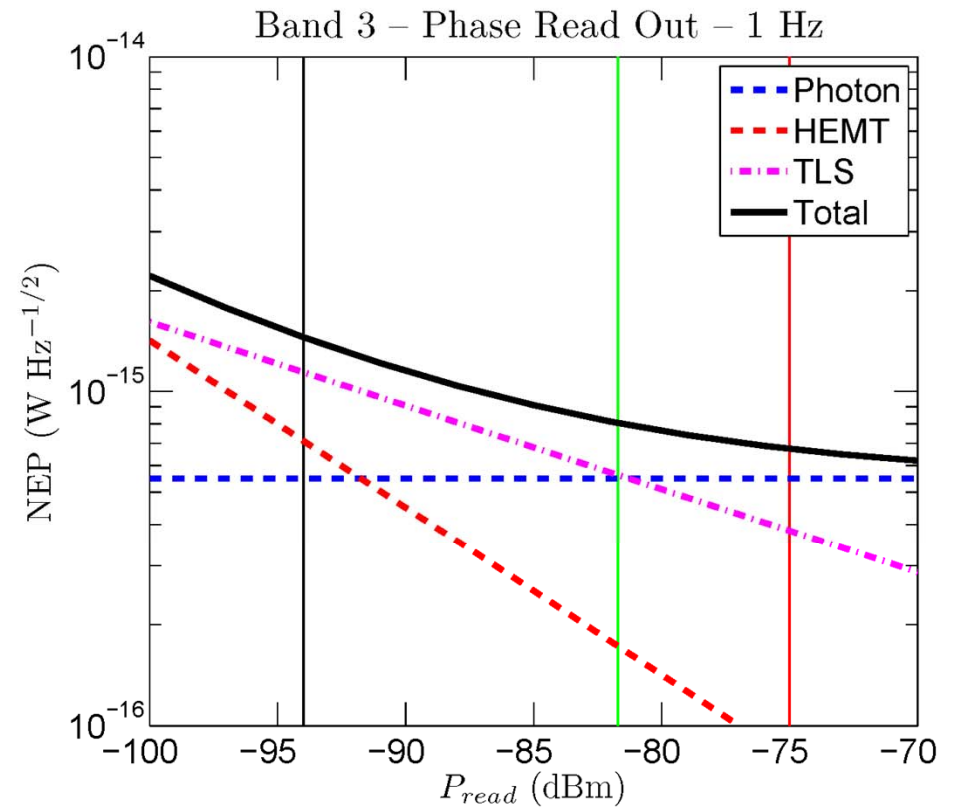
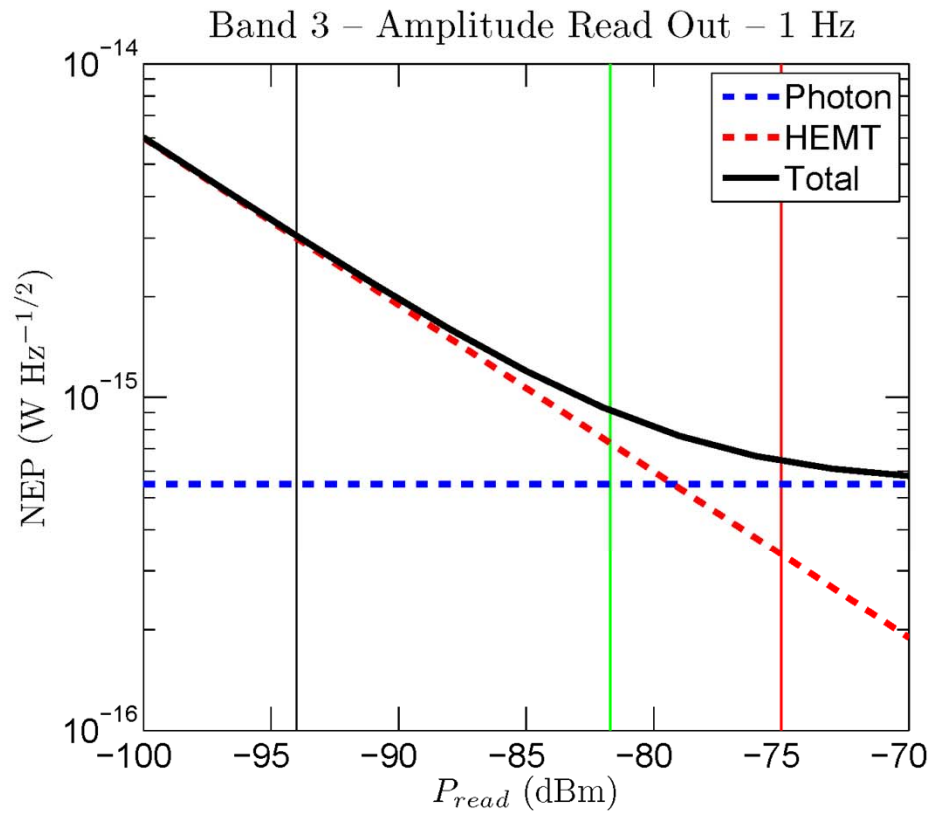
Band	Atmospheric Load
0	16 K
1	27 K
2	42 K
3	80 K

*Credit: Jack Sayers*

### DemoCAM NEP -- No Backshort



## DemoCAM NEP -- Backshort



**In Conclusion...**

# Accomplishments

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- **Reduction of substrate heating** (through the addition of gold wire bonds) **to a negligible level** (as measured by epsilon test devices).
- Better understanding of our optical efficiency.  
**Measurements roughly match expectations!**
- 15-25% improvement in efficiency with the introduction of a Niobium back-short (at the cost of increased excess load).
- Experimental results point to physics that we do not fully understand:
  - Anomalous frequency-to-dissipation ratio
  - Quasi-particle heating

# Challenges

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- Understanding the mechanism for direct pick-up
  - New “direct pick-up” mask currently being tested
- Optimizing MKID read-out power
- Removing correlated, low-frequency electronics noise
- Characterizing and reading out 2304 MKIDs

